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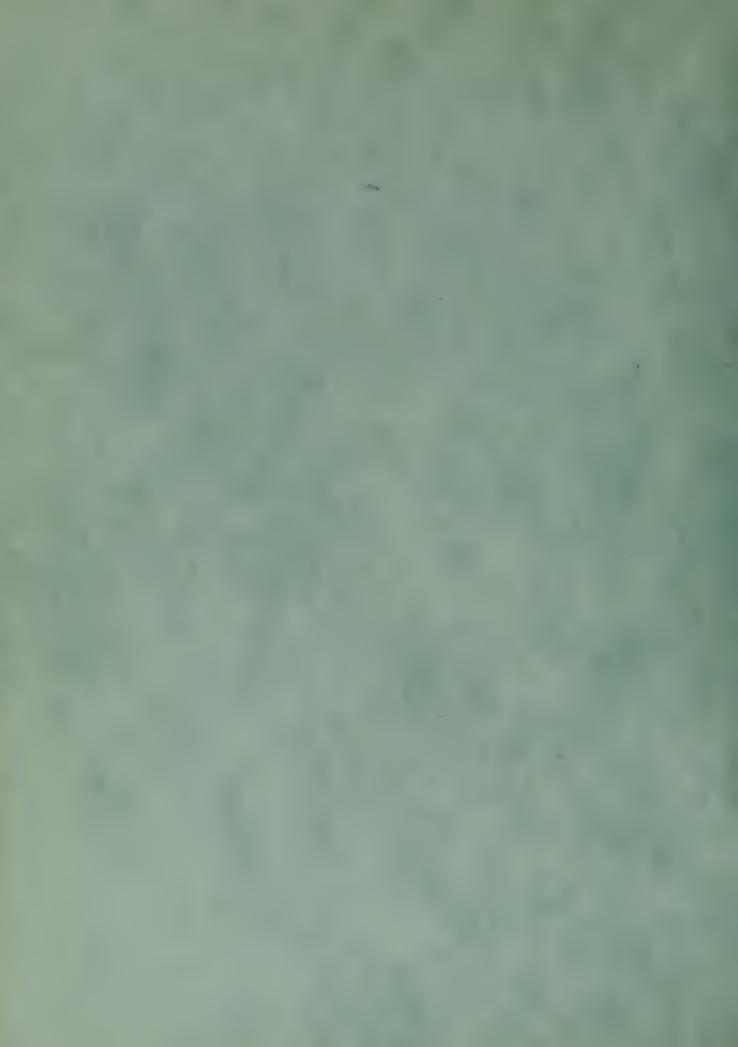
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# HOLOGRAPHIC INTERFEROMETRY OF THE FLOW FIELD BETWEEN A FIN AND FLAT PLATE

Robert Ward Heyer



# NAVAL POSTGRADUATE SCHOOL Monterey, California



# THESIS

Holographic Interferometry of The Flow Field Between a Fin And Flat Plate

by

Robert Ward Heyer

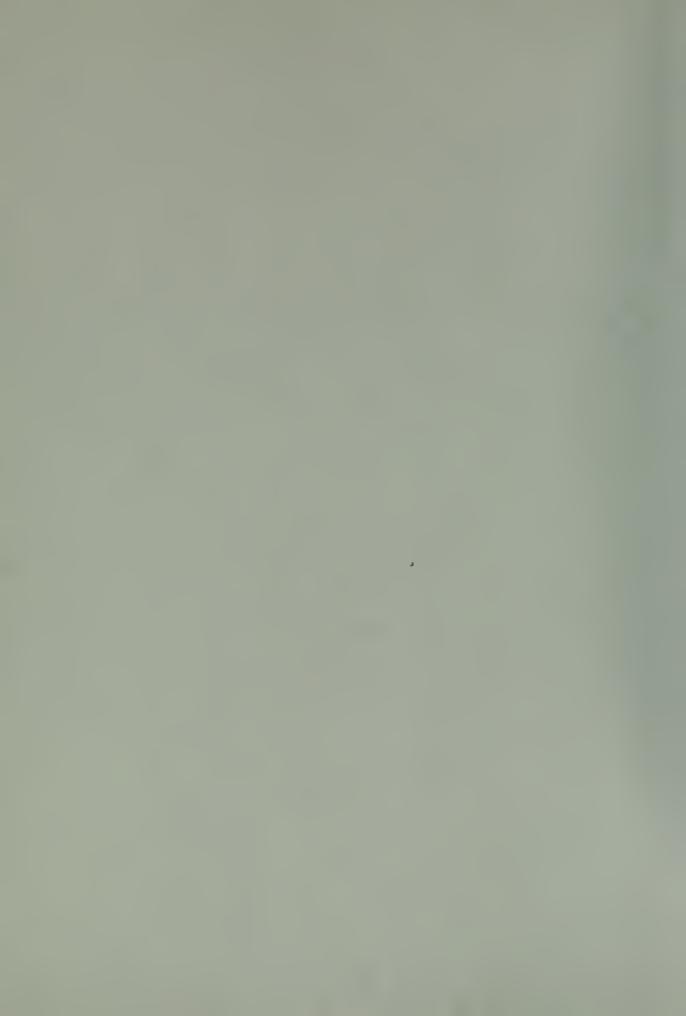
Thesis Advisor:

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March 1972

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Holographic Interferometry of The Flow Field Between a Fin And Flat Plate

bу

Robert Ward Heyer Lieutenant, United States Navy B.S.E.E., Duke University, 1964

Submitted in partial fulfillment of the requirements for the degree of

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from the

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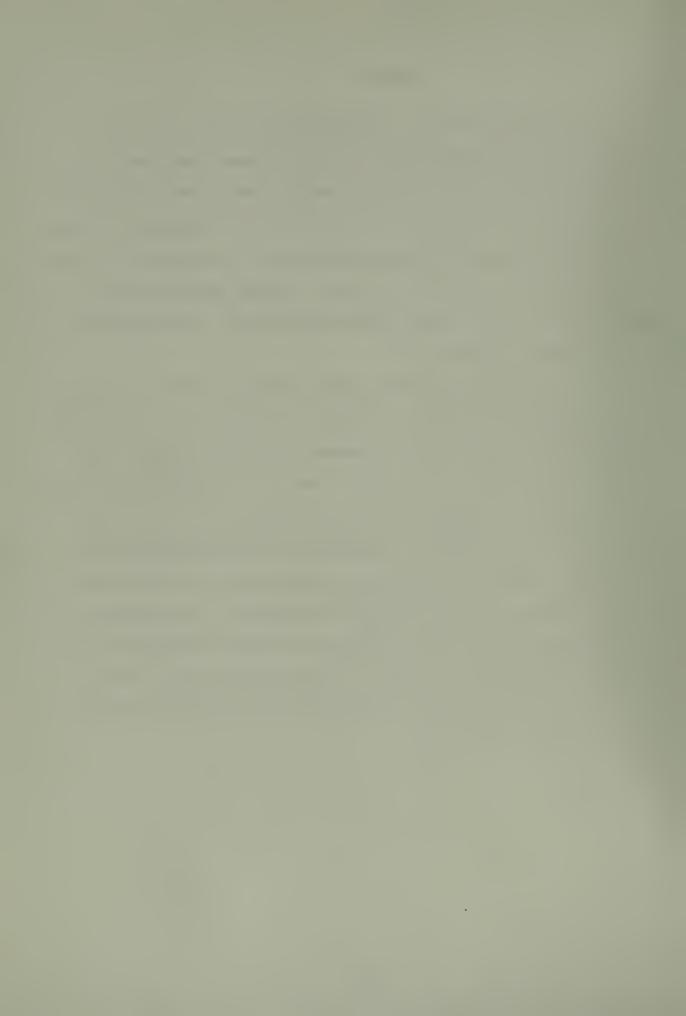
#### **ABSTRACT**

This study was an attempt to map the density field in a fin-flat plate junction three-dimensionally using holographic interferometry.

This investigation has extended the density studies by Matulka [12, 13] and Jagota [3, 4] to include, for the first time, interferometric fringe information obtained through a transparent model in supersonic flow. The fringe information was then inverted by a FORTRAN computer program to produce a plot of the density field around the model. The feasibility of the method was demonstrated.

The factors which are thought to have limited the success of the experiment include vibration of the model, fluctuations of the tunnel flow and the fact that the model was somewhat too large in relation to the size of the wind tunnel test section. Schlieren photography was used to look through and around the model and to verify that the same flow was established as was reported by Thomas [23, 24] and Winkelman [26, 27].

The data reduction of holographic interferograms was, for the first time, accomplished using photographic enlargements. This technique is considered to be much easier and more accurate than the one used in the previous investigations. However, the data reduction step, because of the time and labor involved, is considered to be the rate controlling process of the whole analysis.



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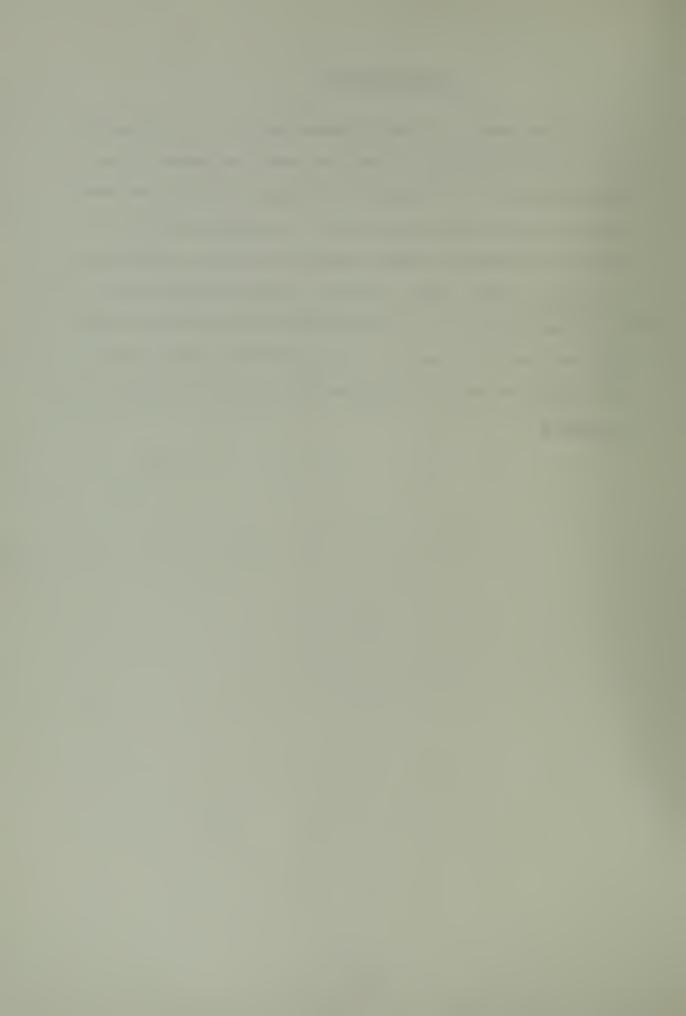


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## I. INTRODUCTION

The determination of the flow field in the wing-body junction of an aircraft in supersonic flight presents many problems. The typical approach has been to measure the static pressure on the surface using many small pressure taps [23, 24] and to determine the velocity field in the junction using translatable pressure probes [18, 19, 25]. Thomas [23, 24] and Winkelmann [26, 27] used azobenzene and oil-smear tests to view the flow field around a flat plate-fin junction. From the streak patterns on the plate and fin, they were able to illustrate the three-dimensional flow field, although in a partially speculative fashion. This study has attempted to map the density field in a fin-flat plate junction three-dimensionally using holographic interferometry. Although this objective was not fully achieved, the feasibility of the method has been demonstrated.

By using a Q-switched laser with exposure times of about twenty nanoseconds, it was possible to obtain three-dimensional holographic interferograms of the density field in the fin-flat plate junction. From holograms taken at a number of viewing angles the fringe shifts in different planes could be obtained. By integrating this information using a FORTRAN computer program, the density field can be determined. This technique has been previously demonstrated for the flow field of a free jet by Matulka [12, 13] and for the supersonic flow field around a cone at angle of attack by Jagota [3, 4].

The tests were performed at the Naval Postgraduate School, using the four-inch supersonic wind tunnel.



# II. EXPERIMENTAL APPARATUS

## A. THE WIND TUNNEL

The investigation was conducted in the Naval Postgraduate School blowdown-to-atmosphere supersonic wind tunnel. The test section is four inches by four inches in cross section and six inches long with two different sets of side walls. The two-inch thick plexiglas side walls, which have a refraction index of 1.49, present a complete field of view of the flow from the nozzle throat to aft of the test section mounting bracket (Figure 1 (a)). The second set of sidewalls used are aluminum with high quality optical glass portholes located in the test section area (Figure 1 (b)). The interchangeable nozzle for a test section Mach number of 2.8 was used for all tests. The nominal run time is five minutes at Mach 2.8 with the maximum stagnation pressure of about 105 pounds per square inch.

# B. THE HOLOGRAPHIC ARRANGEMENT

The holographic arrangement is illustrated in Figure 2 and shown in photographs included as Figures 3, 4, and 5. The equipment stand was rested on a portion of the building floor that was vibrationally isolated. A Konrad K-1 pulsed ruby laser with a Pockels cell Q-switching unit was used to produce monochromatic light at a wave length of 6943 Angstroms and exposure time of twenty nano-seconds. The laser cavity length was seventy-three cm. giving a coherence length of about ten cm. To maintain the laser head and output etalon at a constant temperature of 27.5 degrees centigrade, a Lauda constant temperature circulator Model N was used. This was controlled by an electronic relay type R-10 coupled with a Culligan de-ionizer.



Holograms were obtained by routing the reference beam under the wind tunnel and the scene beam through the test section, and intersecting the two beams on the hologram plate at an intersecting angle of approximately 50 degrees. The beam sizes were controlled by translating the concave lenses located between the beam splitter and hologram plate in each beam (Figure 2). The Q-switched laser and optics were aligned using a continuous wave helium-neon laser. For reference purposes, grids were mounted on the outside tunnel walls and aligned using a surveyor's transit. The holographic stand and test section were completely enclosed in a wooden box to enable holograms to be taken in the daylight (Figure 6).

## C. THE WIND TUNNEL MODELS

The fin-flat plate models used are shown in Figures 7 (a), (b), (c), and (d). The metal portions of both models were stainless steel. The center section, part of one strake, and all of the other strake of the first model in Figure 7 (a) were made of epoxy while the center section of the second model, Figures 7 (b) and 7 (c), was fabricated from plexiglas. The flat plate grids in both models were etched into the plastic surface and coated with a clear plastic to achieve a smooth surface.

The models were rotated about their sting mounts as shown in Figure 7 (d). Alignment for the desired rotation angle was accomplished by aligning prescribed lines on the sting mount collar with a scribed mark on the sting stand using a surveyor's transit.



# III. ANALYTICAL EVALUATION OF THE DENSITY FIELD

# A. THE BASIC INTERFEROMETRIC EQUATION

Interferograms are created when two coherent light beams are superimposed on each other and projected on a viewing screen. The light and dark regions observed correspond to the relative phase difference between the two beams which are caused by a difference in the two optical path lengths. Consider a coherent beam which is split and then recombined on a viewing screen. A difference in optical path lengths of the two beams may be achieved in two ways in order to create an interferogram. The first is to make the physical distance traveled by the two beams different. In a vacuum this path length difference is expressed as  $\mathbf{L} = \mathbf{C_0}\Delta \mathbf{t}$  where  $\mathbf{C_0}$  is the speed of light in a vacuum. The second way is to maintain equal physical path lengths but to have the beams traverse through different media prior to recombining. In this case each light beam will travel at a speed  $\frac{\mathbf{C_0}}{\mathbf{n}}$  where n is the index of refraction for the medium traversed. The optical path length difference then becomes:

$$\Delta L = L (n_2 - n_1) = \mathbf{c}_{o} \Delta t$$
 (1)

The interference pattern or fringes observed may be expressed as a function of the optical path length difference or

$$g = \frac{\Delta L}{\lambda}$$
 (2)

where:

g = fringe shift

 $\lambda$  = wave length of the light source

 $\Delta L$  = change in optical path



Combining equations (1) and (2), the fringe shift is then

$$g = \frac{L}{\lambda} (n_2 - n_1) \tag{3}$$

The index of refraction is known to be a function of density. Since the speed of light is only slightly less in gases than in a vacuum, the index of refraction could be closely approximated by the series expansion [8]

$$n = 1 + \beta \frac{\rho}{\rho_s} \tag{4}$$

where

 $\beta$  = dimensionless constant related to the Gladstone-Dale constant by  $K = \frac{\beta}{\beta_3}$ 

 $\rho_{\rm s}$  = reference density of 0° C, 760 mm. Hg.

The variation of with wavelength is small and has a value of 0.000292 for  $\lambda$  = 5893 angstroms.

Considering a fixed difference in the index of refraction between the two beams in Equation (3), then

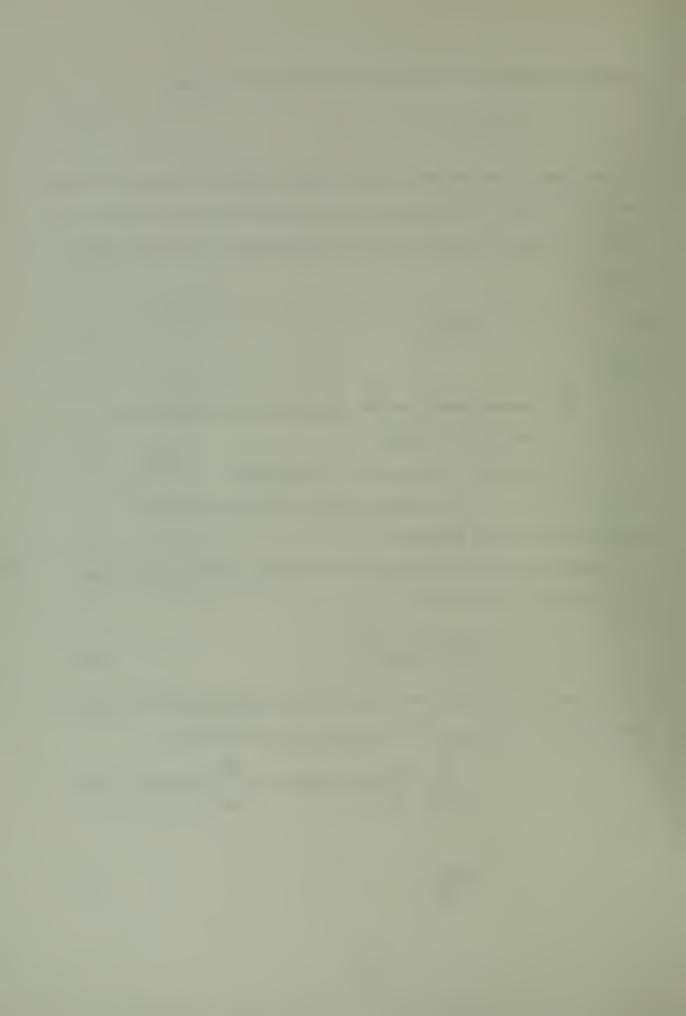
$$g = \beta \frac{L}{\lambda} \left( \frac{\rho_2 - \rho_{\infty}}{\rho_s} \right) \tag{5}$$

If the density varies in a beam path, the net change in the optical path length will be the integrated effect along the beam path or

$$g = \frac{\beta}{\lambda \rho_s} \int_0^L (\rho - \rho_{oo}) ds = Q \int_0^L f(x, y, z_c) ds \qquad (6)$$

where

$$Q = \frac{\beta \rho_{\infty}}{\lambda \rho_{3}}$$
 (6a)



$$f(x,y, Z_c) = \frac{P(x,yZ_c)}{P\infty} - 1$$
 (6b)

 $Z_c$  = a plane of constant Z

ds = incremental distance along the ray

In order to determine the density along the beam path where the fringe shift is known from an interferogram, Equation (6) must be inverted.

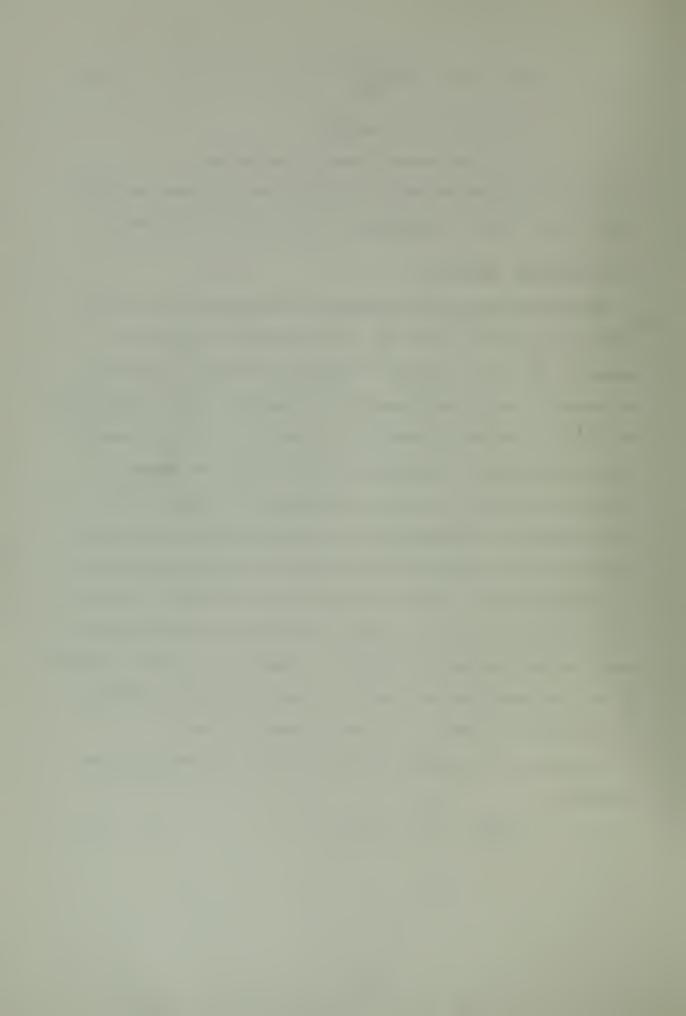
#### B. THE INTEGRAL INVERSION

The integral inversion technique was first reported by C. D. Maldonado et al in 1965 [9, 10, 11]. R.D. Matulka [12, 13] and R. C.

Jagota [3, 4] used this method to determine the density variation in an asymmetric free jet and about a cone at angle of attack, respectively. The technique involves representing the function, f(x,y,z) in Equation (6) by a complete set of orthogonal functions where the unknown coefficients are evaluated using the orthogonality relationship between the set of functions. The functions are orthogonal over the entire plane and also have the propery of being invariant in form to any rotation of the coordinate system. Figure 8 illustrates the coordinate system for the inversion where x and y are the fixed laboratory coordinates and x' and y' are the coordinates in which the fringe number function is defined. As the view through the test section is varied the primed coordinates are rotated with respect to the fixed coordinates x and y.

The fringe shift expressed in Equation (6) may be written as the transform

$$g(\xi, y, z_c) = f(x, y, z_c)$$
 (7)



or, inverting the equation, the density function, f, is equal to:

$$f(x, y, z_c) = T_g(\xi, y, z_c)$$
 (8)

The density function can be expanded in the following manner using a set of polynomial functions,  $U_{m+2k}^{\pm m}$  (x, x, y), and unknown comples coefficients, C m+2k

$$f(x,y,z_{c}) = \sum_{m=0}^{\infty} \sum_{k=0}^{\infty} \epsilon_{m} \left[ \binom{m}{m+2k} (x) \bigcup_{m=2k}^{m} (x,x,x,y) + \binom{-m}{m+2k} (x) \bigcup_{m+2k}^{-m} (x,x,x,y) \right],$$
where
$$\epsilon_{m} = \left\{ \frac{1}{2} m = 0 \\ 1 m = 1,2,3,\cdots \right\}$$
(9)

The polynomial functions,  $U_{m+2k}^{\pm m}$  (4x, 4y), are invariant in form to a rotation of the coordinate system [10, 12, 13]. They also have a Gauss transform which makes them adaptable to the physical situation and to manipulating into the form of Equation (7). The functions are defined as:

$$U_{m+2k}^{\pm m}(\alpha x, \alpha y) = (-1)^{k} \propto \left[ \frac{k! (\alpha^{2} x^{2} + \alpha^{2} y^{2})^{m}}{T(m+k)!} \right]^{\frac{1}{2}} e^{\pm i m \phi} L_{k}^{m} (\alpha^{2} x^{2} + \alpha^{2} y^{2})$$
(10)

where 
$$\phi = \tan^{-1}\left(\frac{y}{x}\right) - \frac{\pi}{2}$$
 (10a)

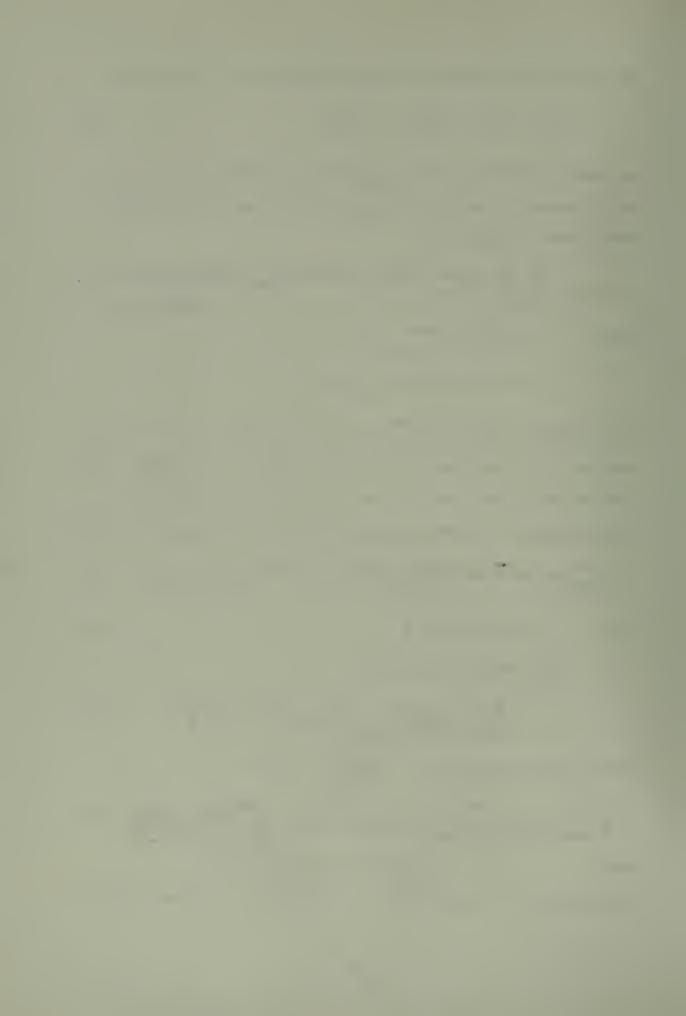
Lm = Laguerre polynomial

$$= \sum_{g=0}^{k} \left[ \frac{(m+k)!}{(k-9)!(m-3)!} \right] \left[ (-1)(\alpha^2 x^2 + \alpha^2 y^2) \right]^{8}$$
 (10b)

Umask And the Gauss Transform of

$$I_{m+2k}^{\pm m}(\alpha y', \xi) = \int_{-\infty}^{\infty} \int_{m+2k}^{\pm m} (\alpha x, \alpha y) e^{-\alpha^2 x'^2} dx' = \frac{e^{\pm im\xi} H_{m+2k}(\alpha y')}{[k! (m+k)!]^{\frac{1}{2}} 2^{m+2k}}$$
where
$$H_{m+2k}^{(\alpha y')= \text{ Hermite polynomials}}$$
(11)

By applying the transform above to Equation (9), the fringe function in



Equation (7) can be written as:

$$9(\$, y_1 = 0) = \sum_{m=0}^{\infty} \sum_{k=0}^{\infty} \frac{e_m \left[ C_{m+2k}^m(\alpha) e^{im\$} + C_{m+2k}^m(\alpha) e^{-im\$} \right] H_{m+2k}(\alpha) e^{-in\$} J_{m+2k}(\alpha) e$$

using the following orthogonality relationship on Equation (12)

$$\int_{-\pi}^{\pi} e^{\pm im\xi} e^{\mp in\xi} \int_{-\infty}^{\infty} H_{m+2k}(\alpha y') H_{n+2k}(\alpha y') e^{-d^{2}y^{2}} dy' = \frac{2\pi^{3/2}}{\alpha} \left[ (m+2k)! (n+2k)! 2^{m+2k} 2^{n+2k} \int_{mn}^{\infty} \delta_{(m+2k)!(n+2k)} \right]$$
(13)

where  $\delta$  is the kroneker delta, the expansion coefficients  $C_{m+2k}^{\pm m}$ , can be determined by:

$$C_{m+2k}^{tm}(x) = \frac{\alpha}{2\pi^3 2} \left[ \frac{(k!(m+k)!)^{\frac{1}{2}}}{(m+2k)!} \right] \int_{-\pi}^{\pi} g(y', \xi, z_c) H_{m+2k}(\alpha y') e^{-im\xi} dy' d\xi$$
 (14)

Substitution of the coefficients in Equation (14) bact into Equation

(9) results in the density variation being expressed as:

$$f(x,y,z_{c}) = (\frac{1}{\pi})^{2} \sum_{m=0}^{\infty} \sum_{k=0}^{\infty} \epsilon_{m} [k!(m+k)!]^{\frac{1}{2}} e^{-(A^{2}x^{2}+A^{2}y^{2})}.$$

$$REAL \left[ \int_{-\pi}^{\pi} \int_{-\infty}^{\infty} g(y',\xi,z_{c}) e^{-im\xi} H_{m+2k}(x'y') dy' d\xi \right] U_{m+2k}^{m}(x',x',y')$$
(15)

or by inserting Equation (10):

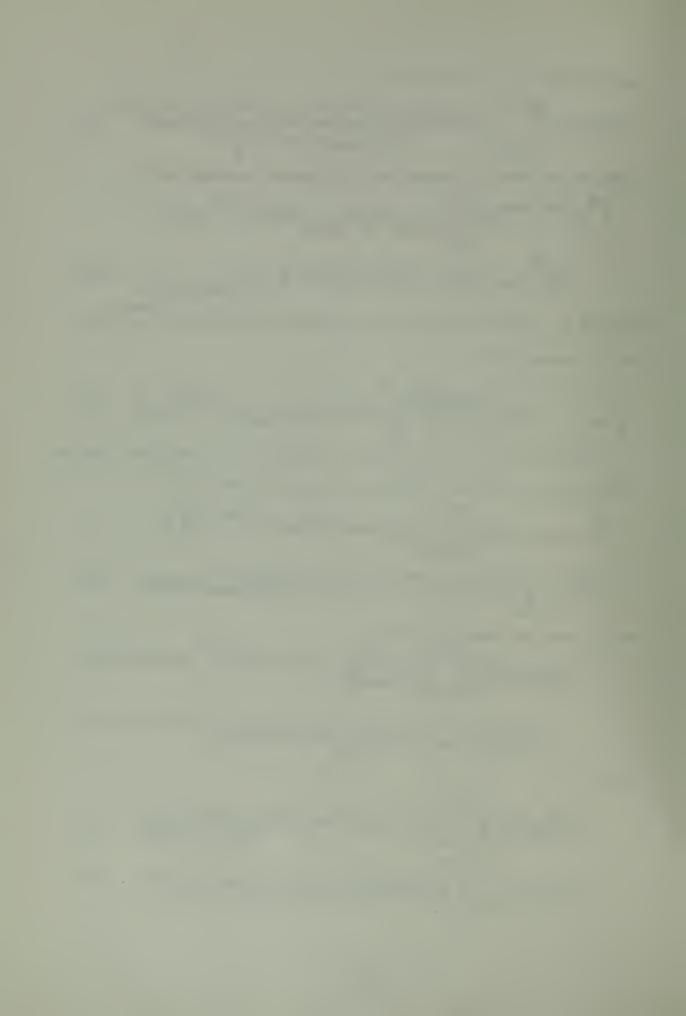
$$\int (x,y_1z_c) = \left(\frac{2}{11}\right)^2 \sum_{m=0}^{\infty} \sum_{k=0}^{\infty} \epsilon_m \frac{(-1)^k k!}{(m+2k)!} \left( d^2x^2 + d^2y^2 \right)^m \sum_{k=0}^{m} \left( d^2x^2 + d^2y^2 \right)^{-m}$$

$$[B_{m+2k}^{m}(a)\cos(m\phi) + D_{m+2k}^{m}(a)\sin(m\phi)]e^{-(a^{2}\chi^{2}+a^{2}y^{2})}$$
(16)

where:

$$B_{m+2k}^{m}(\alpha) = \int_{-\infty}^{\pi} \int_{-\infty}^{\infty} q(y', \xi, \xi_{c}) \cos(m\xi) H_{m+2k}(\alpha y') dy' d\xi \qquad (17)$$

$$D_{m+2k}^{m}(\alpha) = \int_{\pi}^{\pi} \int_{0}^{\infty} g(y', \xi, z_{c}) sim(m\xi) H_{m+2k}(\alpha y') dy' d\xi$$
 (18)



Equations (16), (17), and (18) are the basic equations used to calculate the density distribution from the experimentally determined fringe variations.

## C. THE NUMERICAL PROCEDURE

The form of the density distribution in Equations (16), (17), and (18) must be modified in order to input the experimentally determined fringe distribution. First from Figure 8 and Equation (6b) it can be seen that it is only necessary to integrate Equations (17) and (18) over an area where the density is changing from a known density,  $\rho_{\infty}$ . Outside of this region where there is no change in density, the function  $f(x,y,z_0) = 0$ , i.e. outside the test section. Also since the fringe distribution is taken in small increments over the test area the coefficients, B and D, can be approximated as

$$B_{m+2k}^{m}(x) = \sum_{i=1}^{I-1} \sum_{j=0}^{J-1} g(\xi_{j} + \Delta \xi_{j}) x_{i} + \Delta x_{i}) \int_{\xi_{i}}^{\xi_{j+1}} \cos(m\xi) d\xi \int_{x_{i}}^{x_{i+1}} H_{m+2k}(x) dx$$
 (19)

$$D_{m+2k}^{m}(x) = \sum_{i=1}^{k-1} \sum_{j=0}^{k-1} g(\xi_{j} + \Delta \xi_{j}) x_{i} + \Delta x_{i} \int_{\xi_{i}}^{\xi_{j+1}} \sin(m\xi) d\xi \int_{x_{i}}^{x_{i+1}} H_{m+2k}(xx) dx$$
 (20)

The integral of \$\mathbf{S}\$ is easily determined and by using the derivative formula for Hermite polynomials the integral of x may be manipulated to yield

$$B_{m+2k}^{m}(\alpha) = \left[\frac{1}{2\alpha m} \frac{1}{(m+2k+1)}\right] \sum_{i=0}^{T-1} \sum_{j=0}^{T-1} g(\xi_{j} + \Delta \xi_{j}) \times_{i} + \Delta x_{i} + \Delta x_{i} = 0$$

$$[sin(m\xi_{i+1}) - sin(m\xi_{i})][H_{m+zk+1}(\alpha X_{i+1}) - H_{m+zk+1}(\alpha X_{i})]$$
 (21)

$$D_{m+2k}^{m}(\alpha) = -\left[\frac{1}{2Am(m+2k+1)}\sum_{l=0}^{L-1}\sum_{j=0}^{J-1}q(\xi_{j}+\delta\xi_{j})x_{l}+\Delta x_{l}\right) = (22)$$

[ $\cos(mS_{j+i}) - \cos(mS_j)$ ][ $H_{m+2k+i}(XX_{i+i}) - H_{m+2k+i}(XX_i)$ ] In the computation of the density function from Equation (16), ob-

taining the infinite summations experimentally is not plausible or



possible. It has been demonstrated that by using a finite number of terms and by adjusting the values of  $\Delta \xi$ ,  $\Delta \times$  and  $\Delta$ , it is possible to obtain the density distribution with very good accuracy [3, 4, 12, 13]. Equation (16) then becomes:

$$\int (x_{1}y_{1},z_{c}) = \left(\frac{A}{\pi}\right)^{2} \sum_{k=0}^{K} \sum_{m \in D}^{M} \left(\frac{k!}{(m+2k)!}\right) \left(A^{2}x^{2} + A^{2}y^{2}\right) \bullet$$

$$\left[\sum_{k=0}^{M} (A^{2}x^{2} + A^{2}y^{2}) \left[B_{m+2k}^{M}(A)\cos(m\phi) + D_{m+2k}^{M}(A)\sin(m\phi)\right] \bullet$$

$$= \left(A^{2}x^{2} + A^{2}y^{2}\right)$$

$$= \left(A^{2}x^{2} + A^{2}y^{2}\right)$$
(23)

# IV. EXPERIMENTAL PROCEDURE

# A. LABORATORY TECHNIQUES

The analysis of a free jet by Matulka [12, 13] illustrated how holographic interferometry can be used to obtain a complete three dimensional plot of the density within a moving transparent flow field.

Jagota [3, 4] in his study of a cone at angle of attack in supersonic flow went one step further by introducing an opaque object into an assumed steady state flow field and describing the density field three-dimensionally.

This investigation has attempted to determine the three-dimensional density field around a transparent object in a supersonic flow field by passing a light beam through the object. Specifically the interest was to describe the flow field existing in the junction of a fin-root intersection.

# 1. Model Considerations

In order to obtain uniform flow around the fin-root area, a model of the form shown in Figure 9 was selected. The flat plate has a knife



edge and is intended to remain at zero degrees angle of attack so as to establish the flow conditions illustrated. The fin edges were made circular in order to approach flow conditions similar to those established in Winkelmann's [26, 27] investigation. Plastic and metal strakes were added on the model sides so as to maintain two-dimensional flow as well as to add strength to the flat plate. In the first model constructed (Figure 7(a)), maximum visibility of the plastic fin-flat plate center section was achieved by bolting the leading edge and aft plate together through the plastic center section. Due to model flexure, this design was found to be unsuitable and the second model in Figures 7(b) and 7(c) was constructed. The model was made from a single piece of stainless steel. The model rigidity was satisfactory but unfortunately the strengthening borders around the plastic center section reduced the holographic visibility somewhat.

Since the wind tunnel blocks were fixed, it was possible to make multiple test runs with the same flow conditions over the model provided supersonic flow had been established over the model.

# 2. Holographic Techniques

In order to obtain holographic interferograms it was necessary to ensure that the optical path lengths of the scene and reference beams remained approximately equal. Since the ruby laser is believed to have a coherency length of approximately ten centimeters, the equality of lengths is far less critical than in the classical Mach-Zehner interferometric approach. Consequently a string was used in the experiment to trace the reference beam and then adjust the scene beam. This method kept the two beam lengths within one centimeter of each other. Since the scene beam traversed approximately 4.5 inches of plastic tunnel walls and



grids which the reference beam did not, it was necessary to compensate by making the scene beam physically 2.25 inches shorter than the reference beam. The reference beam varied in length from 61 inches to 68 inches during the experimentation.

To determine the fringe/density field, finite fringe interferograms were made by three different techniques. In the direct fin-root flow approach the diffuser plate was part of the model. Either the mirror, M<sub>5</sub>, in Figure 2 or the hologram plate holder was translated between the no-flow exposure and the flow-established exposure. In the total model flow method the diffuser plate was located between seene beam lens, L<sub>3</sub>, and the test section (See Figure 2) and it was translated horizontally or vertically. Translations were varied from .001 inches to .006 inches with the translation distance of .003 inches yielding the best fringe separation.

Most of the holograms taken using basically the holographic arrangement shown in Figure 2 gave well-defined fringe patterns. In order to improve upon the fringe definition, a variety of techniques was attempted. The transverse mode selector was varied from 1.0 mm. to 2.5 mm. in increments of 0.5 mm. to determine the best lighting of the model. The hologram plate holder was rotated horizontally to various positions. These positions varied between being perpendicular to the seene beam to being perpendicular to the bisection of the angle between the scene and reference beams. Polarization plates were added in both the seene and reference beams between the test section and hologram plate in the seene beam and between the last mirror, M4, and the hologram plate in the reference beam. A one-quarter wave plate was also placed between the first lens, L1, and the beam splitter as recommended by Okayama and Emori [14].



The holograms were taken using  $4" \times 5"$  Agfa-Gavaert 8E-75 hologram plates. The developing process involved:

- 1. Five minutes in Kodak D-19 developer
- 2. Thirty seconds in an acetic acid stop bath
- 3. Five minutes in standard fixer
- 4. Five minutes in a flowing water bath
- 5. One minute emersion in Kodak Photo Flo wetting agent
- 6. Drying using blowing cool air
- a. Direct Fin-Root Flow Method

Since the flow field in the fin-root junction is assumed to be identical on either side of the fin, then it is only necessary to determine the density on one side of the fin. To accomplish this it is necessary to obtain holograms for 180° of view as shown in Figure 10. The holograms for the views from 0° to 90° can be obtained by using a frosted fin and flat plate as shown in Figure 11. The advantage of having the frosted plate as part of the model is that the fringe/density information obtained by the interferogram is believed to be only for the area between the fin-root intersection to the tunnel wall vice the whole test section, but this was not verified. In order to obtain the fringe information for angles greater than 90° but less than 180° the fin would be exchanged for one containing a stainless steel reflective surface. The scene beam would then enter the test area from the viewing port below the tunnel and be reflected to the hologram plate as shown in Figure 12.

#### b. Total Model Flow Method

In this method the diffuser plate was located in the scene beam outside the test section as shown in Figure 14 and the flat plate



center section and fin were made of optically clear plexiglas. By translating the diffuser plate between exposures of the hologram, an interferogram of the whole density field in the test section about the model was obtained. From Figure 13 it can be seen that due to symmetry only 90° of view was required to obtain the density field. This makes it much easier experimentally to take the holograms than the previous method described.

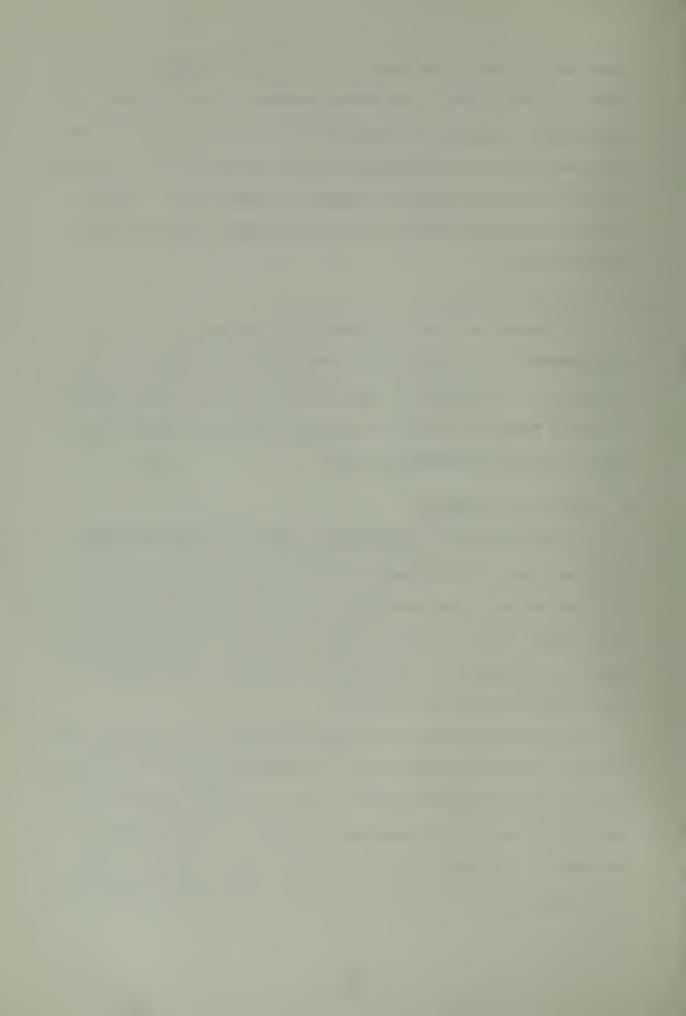
## 3. Schlieren Analysis

A standard Schlieren knife-edge system was used to verify the establishment of the supersonic flow network around the model as shown in Figure 9. Photographs were also taken of the flow with the model at  $0^{\circ}$  and  $90^{\circ}$  rotation in order to compare the fin shock conditions with those obtained by Winkelmann  $\{26, 27\}$ .

# B. PHOTOGRAPHIC TECHNIQUES

By illuminating the holograms with a helium-neon laser beam which has a wave length of 6328 Angstroms, the original scene was reconstructed. Since the original scene beam and the reconstructed beam were of different wave lengths, there is actually a small distortion in the reconstructed scene but of neglectable effect because the hologram plate emulsion in the development process also shrinks.

The typical method for reconstructing the scene is to illuminate the hologram as illustrated in Figure 15. The diffuse glass used in the construction of the hologram appears to act as an infinite light source of non-parallel rays which illuminates the scene. If a small aperture is positioned at the focal plane of the imaging lense, an almost parallel set of rays may then be selected as shown in Figure 16. A third method



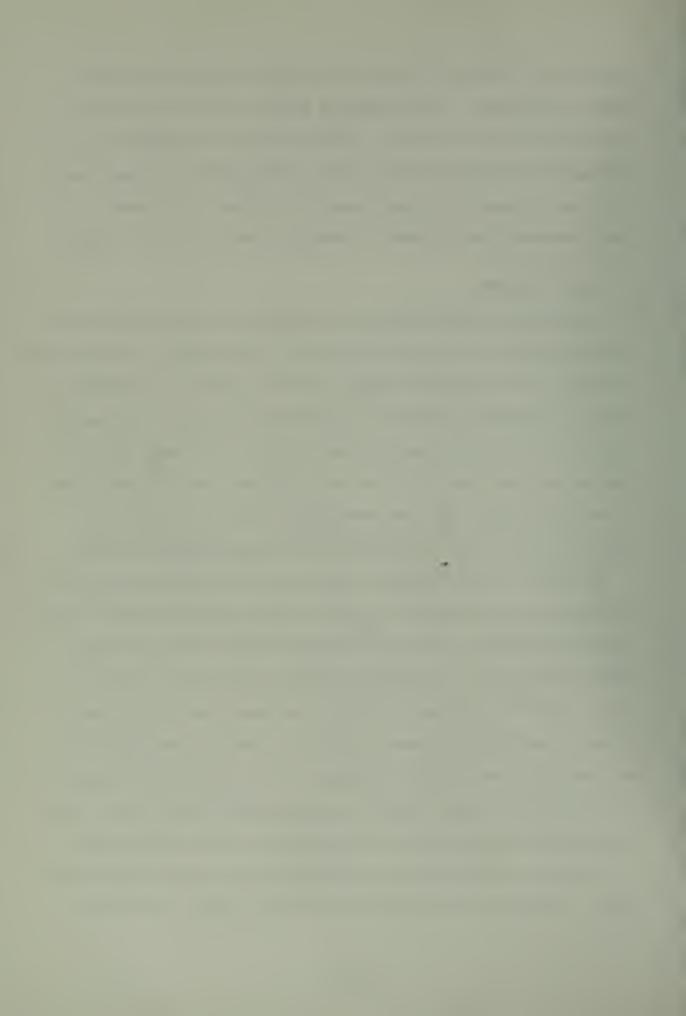
illustrated in Figure 17 uses a small diameter conjugate beam to illuminate the hologram. A large depth of field is achieved because the narrow beam acts as an aperture. This effect was of considerable advantage since it enabled both the front and rear grids, the model, and the fringe patterns to be simultaneously projected on the screen. The best photographs were obtained by focusing on the plane of the fringes.

### C. DATA REDUCTION

Photographic interferograms were obtained by illuminating the hologram scene with a thin laser beam and using a camera with a viewing screen located in the film plane as shown in Figure 15. The line of sight in the plane desired was achieved by translating the hologram until common points on the front and rear grids were aligned. The camera, with the aperture set wide open at 7.7, was then adjusted to give the best focus on the fringe plane. The best photographic results were achieved by using an exposure time of 1/10 second with Polaroid Type 55 P/N film.

It was felt that the density field could be well defined three-dimensionally along the fin if the density fields were determined in four planes perpendicular to the Z-axis and equally spaced along the fin as shown in Figure 18. In obtaining the fringe data across a constant Z-plane it would be necessary to take six photographs per rotation angle, aligned at appropriate intervals down the y' axis, and then graphically mate the fringe data to form one complete set. The six photos across the field were felt necessary because the optical path length from the model to the hologram plate varied for those points not on the aligned plane.

The fringe shift reduction was accomplished using two different techniques. The first was to project the negative, using a photo enlarger,



onto a sheet of paper and trace out the fringe pattern, model surfaces and grid lines. The light fringes were traced out since it was much easier to judge their center line. From the fringe lines forward of the fin in the region of uniform flow one fringe line which appeared the straightest and compatible with most others was selected as the bench mark. A straight line was then drawn over that fringe and extended past the aligned Z plane (i.e. y' axis). Lines parallel to the bench mark line were then drawn likewise over the remaining fringe lines. The fringe displacements were then read relative to the lines drawn at the points of intersection of the fringes with the y' axis. The radius of the inversion circle was selected so that the fin-root intersection was the origin and the fin tip was the 100 percent point.

In the second technique an enlarged positive photograph was made from the negative. Again one fringe line in the uniform flow region just forward of the fin which appeared to be parallel with the majority of the fringes was selected as the bench mark. The remaining fringes were likewise traced over with lines parallel to the first. The fringe displacements were then measured relative to the lines drawn. For further details see Appendix A.

The locations at which reference lines crossed the y' axis in both techniques above were further adjusted to account for the tunnel wall refraction displacement as shown in Figure 19 and computed by a computer program in Appendix B. Once these corrections were made, the radial variation of the fringe number could then be plotted for the various y' alignment planes and a smooth curve drawn through the data points. The fringe number at 201 equidistant points across the field can then be obtained for input into the computer program, HOLOFER, in MODE 3. For



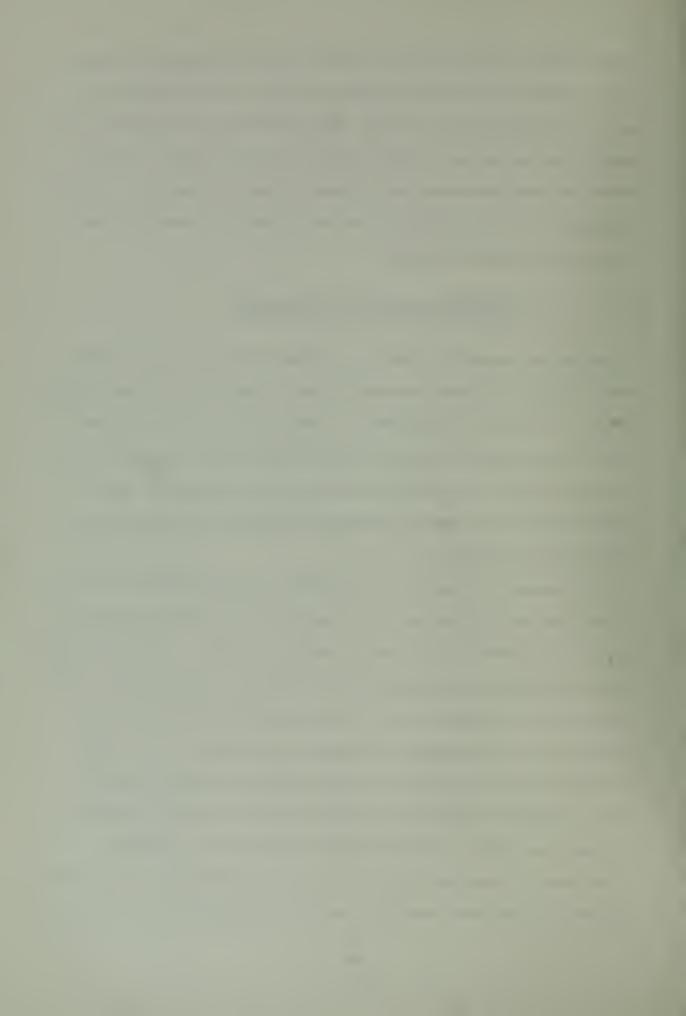
further details on how to use the computer program, HOLOFER, see Appendix C. Once the data from all the rotation angles of the model have been put into the computer program, the program will then calculate the density field across the inversion circle for that Z-plane. After the density has been calculated for all four Z-planes in Figure 18, a three-dimensional plot of the density field can be made by connecting points of equal density across the fin.

# V. EXPERIMENTAL RESULTS AND DISCUSSION

The initial attempts to establish uniform flow over the flat plate shown in Figure 7(a) were unsuccessful due to model vibration and flexure. Movement of the model sting within its holder and flexure of the model plastic center section allowed the model leading edge to establish a little over 1° angle of attack upward when flow was established. Due to various modifications made in attempts to eliminate the vibration, the sting finally fractured.

In an attempt to eliminate these problems, the second model shown in Figures 7(b) and 7(c) was made of a single piece of stainless steel and the sting was mated to its holder to within .001 inches. The model center section was made of poured epoxy and the strakes were both made of stainless steel with plexiglas inserts. The model rotation about the sting was reduced to approximately 0.3° angle of attack upward. Due to the vibration in passing through the transonic range and a weak glue seal between the metal strakes and plexiglas inserts, the inserts were found to break loose. They were subsequently removed and not replaced.

In order to determine the flow field using the direct fin-root approach, the epoxy fin and model center section were frosted on one side using fine



emery paper (see Figure 20). Holograms were taken of the model at rotation angles of  $0^{\circ}$ ,  $12^{\circ}$ ,  $45^{\circ}$  and  $90^{\circ}$  using the holographic arrangement in Figure 2 excluding the diffuser plate between lens,  $L_3$ , and the test section. The mirror,  $M_5$ , was translated in various directions from towards to parallel to the test section between the exposures without flow and with flow established. Fringe patterns were obtained around the model, but only at  $0^{\circ}$  and  $12^{\circ}$  rotation could any fringe patterns be observed across the fin. It was found that the fin fringe pattern appeared to remain almost unchanged no matter how or how much the mirror,  $M_5$ , was translated while the fringe around the model changed appropriately. For instance, in Figure 21 the mirror was not translated and in Figure 22 the mirror was translated .006 inches horizontally parallel to the tunnel.

Since fringes could not be observed across the flat plate at  $45^{\circ}$  and  $90^{\circ}$ , it was felt that the epoxy center section might be too imperfect optically. Consequently double-exposed holograms were taken with no flow through the test section and various translation distances from 0 to .005 inches. The fringe patterns were excellent across the whole model and their spacing decreased according to the increase in the translation of  $M_5$ . Next the diffuser plate in the scene beam in Figure 2 was inserted with the model at  $0^{\circ}$  rotation angle and translated between no flow and flow exposures of the hologram. The fringes about the model were of excellent quality but the double diffusion of the scene beam through the model caused all fringe patterns on the model (fin) to disappear.

It was felt at this point that the fringe pattern obtained across the fin was caused by the movement of the model to an angle of attack and possibly by model vibration, although none was observed visually. The



lack of any fringes across the flat plate center section is not well understood but is believed to be caused by vibration of the model.

Since it was not possible to obtain acceptable interferograms with the diffuser plate as part of the model due to model motion, it was felt that the effect of minor model movements could be eliminated by using an optically clear model and an external translating diffuser plate. Therefore the fin and model center sections were replaced with optically clear plexiglas. With these changes it was found that the flat-plate leading-edge angle of attack had been reduced to approximately 0.1°.

Initially double-exposure holograms were taken using the arrangement in Figure 2 and translating the diffuser horizontally. At 0° model rotation the holographic interferograms were excellent. But as observed in Figure 23 it would be extremely difficult to determine the fringe change across the fin since no free stream reference fringes were available, due to the flat plate leading edge Prandtl-Meyer expansion and fin shock intersecting the tunnel top just above the fin. With a larger tunnel or smaller model this would be an excellent technique.

The diffuser plate was then translated vertically between exposures and excellent horizontal fringe patterns were obtained as shown in Figure 24. The model was then rotated to  $22\frac{10}{2}$  and the same holographic technique was used. Excellent fringe patterns were obtained above and below the flat plate center section. Fringe patterns across the flat plate center section and fin in this area were very light and usable interferograms could not be photographed with the polaroid camera. In an attempt to improve on the fringe quality, polarizer plates were inserted in the reference beam between the mirror,  $M_4$ , and the hologram plate and in the scene beam between the lens,  $L_3$ , and the diffuser plate in order to ensure

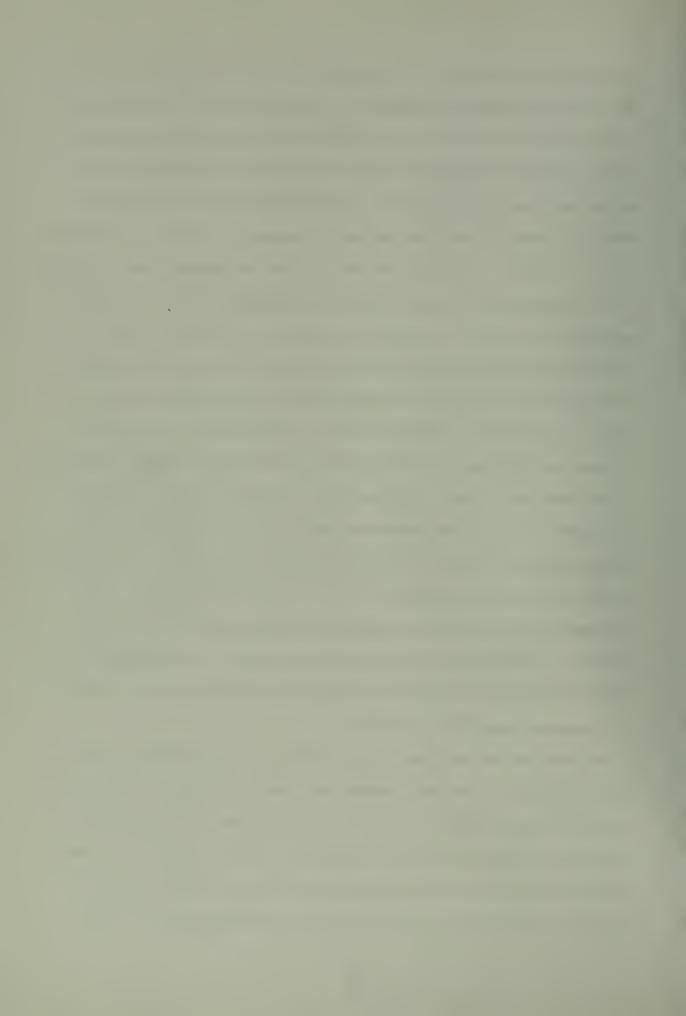


polarization of both beams. No significant improvement could be noticed and they were subsequently removed. A one-quarter wave polarizer plate was then placed between the lens,  $L_1$  and the beam splitter in order to utilize circularly polarized light for the reference and scene beams. Okayama and Emori [14] found that their image resolution improved considerably; however, with this particular arrangement little to no improvement in the fringe resolution was observed and the approach was abandoned.

The transverse mode selector was then varied from 2.0 mm. to 1.5 mm. and later to 1.0 mm. in an effort to improve the coherency length of the laser light and consequently the image resolution. Due to the decrease in output light intensity, up to six exposures were taken during a run with flow established. Image and fringe definition were not found to increase possibly due to model vibration which was not visible to the eye.

The model was rotated to  $45^{\circ}$  and  $67^{\circ}_{2}^{\circ}$  and double exposure holograms were taken with a 2.5 mm. transverse mode selector. There were no observable fringe patterns in any portion of the test section. Consequently it was believed that supersonic flow was not established due to tunnel blockage caused by the shock wave from the model and by slight model vibrations in passing through the transonic range. In the transition to supersonic flow, the model leading edge would sometimes flex as much as  $0.2^{\circ}$  depending upon the transition time.

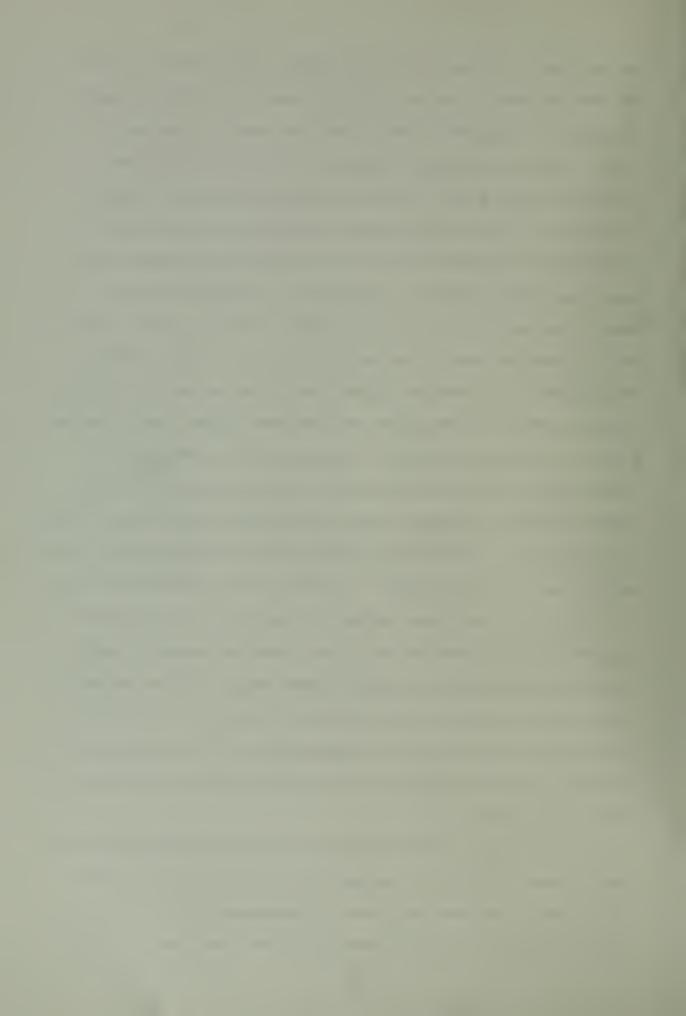
The test section walls were changed from the total plexiglas side walls shown in Figure 1(a) to the aluminum walls with the optical quality glass port holes (Figure 1(b)). The better quality glass would hopefully improve the viewing and the port holes made the model much more accessible. The metal strakes were also removed from the model since they appeared to have a minimal effect on the flow, were an interference optically,



and appeared to add very little structurally. Flow tests were run and the model was found to attain an angle of attack of about 0.14° upward when flow was established at Mach 2.84. The angle of attack was tolerable, however the model also maintained a visible high frequency vibration of about + 0.05°. Double exposure holograms were taken with the model at 0° rotation and fringes could only be observed forward of the Prandtl-Meyer expansion on the top of the model and forward of the flat plate leading edge shock on the bottom. The tunnel seals were found to leak and were replaced and resealed. The same angle of attack and high frequency model vibration still persisted. Double exposure holograms were again taken and a very faint fringe pattern was observed across the model. In reproducing the scene when aligned along a constant Z-plane passing through the fin, it was impossible to photograph the fringe pattern across the fin because of a bad speckle effect. Double exposure holograms were made at model rotation angles of 45° and 90° without and with flow. The holograms taken without flow gave excellent fringe patterns across the model however no fringes could be observed with flow.

At this point a knife edge Schlieren system was set up to determine just what the flow conditions were in the tunnel and around the model. The first runs were made at a model rotation angle of  $90^{\circ}$  and presented the shock knee in front of the fin (Figures 25(a) and 25 (b)) which was pointed out by Thomas [23, 24] and Winkelman [26, 27]. Due to the model vibration, the shock extending out from the shock knee was observed to flutter about  $\pm \frac{1}{2}$  cm.

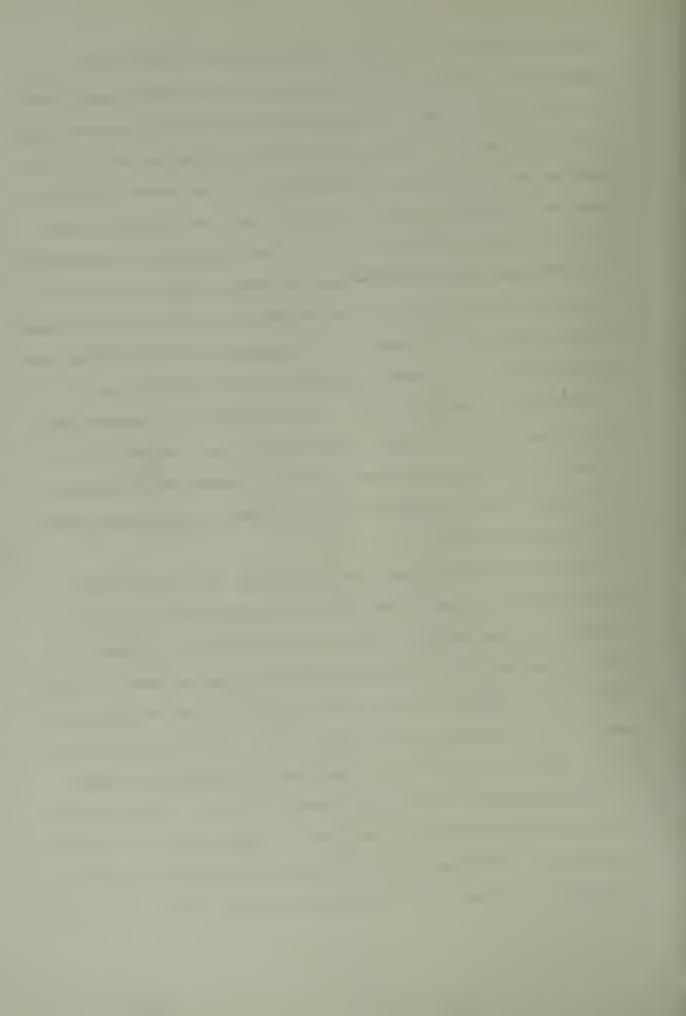
The model was then rotated back to  $0^{\circ}$  and Schlieren photographs were taken as shown in Figure 26. During all but one of the runs at this rotation angle, the model would pitch up approximately  $0.14^{\circ}$  angle of attack and continue to vibrate about  $\pm~0.05^{\circ}$  at a high frequency when



flow was established. The plate leading edge shock and fin shock appeared quite fuzzy and light. Due to oil from the tunnel supply reservoir mixing in the flow, a light oil smear pattern can be observed across the fin in Figure 26. During one run, which could not be duplicated, the model pitched up as flow was established but did not vibrate. The plate shock and Prandtl-Meyer waves, fin shock system, and tunnel Mach lines became very distinct and well defined as seen in Figure 27. The schematic of the Schlieren photographs in Figure 28 points out the cause for the various flow lines observed. The non-uniformity of the free stream caused by tunnel leakage can be easily seen. Consequently the experimental work was discontinued due to tunnel conditions and time considerations.

It was felt at this point that if the holographic interferogram taken of the clear plexiglas fin at  $0^{\circ}$  rotation angle could be reduced to useful data for the computer program then the holographic method would be to a certain extent verified even though the actual density field could not yet be determined.

In obtaining interferograms from the hologram, the reconstruction technique shown in Figure 15 was used. The plane of constant Z across the model to be reduced was chosen to be the forward most vertical grid line crossing the fin. It should be pointed out that in order to simplify the hologram alignment process, all four reduction planes across the model should have been scribed on the exterior grids. The three photographic points across the Z plane on the y' axis were, for convenience, chosen to be where horizontal exterior grid lines crossed the y' axis on the fin. Five interferogram photographs were taken at points A, B, and C as shown in Figure 29. Photographs 2 and 5 were taken using a Kodak Wratten Gelatin Filter N.D. 2.0 placed in the reconstruction beam in order to reduce the



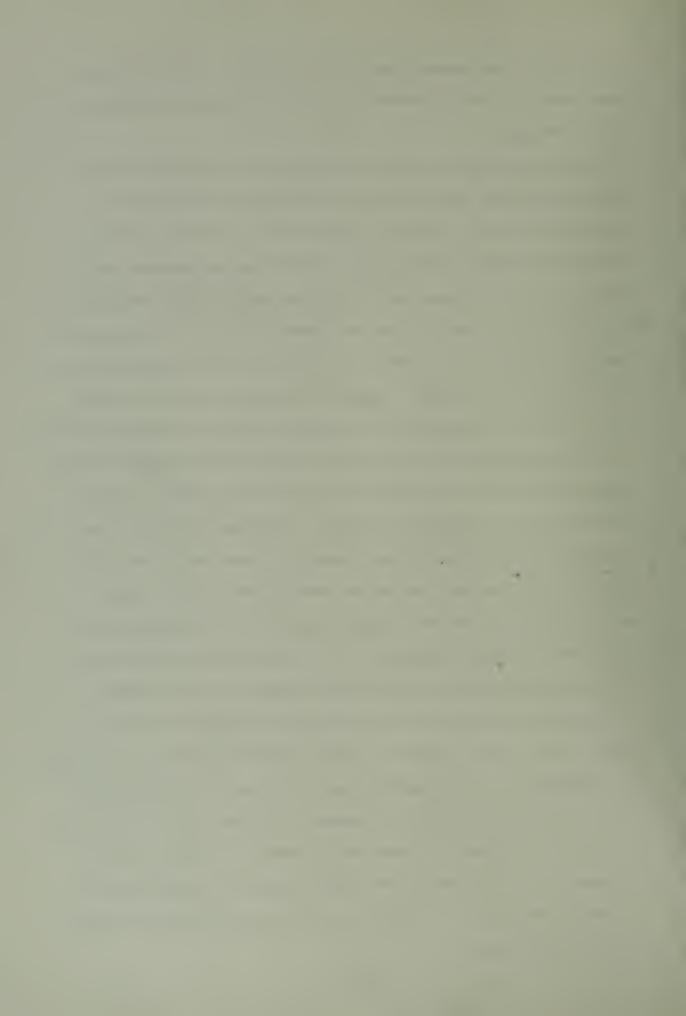
beam intensity and increase the fringe definition. These two photographs were also used to provide a check on the consistency of the reduction process.

All five negatives were blown up on the photo enlarger in the dark room and the plate, fin, grid lines and fringes were traced out on a sheet of white paper as shown in Figures 30-34. It was very difficult to trace the fringes in the region of the fin tip and fin root due to the photographic resolution and the fin tip shadow. Also fringes were not visible in the region of the fin leading edge shock. Consequently, connecting the fringe in front of the fin to the correct fringe on the fin was a best guess effort. Typical of the problem was that fringe lines in two of the drawings were initially improperly connected across the fin leading edge shock. After being checked against the photograph negatives, the fringes were reconnected and the data taken correlated well with the data from other drawings. The fringes in the fin root area were extremely light in some photographs which made it quite difficult to determine their centerline crossing the y' axis. Another point of difficulty was determing the location of the top and bottom of the fin. An error in drawing here will have an effect later when the fringe locations are normalized with respect to the fin height.

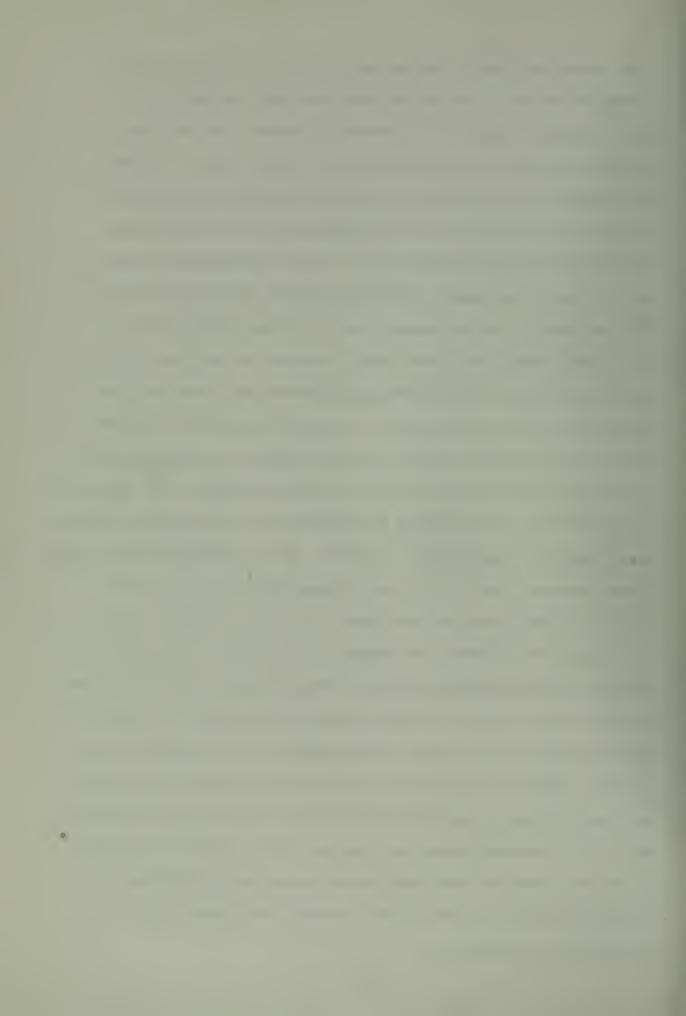
Enlarged positives were then made from the negatives as shown in

Figures 35-39 to see if the data accuracy could be improved by eliminating
the difficulties of tracing the fringes in the dark room. This method
also made it possible usually to reassess the assumed path of the fringe
lines at a later time if the data did no appear to correlate properly.

The same problems of locating the fringe lines in the region of the fin
tip and of connecting the fringe lines across the fin shock are readily
apparent in the figures.



To obtain the fringe change across the fin one of the straightest fringes in the free stream region which paralleled the majority of other free stream fringes was selected. A straight line was drawn through its centerline and extended to cross the y' axis. The other free stream reference lines were then drawn parallel to the first and adjusted so as to best follow the centerline of the selected fringe. This method actually averages the free stream conditions and is only valid if the free stream is essentially uniform. Since it was not possible to obtain in the photographs the free stream fringe pattern forward of the leading edge Prandtl-Meyer expansion for the lower fin region, the fringe reference lines in photo enlargements were drawn over the fringes just prior to the fin shock. In order to adjust them to free stream conditions their location was moved downward a distance equal to 1.1 times the average fringe interval for that photograph. The figure 1.1 was observed in all the enlarged photographs to be the approximate fringe change across the Prandtl-Meyer expansion. In the drawings (Figures 30-34), fringe reference lines for the lower fringes (generally y' & .8) were initially aligned along the straightest portion of the fringe prior to intercepting the fin shock. By comparing the reference line location on the drawing with the fringe pattern in the photographic enlargements the reference line location and fringe change were adjusted by an appropriate portion or all of the 1.1 fringe change caused by the Prandtl-Meyer expansion. All of the reference lines were then corrected for the tunnel wall and grid plastic parrallax shown in Figure 19 and computed in Appendix B. The reference fringe locations were then normalized with respect to the wing height and the fringe numbers calculated by dividing the fringe change by the average fringe interval. For further details and calculations see Appendix A.



The data obtained for alignment points A and C by the two different reduction processes are plotted in Figures 40-43. The data obtained from the photo enlargements aligned at point C and shown in Figure 43 gave the best data agreement between two photographs. The worst data agreement was obtained from the reduction of the drawings aligned at the same point. The inconsistency in the data was probably caused by connecting the wrong fringe lines across the fin leading edge shock. The data fluctuations and discrepencies between the curves in the figures could have been caused by a number of things. It could have been caused, for instance, by not drawing the fringe reference line parallel to or exactly on the free stream fringe center line or slightly missing the fringe center as it crosses the y' axis or by misconnecting fringe lines across the fin shock as was illustrated in Figure 42. Fringe location errors could be caused by misdrawing the fin tip and root lines as was mentioned earlier. Another contributer would be measurement errors.

To analyze these sources for error, first consider that the typical fin size in the drawings and photo enlargements averaged about 2.5 inches high and 2.75 inches wide and that the fringe spacing averaged about 0.10 inches. All measurements were taken using a ruler graduated in 0.01 inches and readings were made to the nearest .005 inches. Since the average difference between the data points and curves ran around ½ fringe, some figures were calculated to determine what measurement errors could produce this fringe error. It was found that an error of .025 inches in alignment of the fringe reference line with the free stream fringe center line and/or fringe center line crossing the y' axis could produce ¼ fringe error. This fringe error will also occur if the fringe reference line differs from the free stream center line by more than 1.4° when drawn 1 inch from the y' axis or by more than 0.36° when drawn 4 inches from the



y' axis. A difference of .01 inches in the average fringe interval could also produce a  $\frac{1}{4}$ -fringe error; however, this is not too likely since it is an average of fifteen to twenty-five intervals. By comparing the actual height-to-width ratio with those found in all the drawings and photo enlargements it was found that average error was around 2% or a distance of .02 on the y'-axis.

In order to compare the interferogram negative quality and the two different reduction techniques, a plot of the data for each negative was made as shown in Figures 44-48. Photographs 1 and 3 in Figures 44 and 46 gave the smoothest curves indicating the highest interferogram resolution. Photograph 5 (Figure 48) gave the worst dispersion indicating poor interferogram resolution; yet looking at the enlargement in Figure 39, the fringe line contrast is very good. In comparing Figures 40-48 it appears that the reduction technique using the photographic enlargements gave the most consistent data. This technique also provides a much easier and faster recheck on the proper tracing of fringe lines because it is extremely difficult and tedious to duplicate the fringe pattern to the same scale over the drawings using the photo-enlarger in the dark room.

All the data obtained by either method for one alignment point were then plotted in Figures 49 and 50 in order to observe the dispersion.

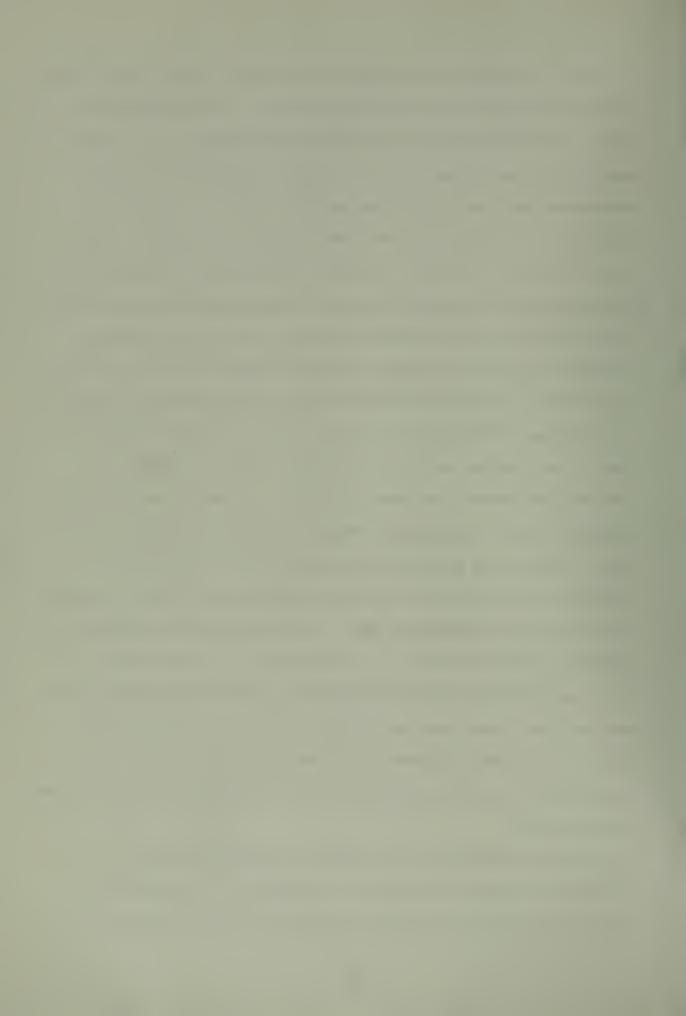
A fringe number dispersion of about 0.6 fringes was observed in the data taken from the negatives aligned at point A and a dispersion of about 0.8 fringes for the data about point C. The dispersion is attributable to the inaccuracies in the drawings, to the photographic quality of the interferograms, and to the inaccuracies in correcting the lower fringe reference lines to free stream conditions. The last point is based upon the increased dispersion between the fin tip and fin root data.



Figures 51 and 52 show an integrated curve of the fringe change across the wing as determined by each reduction method. In constructing the curve, the data from the three aligned points was plotted so as to just overlap each other. These two curves were then compared in Figure 53. The maximum variation in the fringe number is about ½ a fringe but the variations in the location of the fringe maximums and minimums average about 0.1 inches on the fin. The location difference is probably due to improper drawing of the fringe reference lines compounded with not being able to measure the wing height accurately. Due to the inaccuracies introduced in tracing the negative and then reducing the data it is felt that the photo enlargement method is the more accurate reduction method.

With fringe data from only one field of view, the density field, obviously, could not be obtained. It was felt useful to consider the flow field axisymmetric in order to exercise the computer program and to provide a check on the program's ability to handle these particular curve shapes. Fringe data at 101 equidistant points across the fin from  $0 \le Y' \le 1.0$  was obtained from Figure 53 for both curves and fed in HOLOFER in Mode 3 for the axisymmetric case. For further details on HOLOFER see Appendix C. The scale factor,  $\infty$ , in Equation 9 was then varied from 0.2 to 2.5 and a value to 1.0 was determined to yield the most accurate density solutions. This value was verified by feeding the function data,  $( ?_{p_{\infty}} -1)$ , calculated by Mode 3 back into the program in Mode 1 and comparing the fringe data calculated with the original fringe data obtained from Figure 53.

The density distribution for both the drawing and photographic reduction cases is plotted in Figure 54. The density as Y' approaches zero actually goes as the drawn lines even though the points indicate a dip.



Matulka [12, 13] pointed out that the computer program accuracy does not converge at the origin. The large variations of the density curves at values of Y' > 0.8 are caused by the program trying to adapt to a step or shock wave type function at Y' = 1.0. If the remainder of the fringe data in Figure 53 for values of Y' > 1.0 had been included as input data the density curves would have smoothed out. This shock wave step function effect was demonstrated and analyzed by Matulka [12, 13].

The low values of density for the photographic data curve around Y' = 0.38 were unexpected but not surprising. First, the photographic fringe data curve is more extreme than the drawing fringe data curve and second, the density curves are not true values anyway since the field was considered axisymmetric and this is not the case in reality.

In general the density curves are felt to be reasonable under the assumed conditions. Consequently it is believed that the computer program could very easily and accurately handle a complete analysis of the flow field around the fin-flat plate.

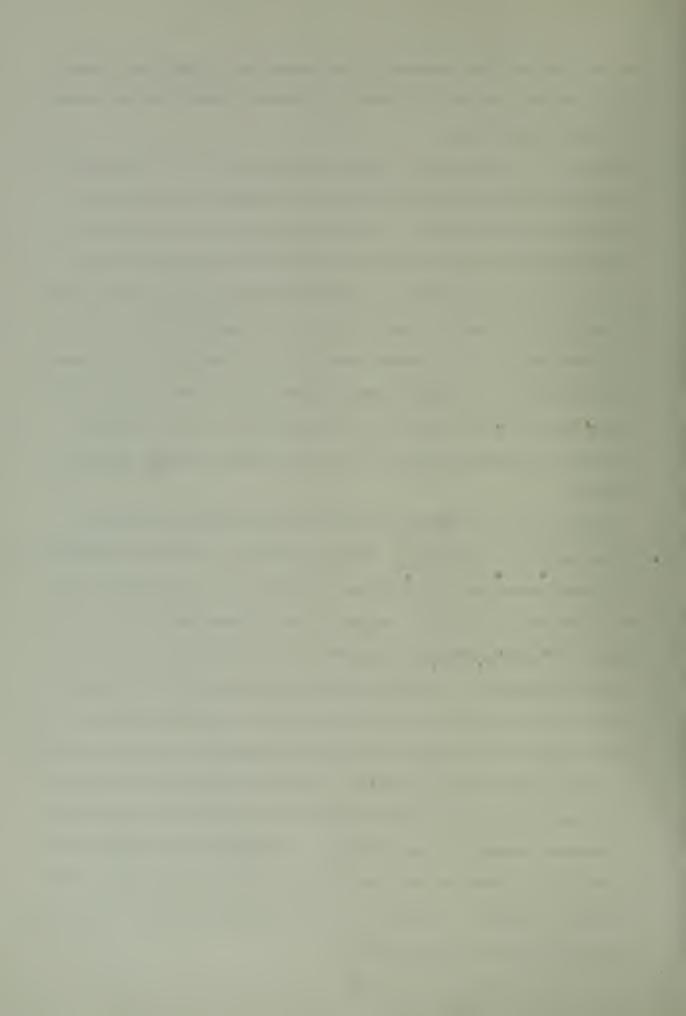
In completing the analysis there are some data reduction problems which would have been encountered in reducing the interferograms at other model rotation angles which merit discussion. With no model rotation, adjusting the fringe reference lines close to the fin root to account for the leading edge Prandtl-Meyer expansion was relatively easy. However, when the model is rotated, the fringe shift across the Prandtl-Meyer expansion can no longer be considered a constant and it will be very difficult to adjust the fringe lines at lower values of Y' to free stream conditions. The best solution would be to increase the scene beam diameter and/or reduce the model size in order to photograph the free stream fringe lines forward of the plate leading edge. The free



stream lines could be connected to the appropriate fringe lines forward of the fin and this would eliminate the numerical correction and increase the fringe data accuracy.

Also at the rotated angles, fringe information will not be available for portions of the reduction plane due to shadows cast by the model sides as shown in Figure 55. In the lower portion of the hologram the fringe curve can be connected with a smooth line because the density field in that portion should be essentially constant. Care must be taken in completing the curve, though, since this information will be used in the integration of other rotation angles. The fringe curve in the upper portion must also be completed carefully but should be easier since the shadow will be smaller. For this particular model it was calculated that for angles greater than 46.6° the upper shadow would not penetrate the fin.

There are two more problems to be coped with which do not have apparent solutions at this time. The first has to do with the superimposing of fringe information from different locations onto one scene beam line as illustrated in Figure 56. In case I the superimposing of the fringe change in Region A onto that in Region C and locating the fringe reference line at point C' is tolerable because the density field in Region A should be fairly constant and uniform. However in case II where the fringe change in all three regions are superimposed and located at point C', the answer is not readily apparent. If the plate and fin were of equal thicknesses then points A' and C' would be the same and the error would be somewhat reduced. If, in addition, the thicknesses were made as thin as structurally possible, the error would be reduced to a minimum. Also it might be possible to integrate the model geometry into the computer program but this has not been attempted.



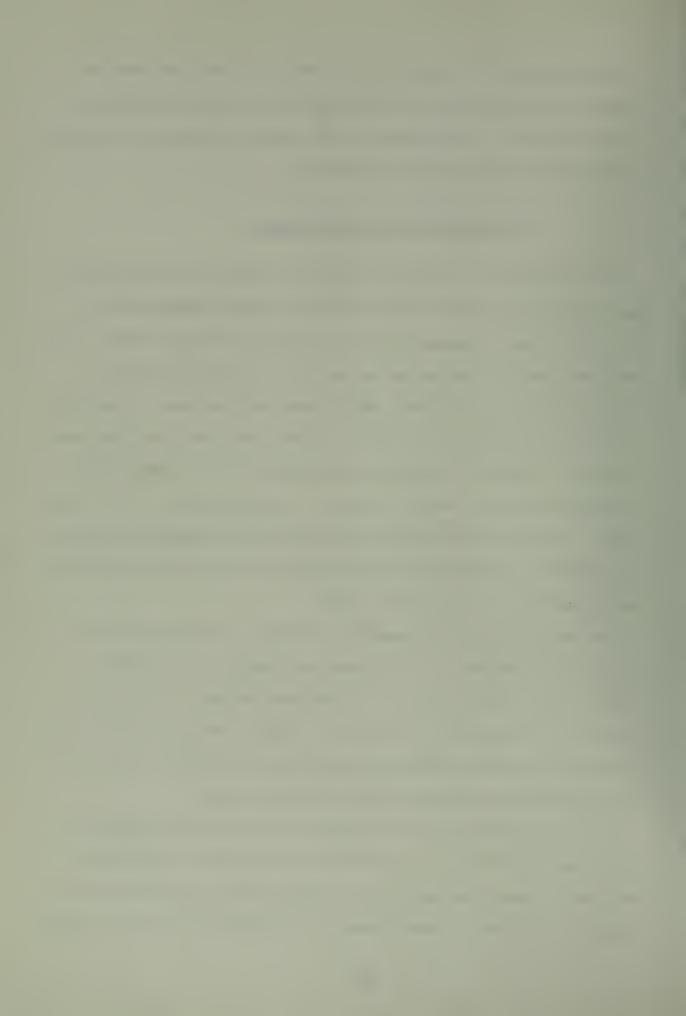
With the model at rotation angles between 0° and 90° scene beam defraction at the fin root and tip areas would also present problems as seen in Figure 57. A minimization of the problem could again be achieved by minimizing the fin and plate thicknesses.

## VI. CONCLUSIONS AND RECOMMENDATIONS

The investigation, although not totally successful, has demonstrated the feasibility of using holographic interferometry to determine the flow field around a transparent model by looking through the model. The problems of model vibration and movement, of superimposing different fringe information on one beam, and of scene beam divergence through the fin root and tip regions must still be solved. The model vibrations and rotation to an angle of attack are attributable to the tunnel pressure fluctuations caused by tunnel leaks and to possible movement of the sting holder. It is felt that all three problems could be decreased to neglectable effects by reducing the overall model size and by incorporating the model geometry into the computer program.

The basic holographic arrangement was found to work quite well for this type of experiment. The holograms were generally high quality except when unfavorable tunnel flow conditions existed. The use of circularly polarized light recommended by Okayama and Emori [14] did not increase the hologram resolution appreciably but the method merits further considerations because of their excellent results.

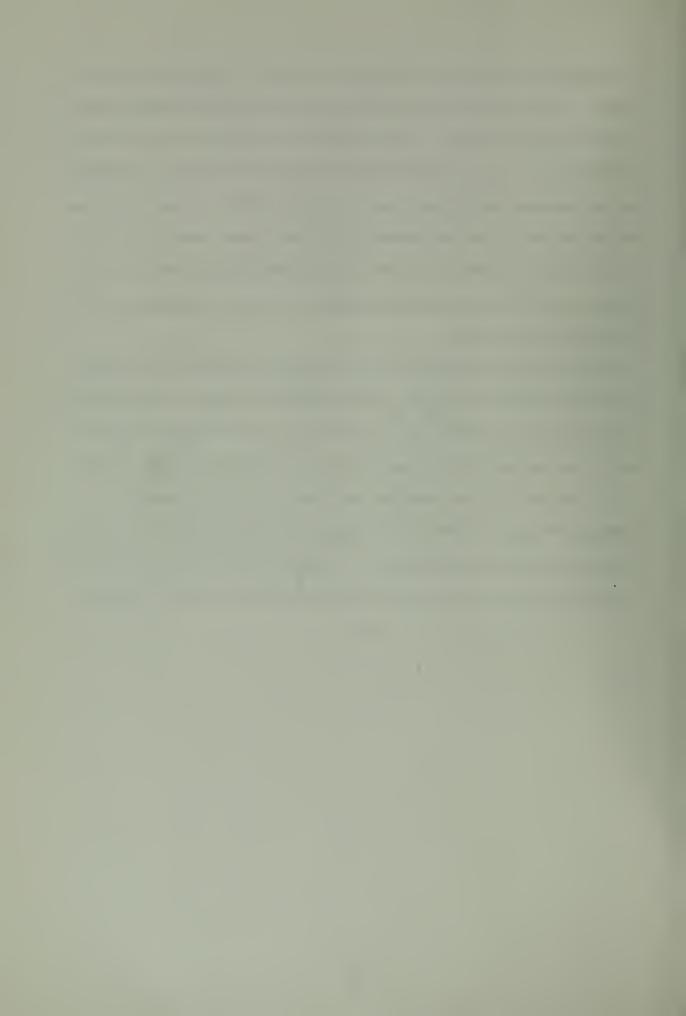
The data reduction process was found to be the rate-controlling step in the investigation due to the time and labor involved. Reducing the data from enlarged photographs of the interferogram saved some time and appeared to increase the data accuracy. The data scatter of  $\pm$  1/8 fringe



was considered acceptable considering the fringe resolution in the holograms. It was felt that this could be reduced by using either a larger tunnel or a smaller model. With either of the changes it would be possible either to use the free stream fringe pattern forward of the model as the reference conditions across the whole model or to use a vertical fringe pattern, since the leading edge Prandtl-Meyer expansion and fin shock would not block out the free stream fringes above the fin. It appears that the use of vertical fringes would also considerably reduce the data reduction time.

With the use of Schlieren the flow network described by both Thomas  $\begin{bmatrix} 23, 24 \end{bmatrix}$  and Winkelmann  $\begin{bmatrix} 26, 27 \end{bmatrix}$  was verified to exist. From the top view (90° model rotation) the fin shock and its fluctuations were observed and photographed through the flat plate plastic center section.

The computer program, HOLOFER, was found to be quite capable of handling the type of flow field fringe data which would be generated in a complete analysis of this type. As pointed out before it is believed that the program should be modified to incorporate the model geometry.



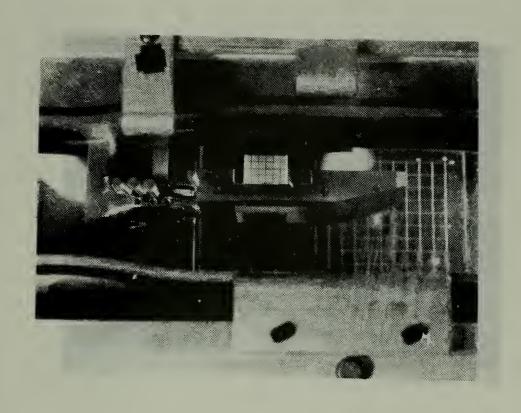


Figure 1(a). Wind Tunnel Test Section with Clear Plastic Side Walls



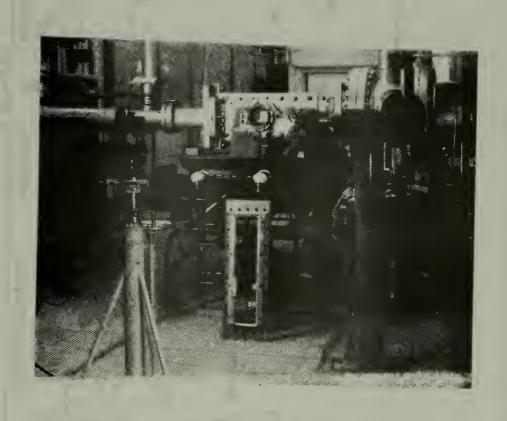
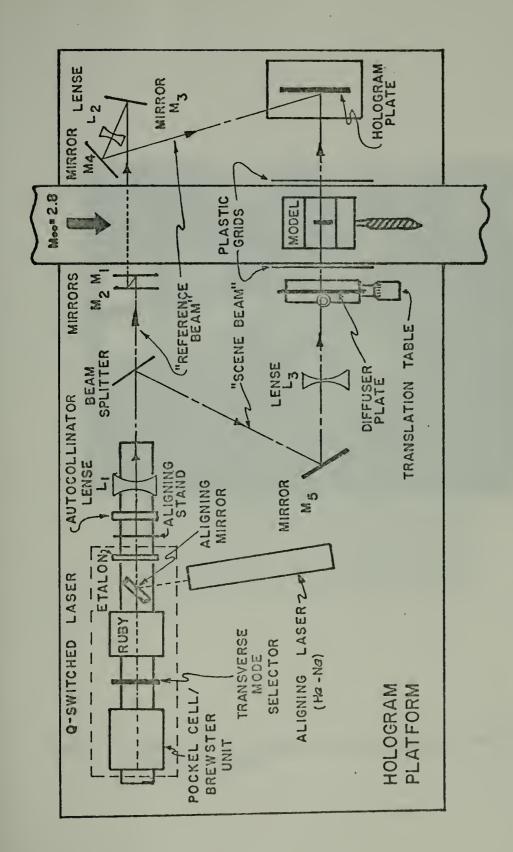
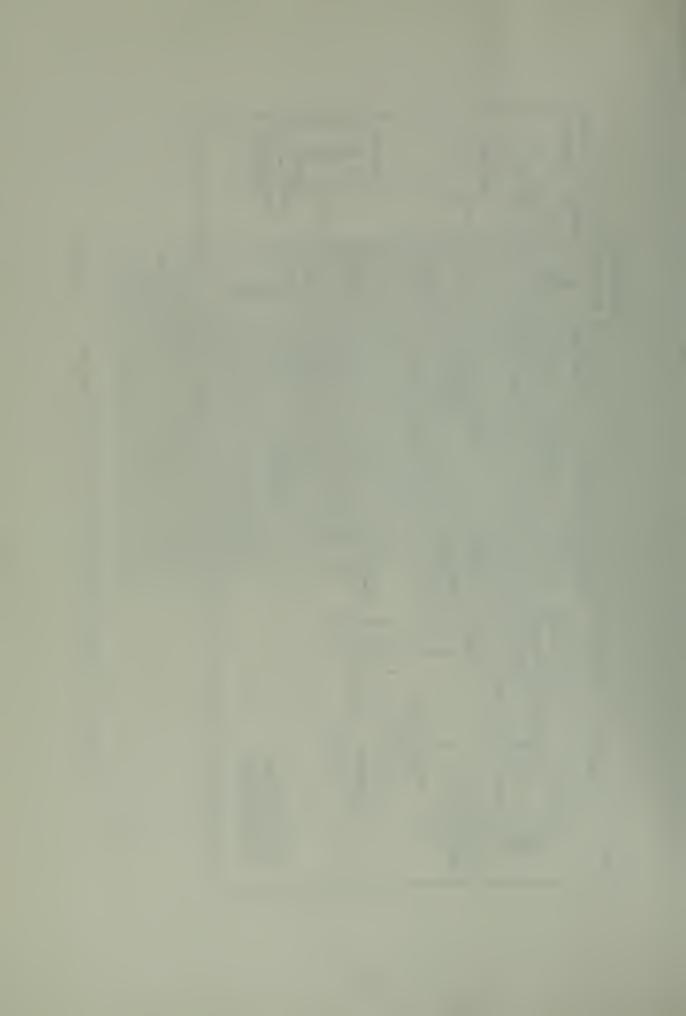


Figure 1(b). Wind Tunnel Test Section with Aluminum Side Walls





Schematic Representation of the Holographic Arrangement Figure 2.



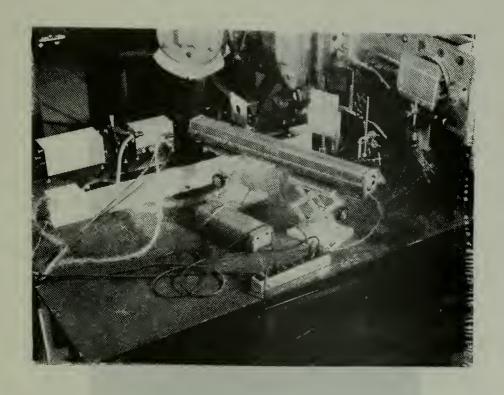


Figure 3. Holographic Arrangement Including the Laser Cooling and Aligning Equipment



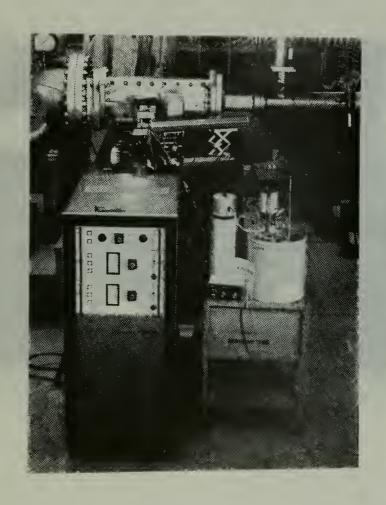


Figure 4. Ruby Laser Power Supply and Cooling Unit



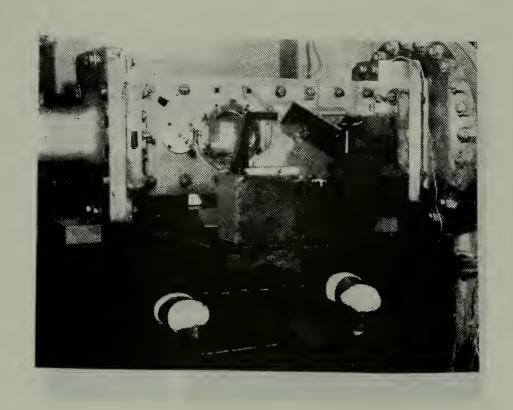


Figure 5. Hologram Plate Holder and Mirrors on the Reverse Side of the Wind Tunnel



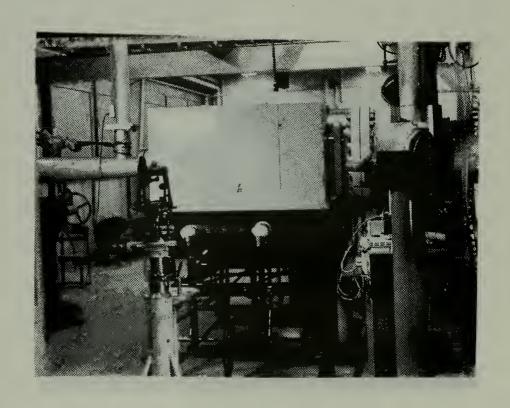


Figure 6. Hologram Platform Box Cover for Daylight Photography



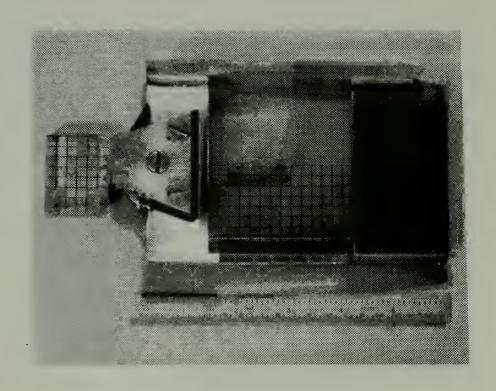


Figure 7(a). Initial Fin-Flat Plate Model Used in the Experiment



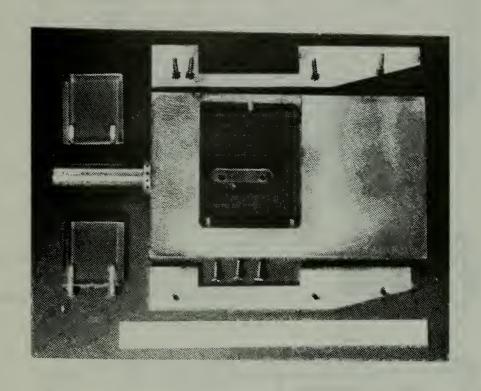


Figure 7(b). Second Fin-Flat Plate Model Used in the Experiment



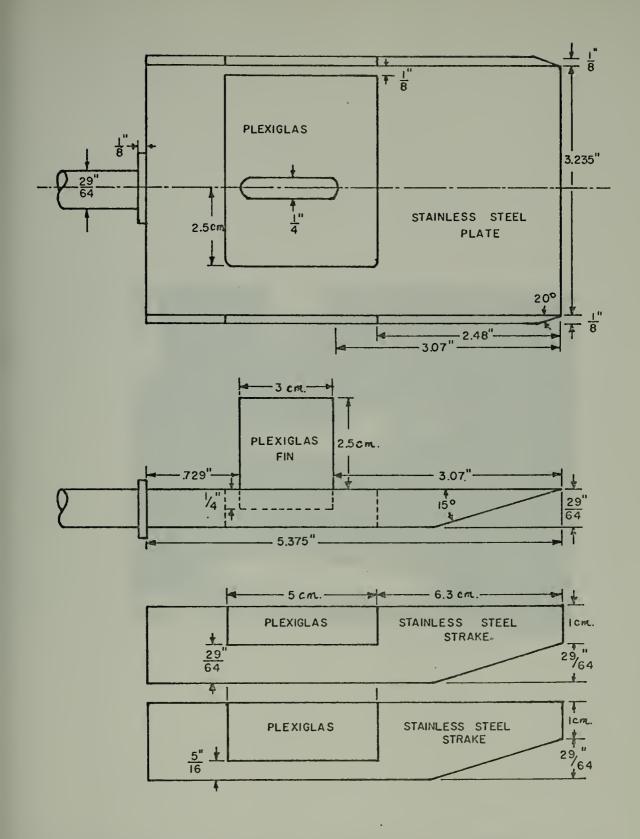
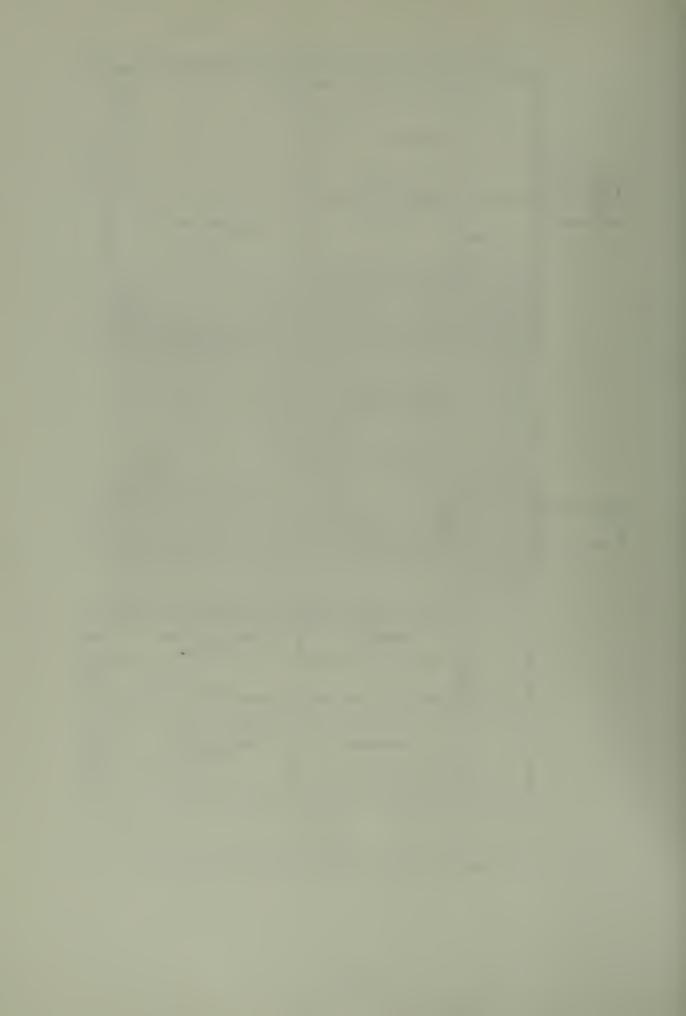


Figure 7(c). Details of the Second Fin-Flat Plate Model



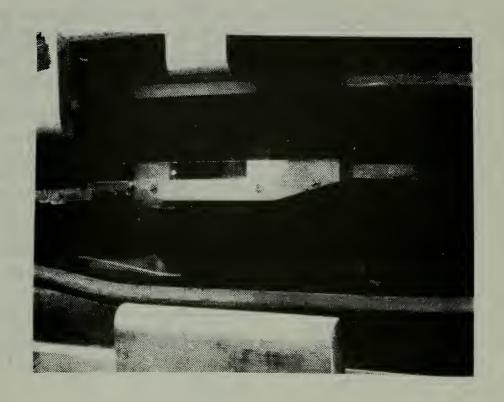


Figure 7(d). Model Mounting in the Wind Tunnel Test Section



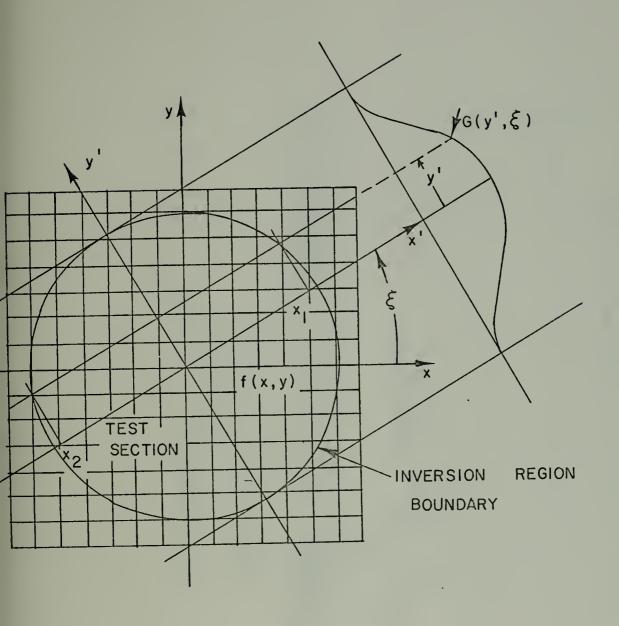
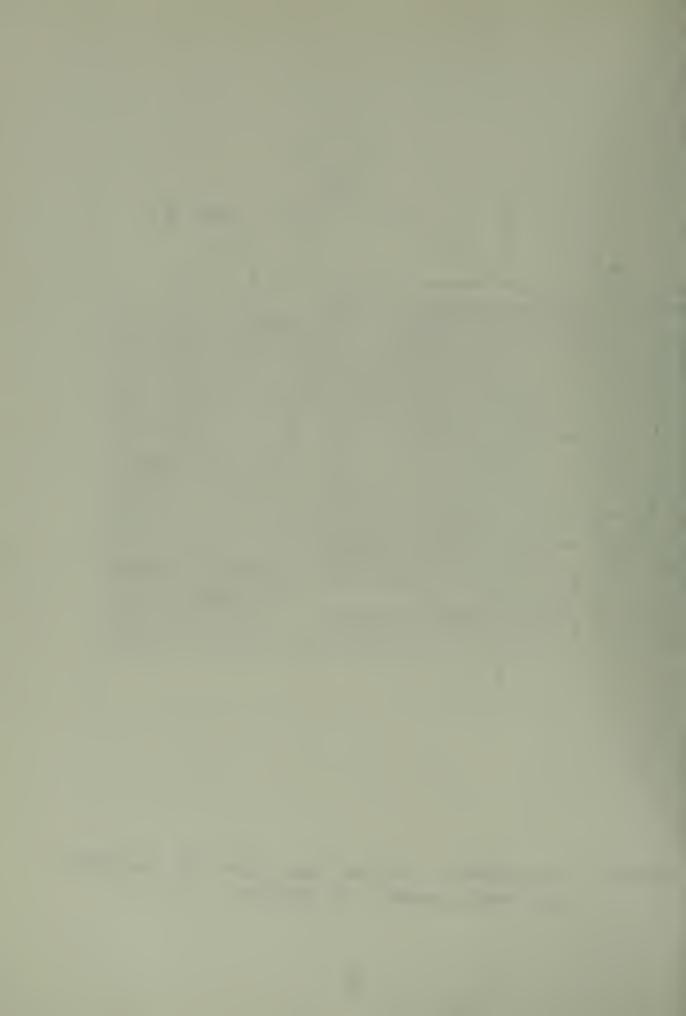


FIGURE 8. CO-ORDINATE SYSTEM USED FOR THE INVERSION OF FRINGE NUMBER TO DENSITY



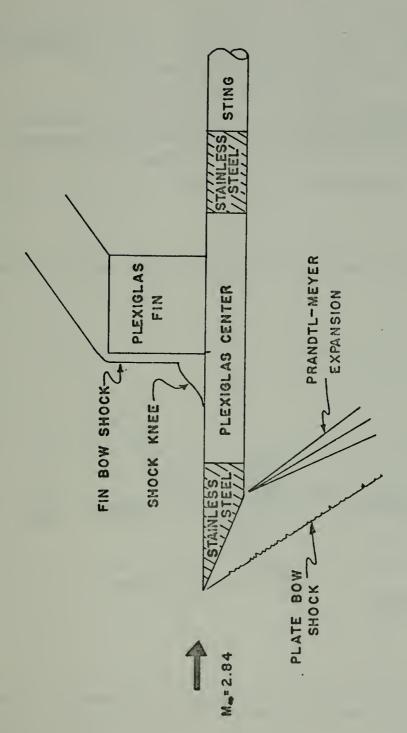
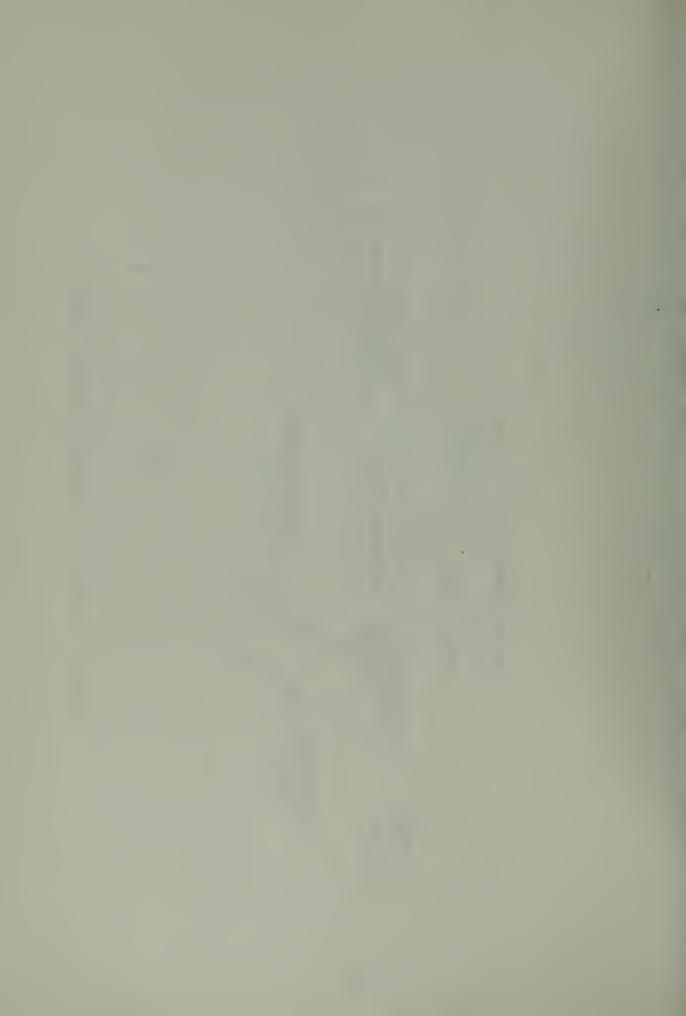


Figure 9. Schematic of the Desircd Model Flow Network



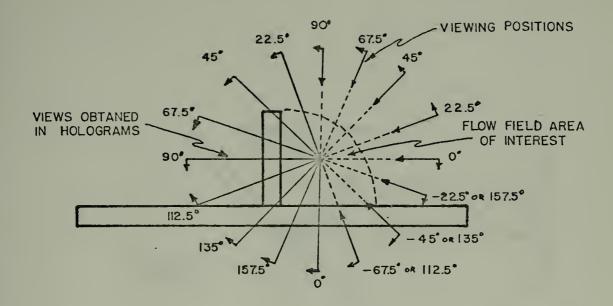


Figure 10. Desired Holographic Views in the Direct Fin-Root Method

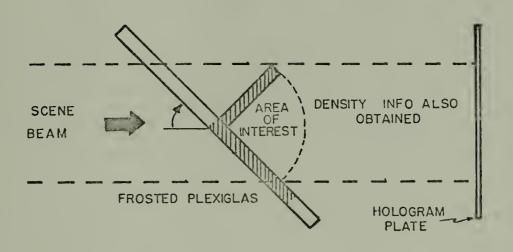
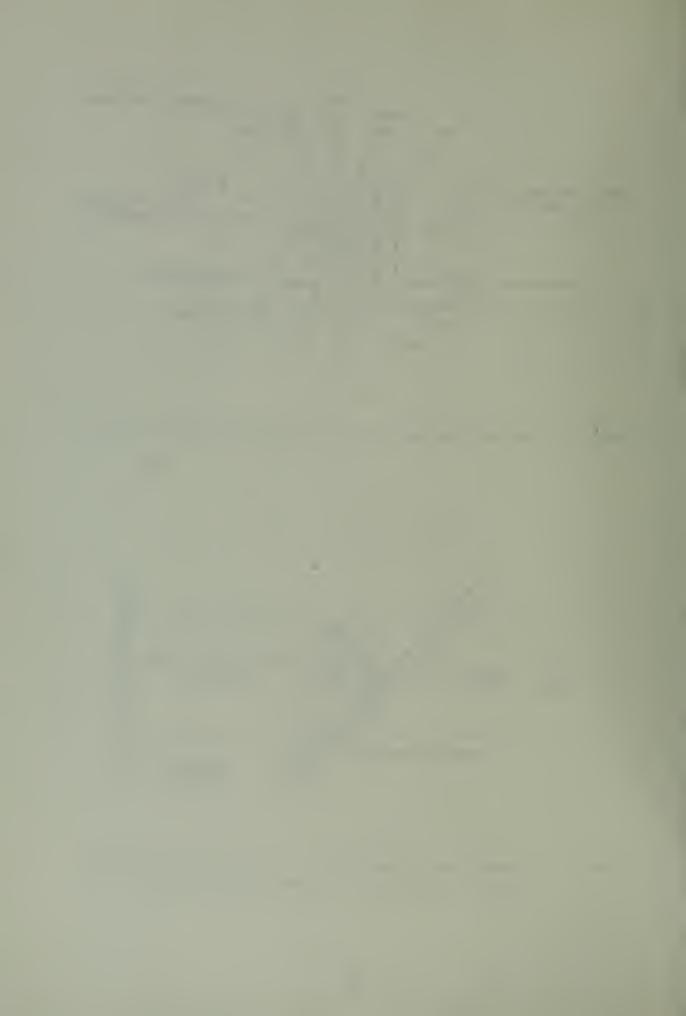


Figure 11. schematic of the Model Used to Obtain Polographic Views Between 0° and 90° in the Direct Fin-Poot Method



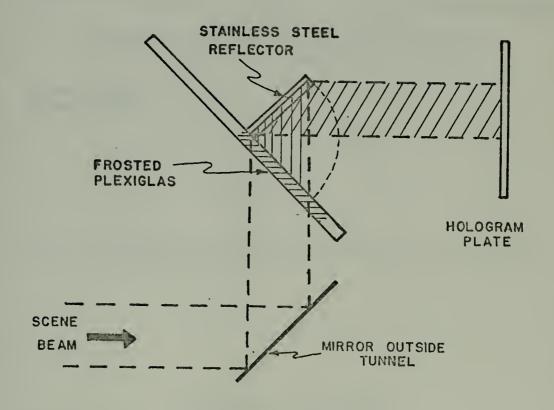
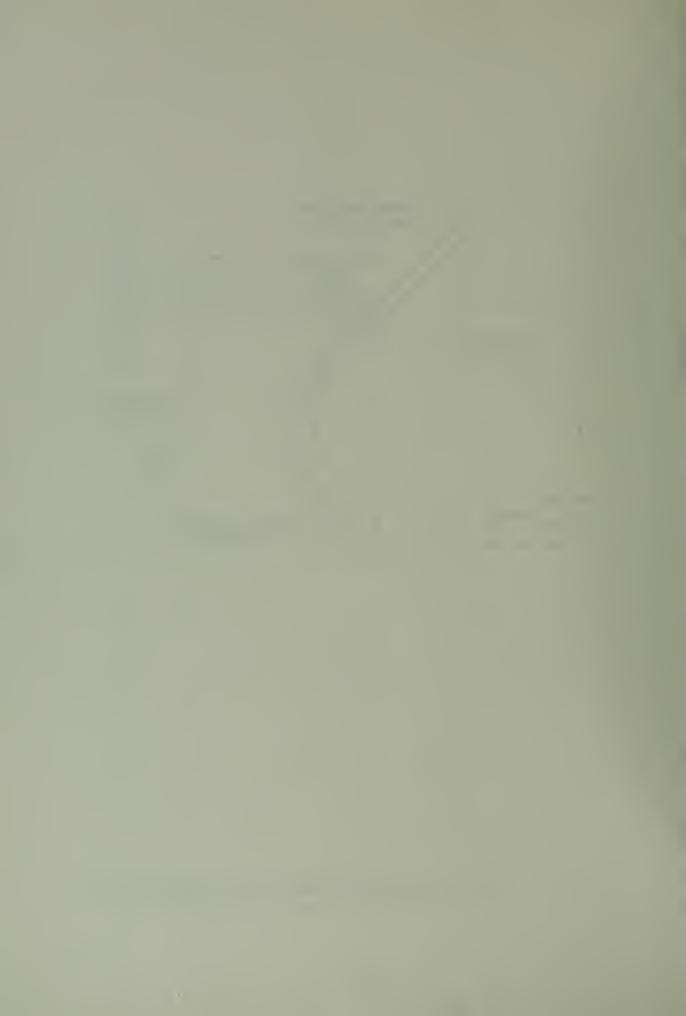


Figure 12. Schematic of the Model Used to Obtain Holographic Views Between 90° and 180° in the Direct Fin-Root Method



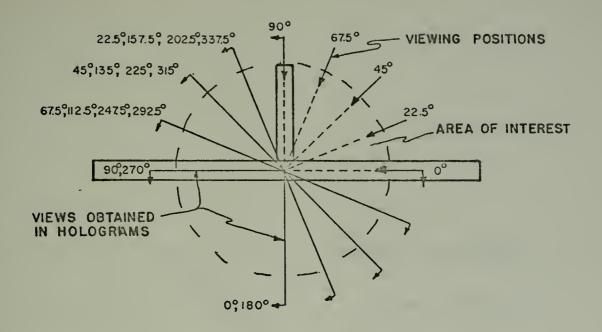


Figure 13. Holographic Viewing Angles Required in the Total Model Flow Method

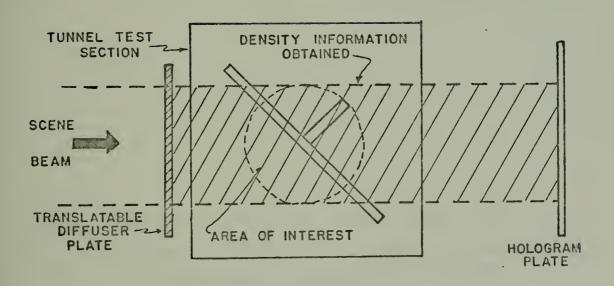
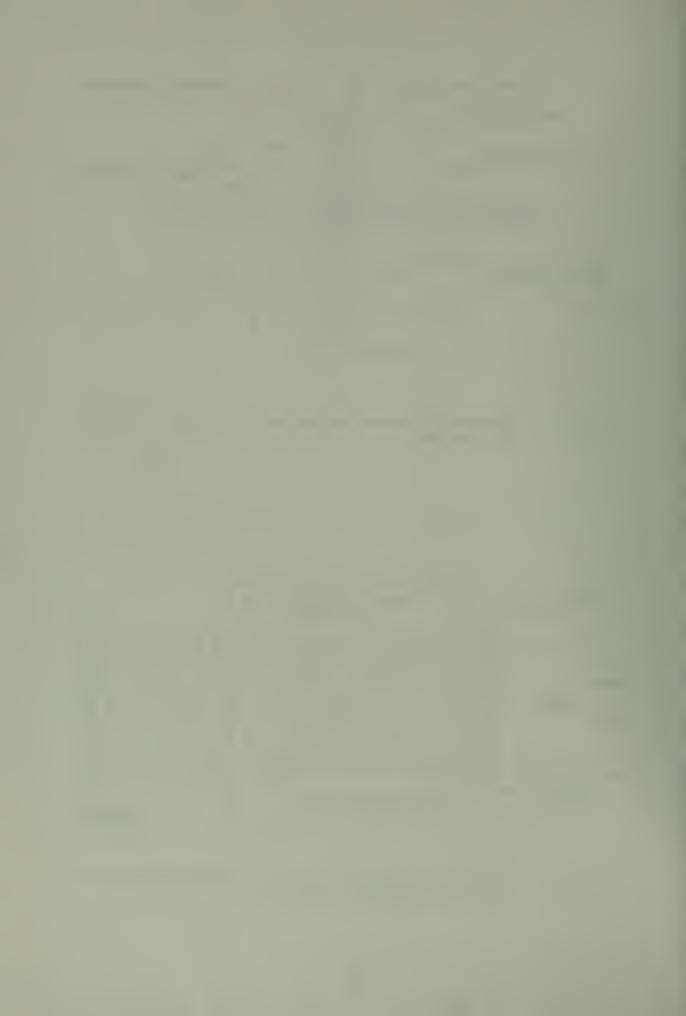


Figure 14. Schematic of the Technique Used to Obtain Holograms in the Total Model Flow Method



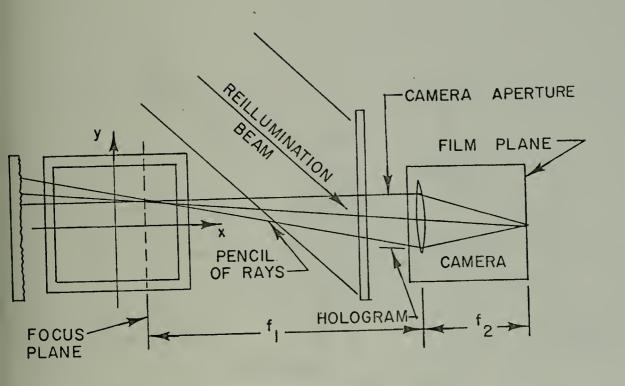
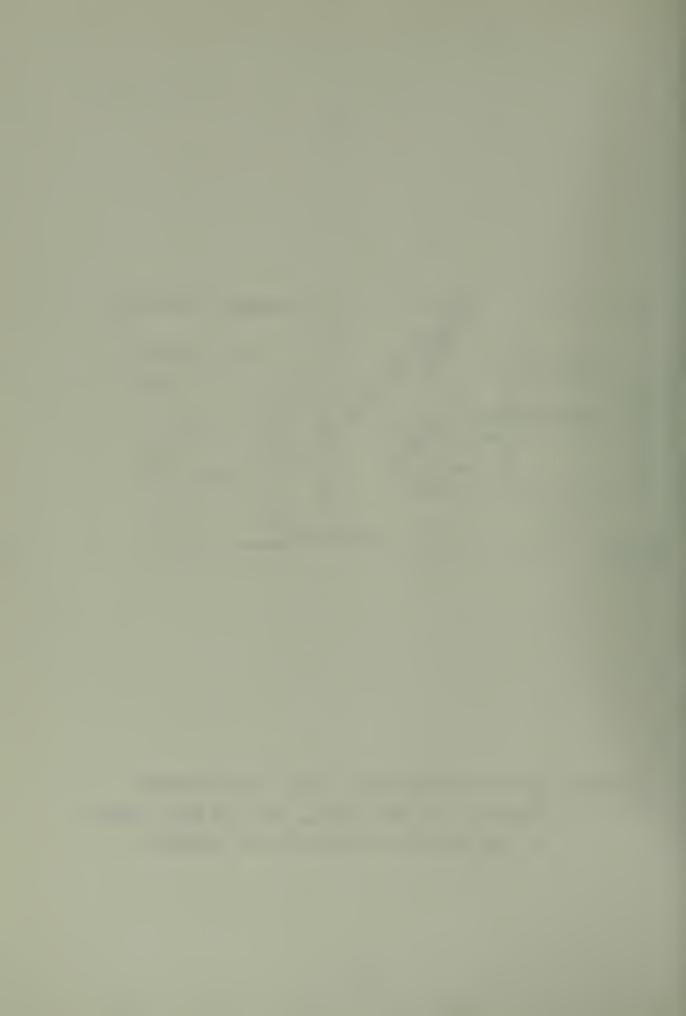


FIGURE 15. EFFECT OF APERTURE SIZE FOCUS PLANE
POSITION ON THE PENCIL SIZE OF RAYS ABOUT
A LINE OF SIGHT RECORDED BY CAMERA



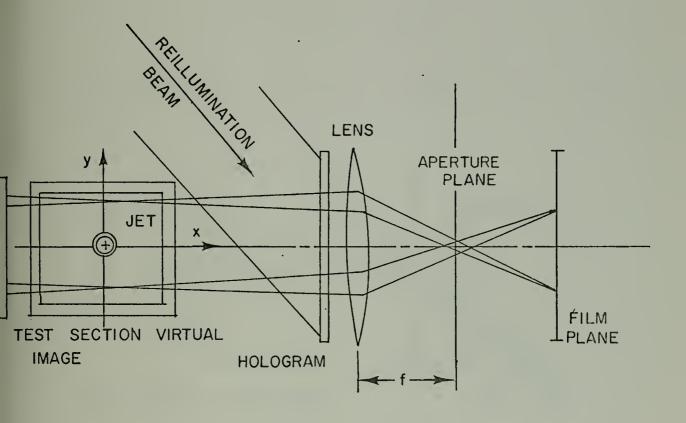
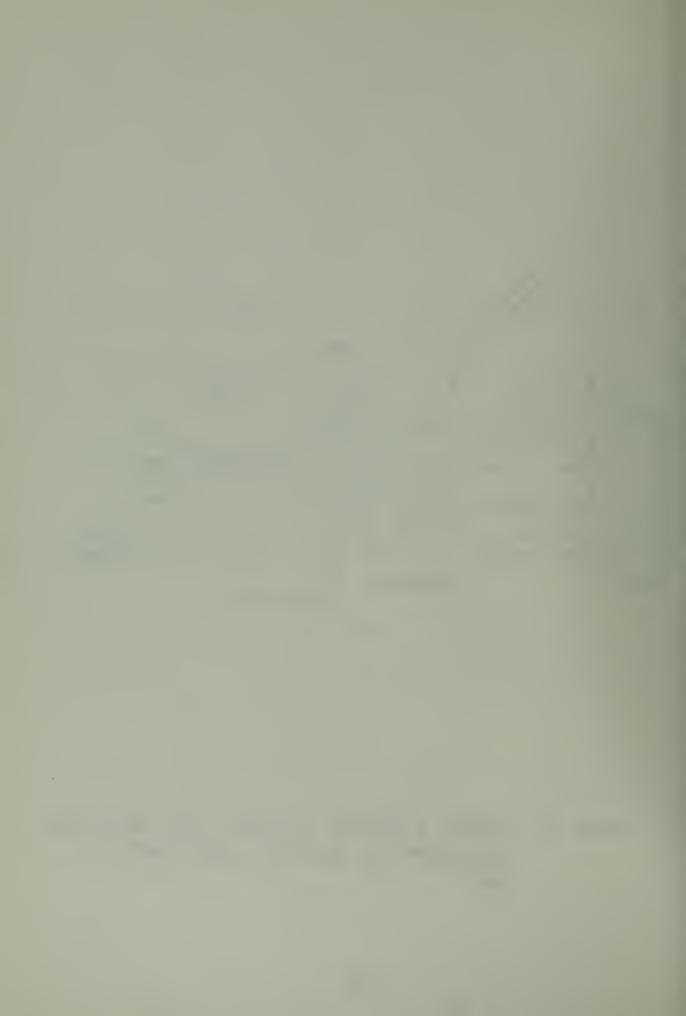


FIGURE 16. SPATIAL FILTERING TECHNIQUE FOR SELECTING PHOTOGRAPH OF CONSTANT ANGLE LINES OF LIGHT



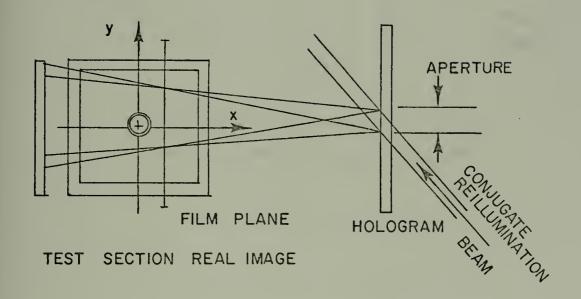


FIGURE 17. LENSLESS PHOTOGRAPHIC TECHNIQUE USING A CONJUGATE REFERENCE BEAM OF SMALL DIAMETER



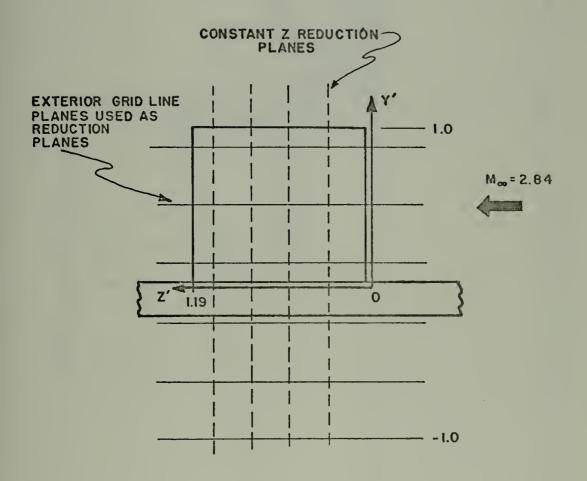
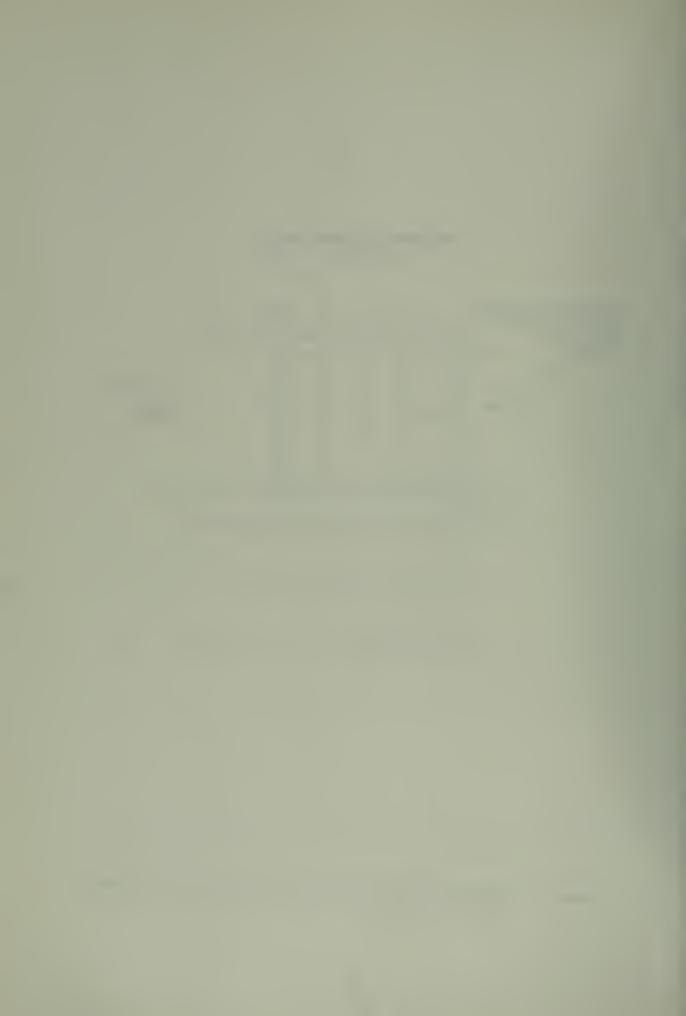
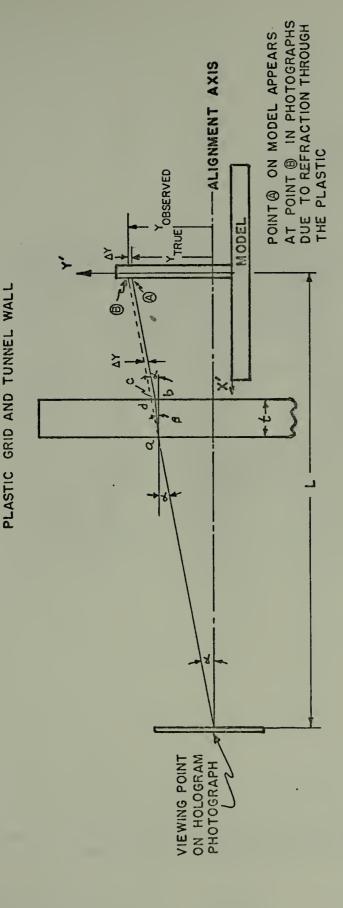
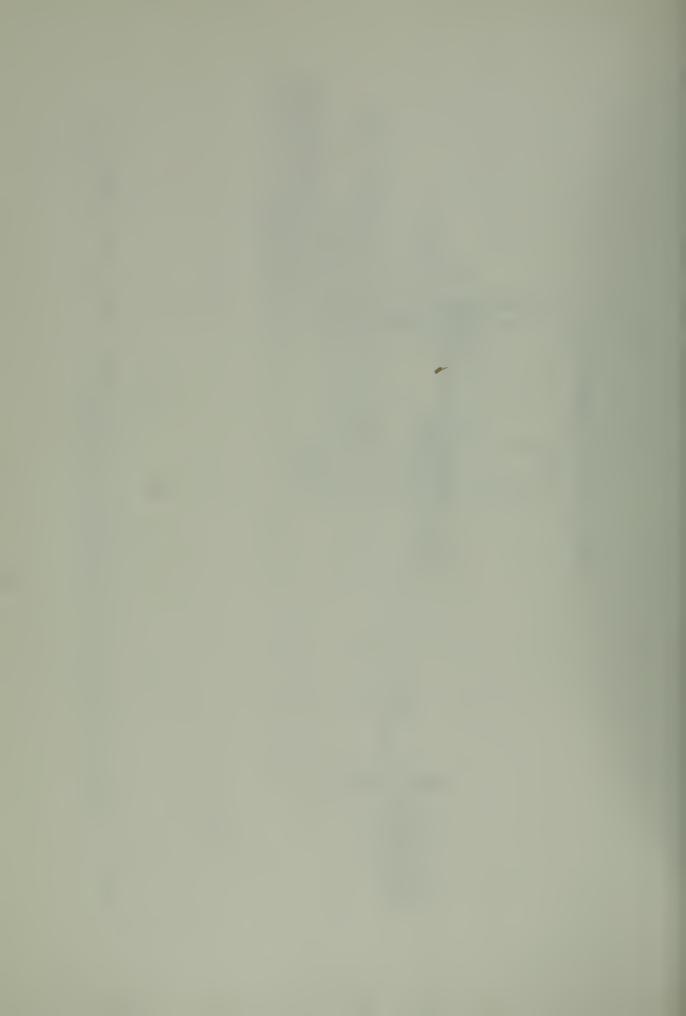


Figure 18. Data Reduction Planes Desired to Describe the Density Field Down the Fin





Schematic of the Refraction Displacement in the Photographs of Points Not Located on the Aligned Axis Caused by the Plestic Grid and Tunnel Wall Figure 19.



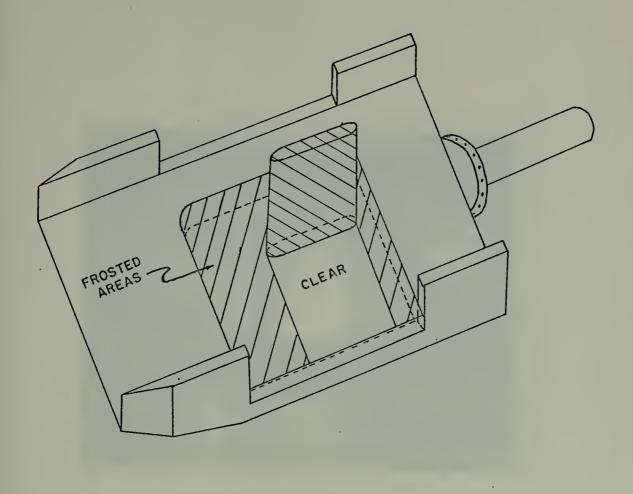
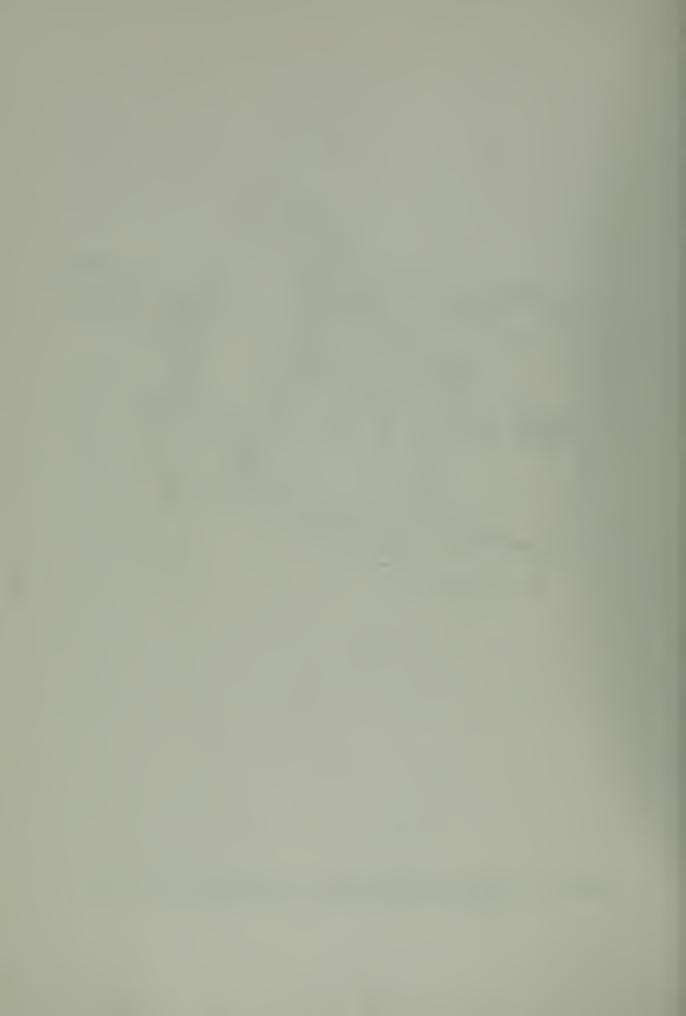


Figure 20. Schematic of the Diffused Plastic Portion of the Model Required in the Direct Fin Flow Method



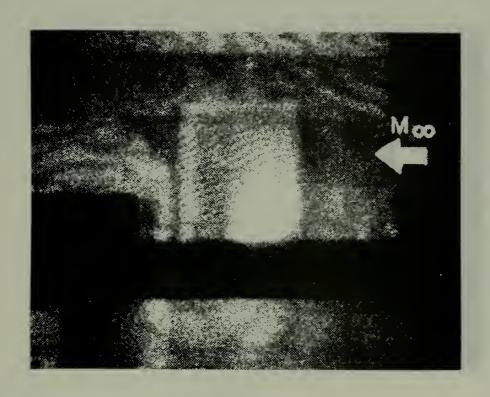


Figure 21. Interferogram Obtained in the Direct Fin Flow Method with the Model at 0° Rotation, Mach 2.84 and no Translation of Mirror,  $\rm M_{\mbox{\scriptsize 5}}$ 



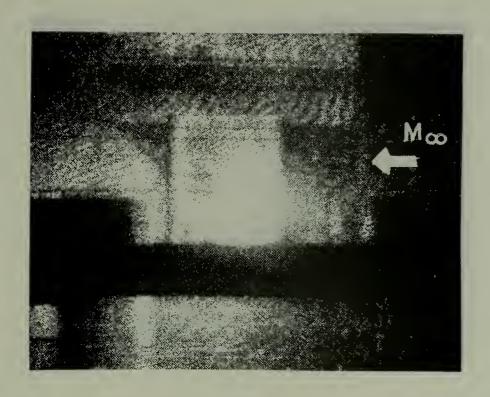
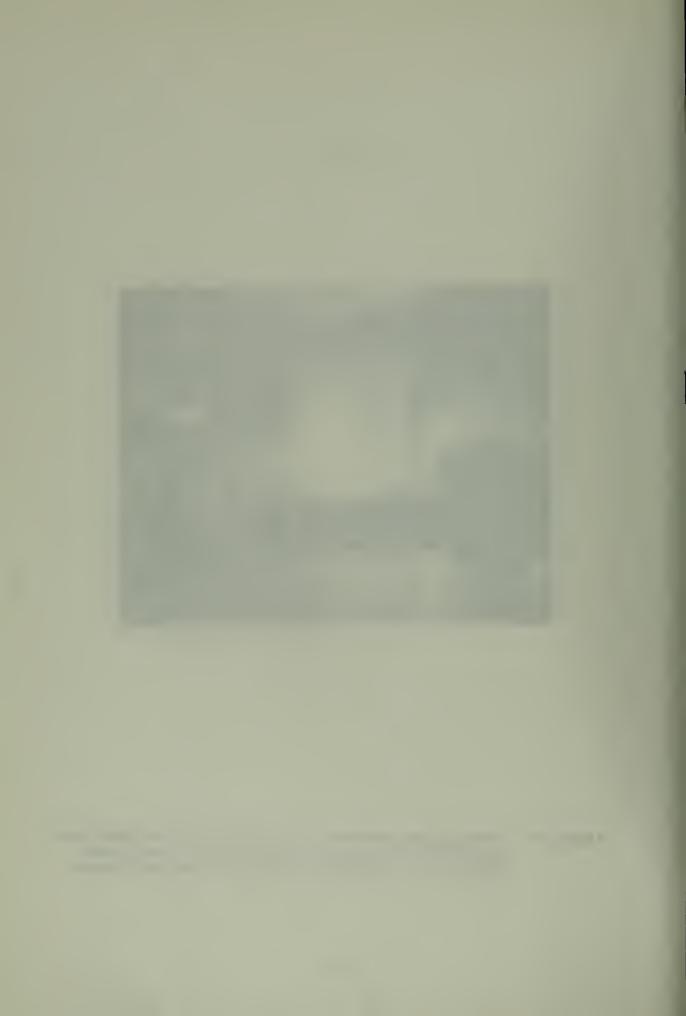


Figure 22. Interferogram Obtained in the Direct Fin Flow Method with the Model at 0° Rotation, Mach 2.84 and a .006 inches Translation of Mirror, M<sub>5</sub>, Parallel to the Test Section



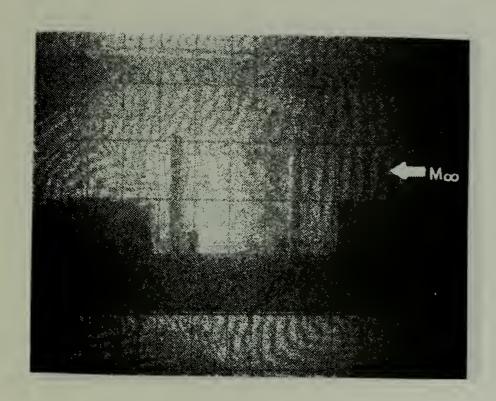
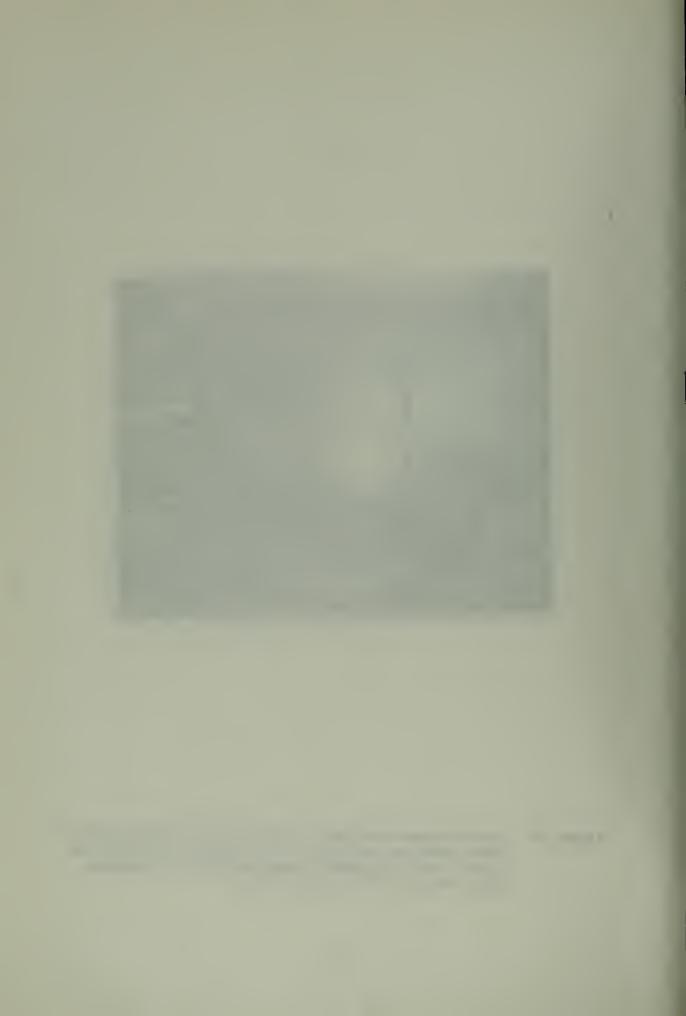


Figure 23. Interferogram Obtained in the Total Flow Method with the Clear Plexiglas Fin Model at 0° Rotation, Mach 2.84 and a .0015 inches Horizontal Translation of the Diffuser Plate Parallel to the Test Section



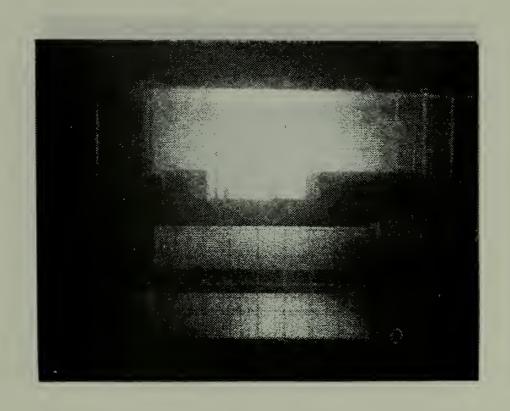


Figure 24. Interferogram Obtained in the Total Flow Method with the Clear Plexiglas Fin Model at 0° Rotation, Mach 2.84 and a .0045 Vertical Translation of the Diffuser Plate Parallel to the Test Section



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Figure 25(a). Illustration of the Shock Network Around the Fin at Mach 2.84 with a 90° Model Rotation Angle Using a Horizontal Knife Edge Schlieren System





Figure 25(b). Illustration of the Shock Network Around the Fin at Mach 2.84 with a 90° Model Rotation Angle Using a Vertical Knife Edge Schlieren System



Figure - 1997 at 1998 at 1998

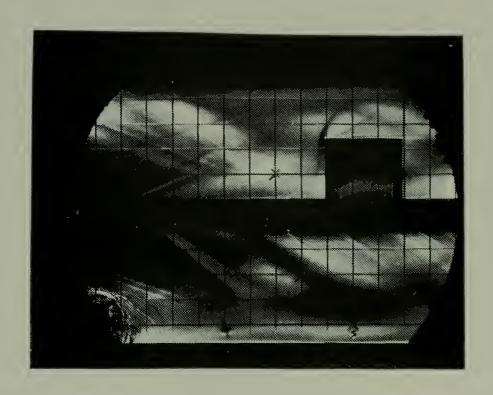
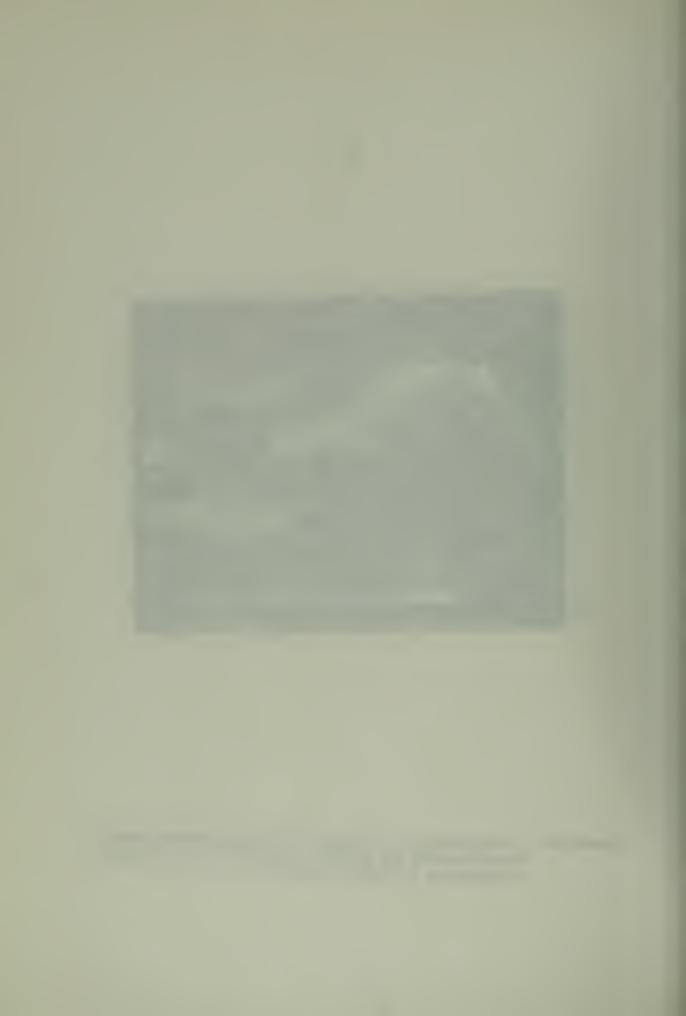


Figure 26. Illustration of the Model Flow Network at Mach 2.84, 0° Model Rotation and a Plate Vibration of ±.05 Angle of Attack Using a Schlieren System



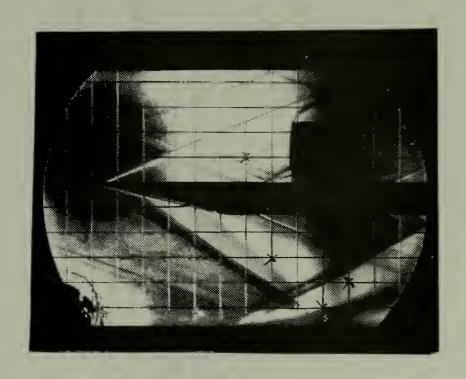
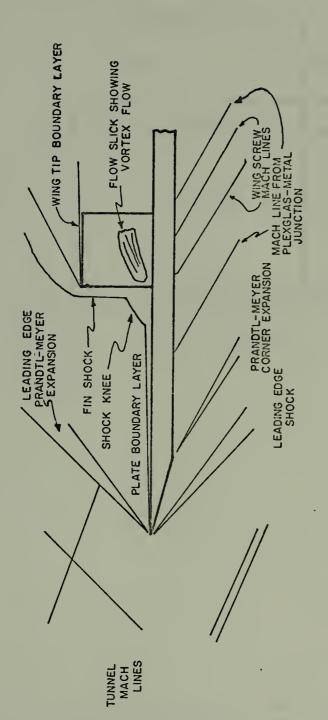
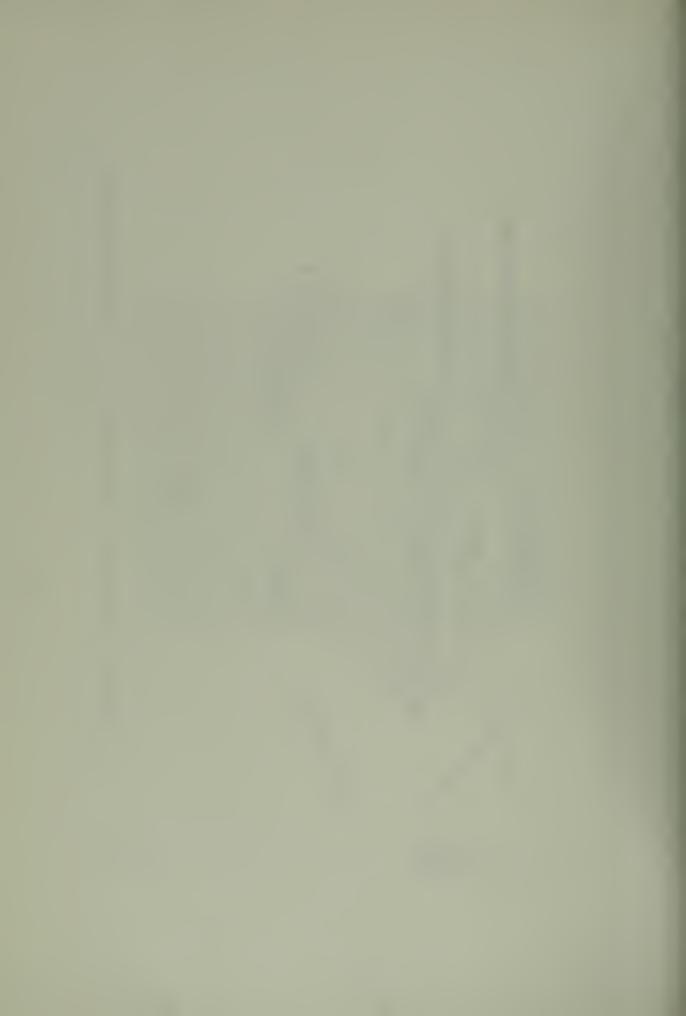


Figure 27. Illustration of the Model Flow Network at Mach 2.84, 0° Model Rotation and No Plate Vibration Using a Schlieren System





Schematic of the Flow Observed in the Schlieren Photographs Figure 28.



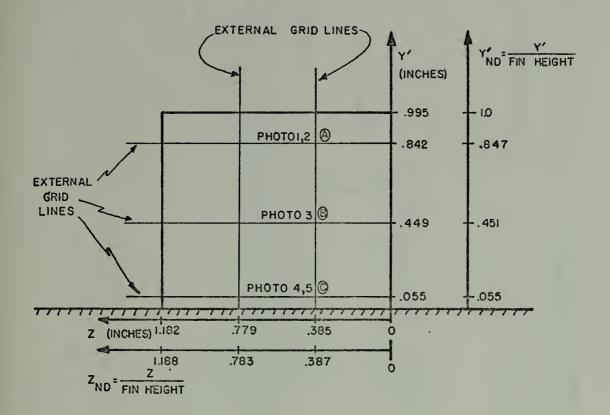
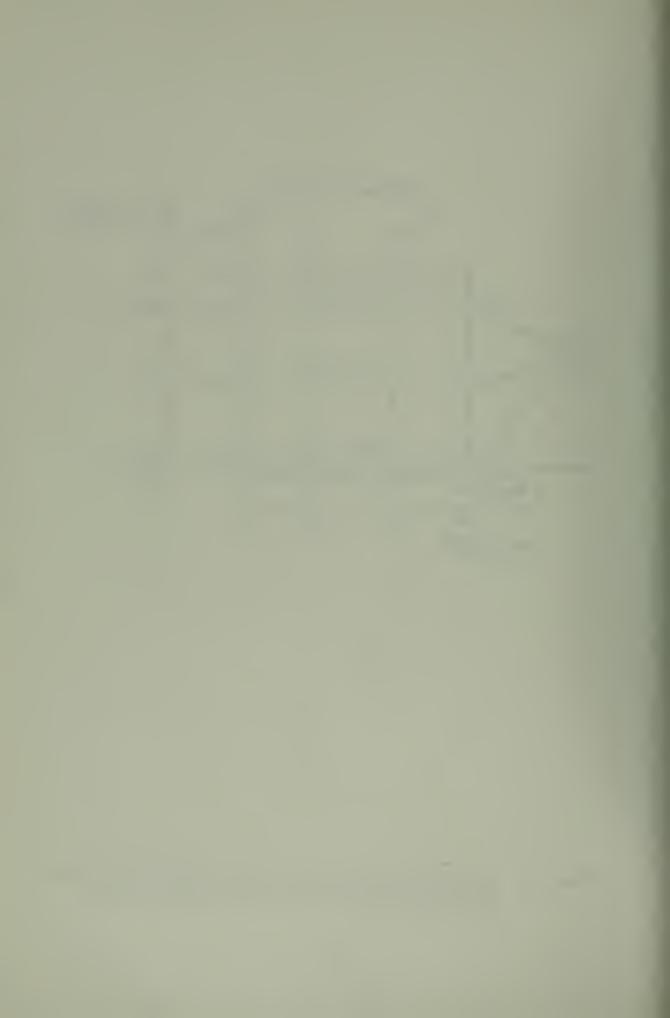
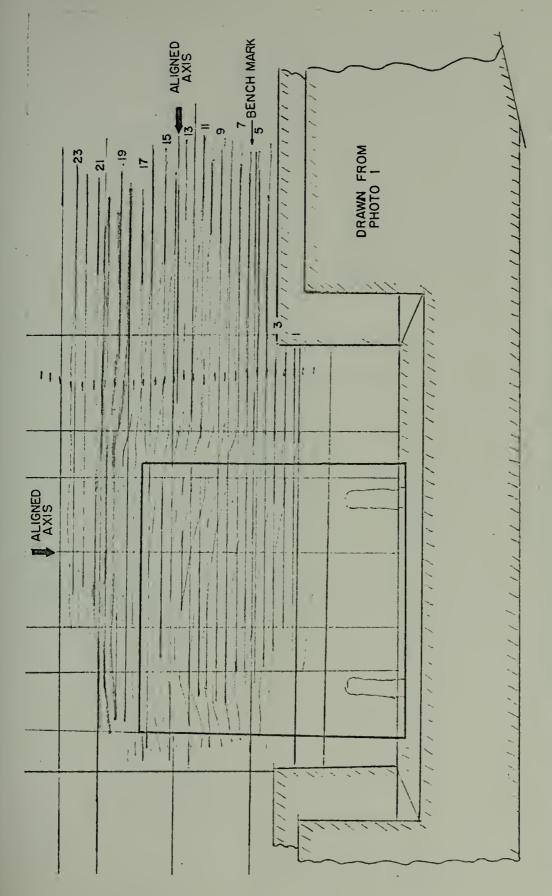
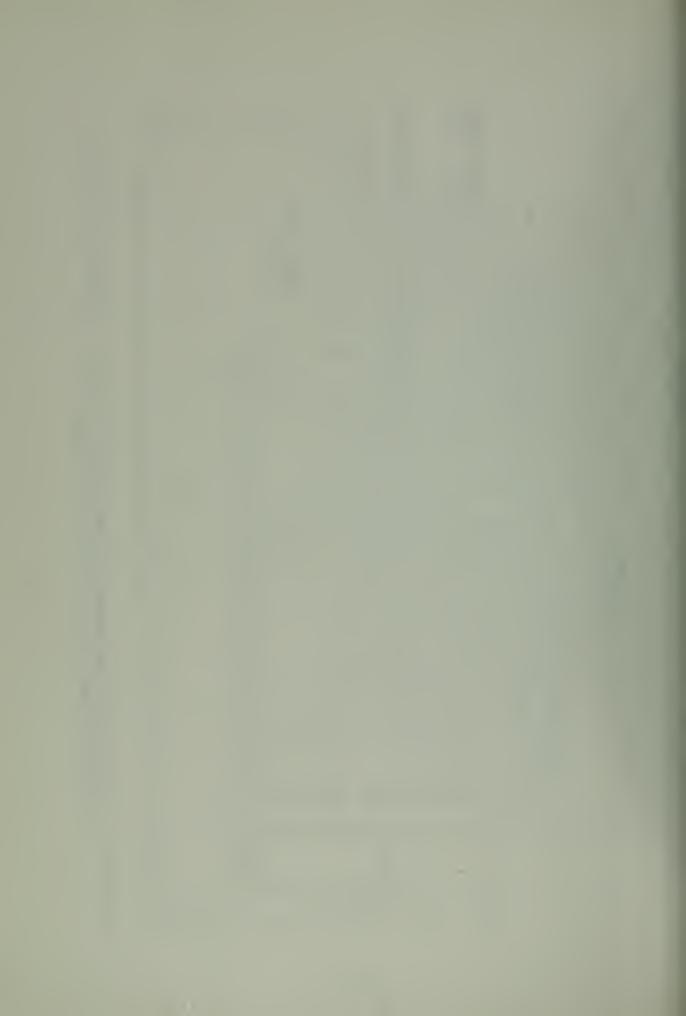


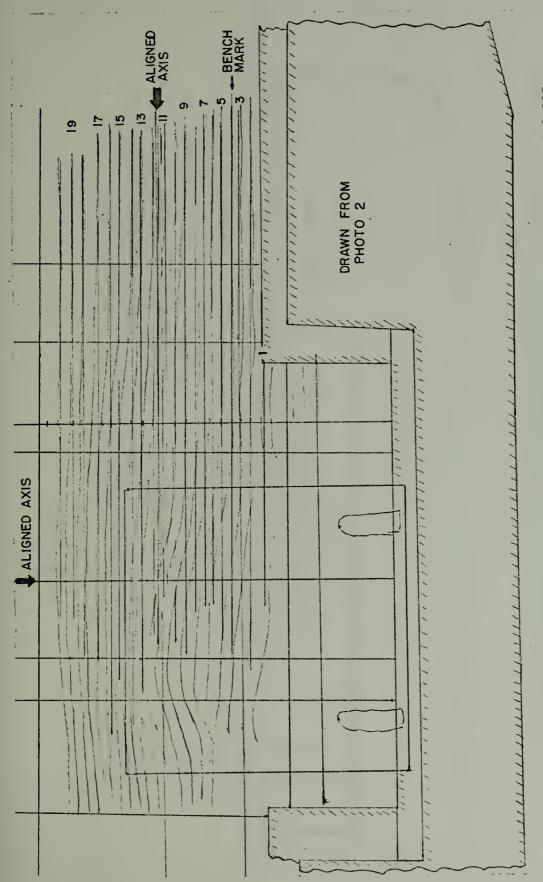
Figure 29. Schematic of the Fin at 0° Rotation Showing the Locations of the External Grid Lines and of the Interferogram Photographic Points





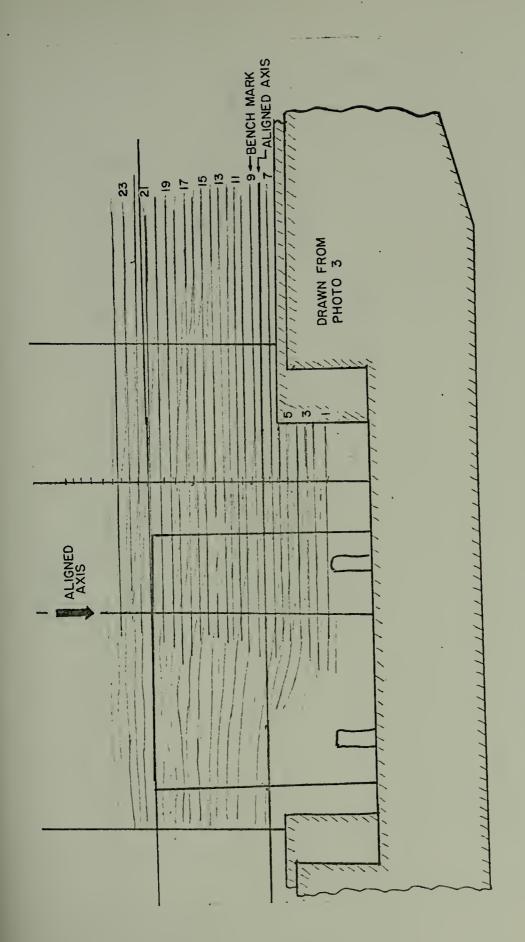
Actual Drawing and Data Reduction of Interferogram Photograph 1 Aligned at Z = 0.387, Y' = 0.847 for Mach 2.84 and 0° Model Rotation Angle Figure 30.



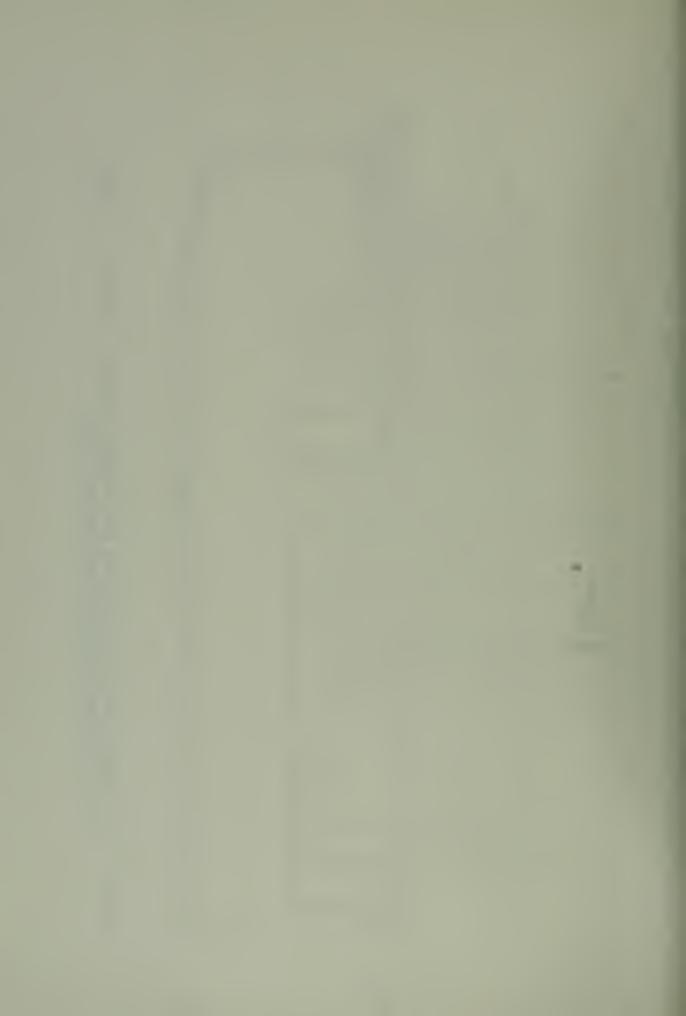


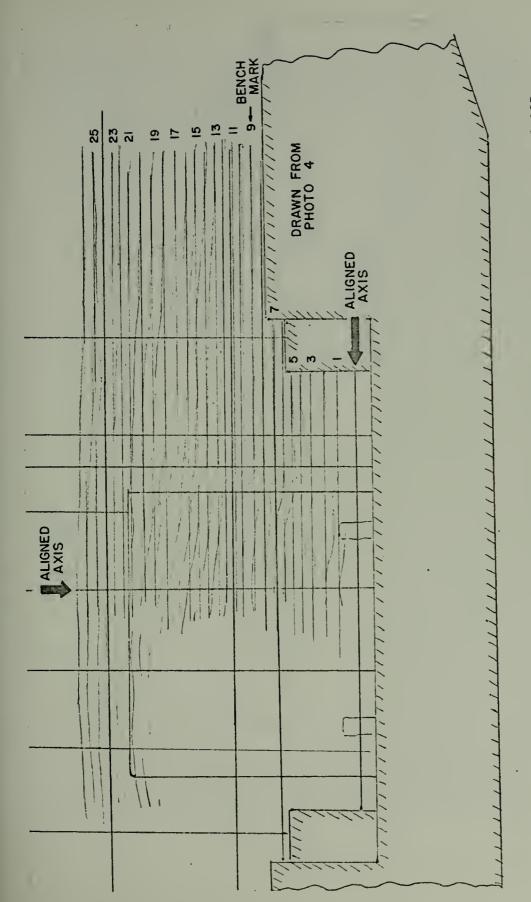
Actual Drawing and Data Reduction of Interferogram Photograph 2 Aligned at Z=0.387, Y'=0.847 for Mach 2.84 and 0° Model Rotation Angle Figure 31.



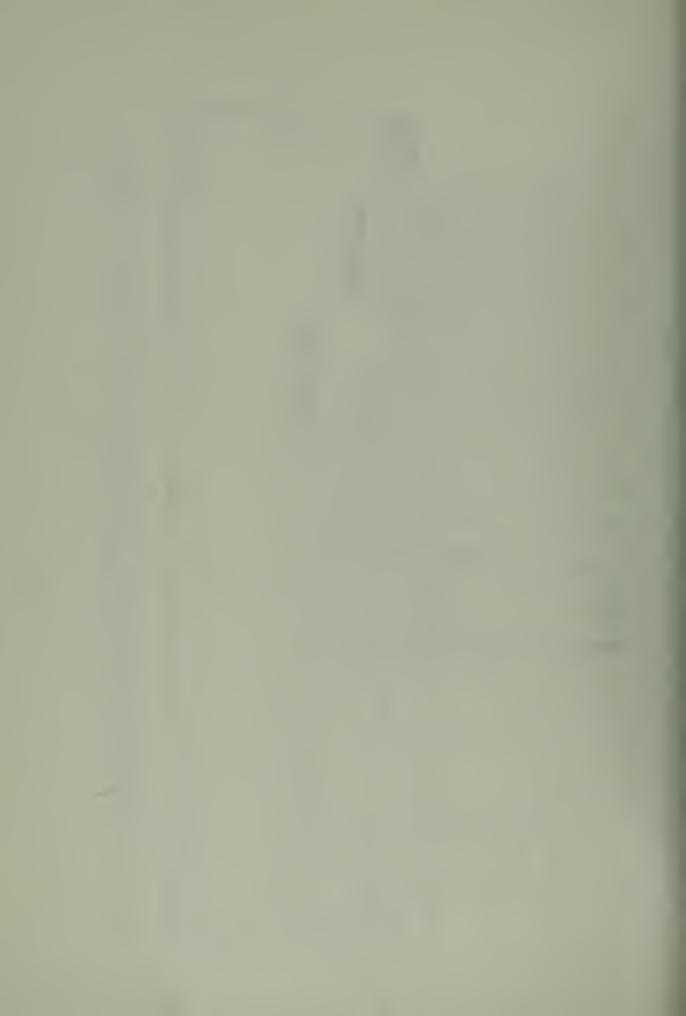


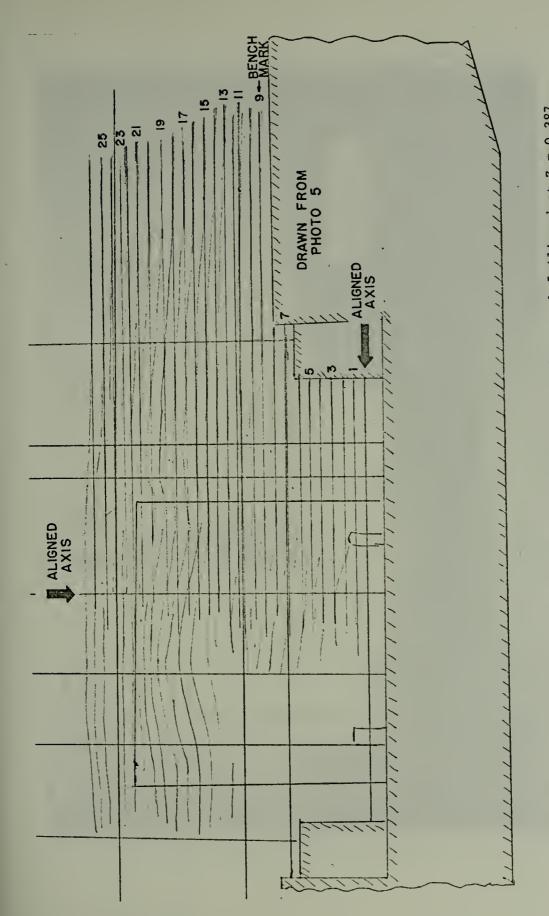
Actual Drawing and Data Reduction of Interferogram Photograph 3 Aligned at Z=0.381, Y'=0.451 for Mach 2.84 and 0° Model Rotation Angle Figure 32.



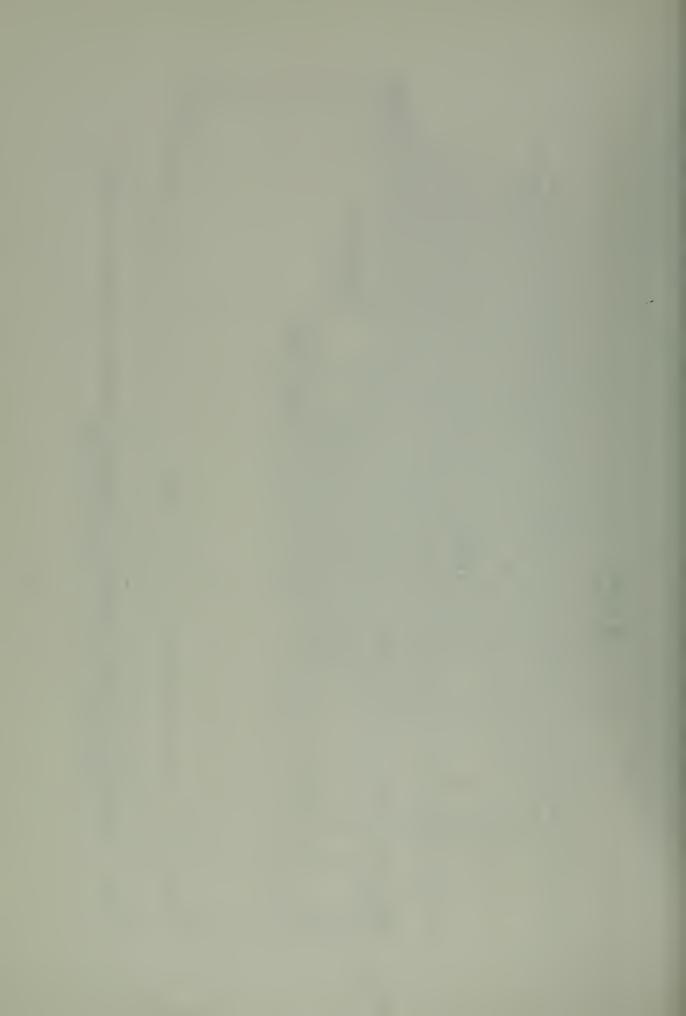


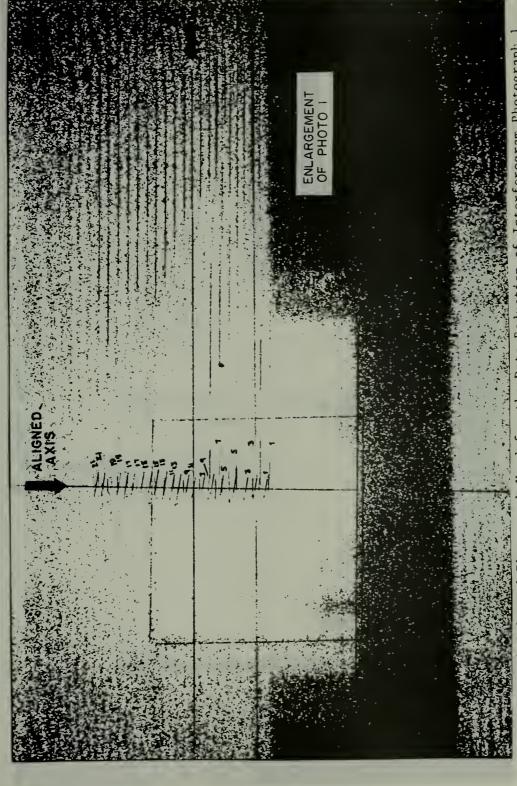
Actual Drawing and Data Reduction of Interferogram Photograph 4 Aligned at Z=0.387, Y'=0.055 for Mach 2.84 and 0° Model Rotation Angle Figure 33.



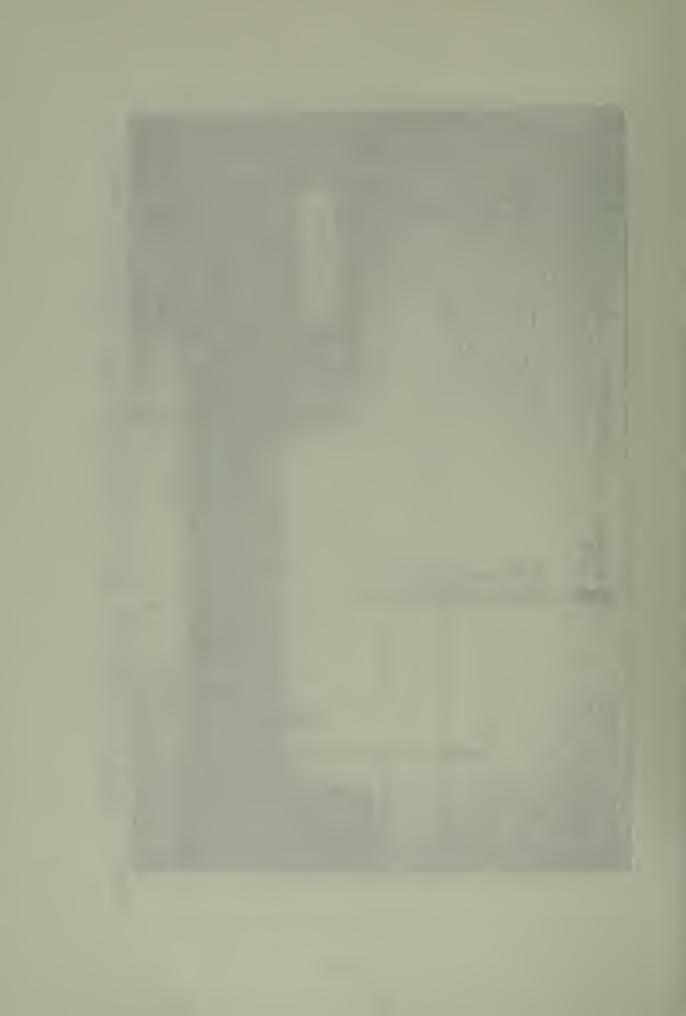


Actual Drawing and Data Reduction of Interferogram Photograph 5 Aligned at Z = 0.387, Y' = 0.055 for Mach 2.84 and 0° Model Rotation Angle Figure 34.





Actual Enlarged Photograph Used for the Data Reduction of Interferogram Photograph 1 Aligned at Z = 0.387, Y' = 0.847 for Mach 2.84 and 0° Model Rotation Angle Figure 35.



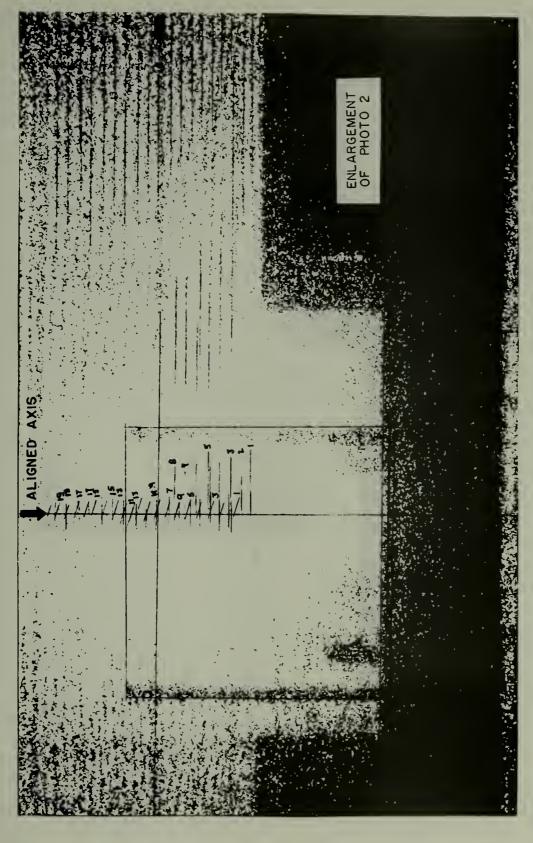
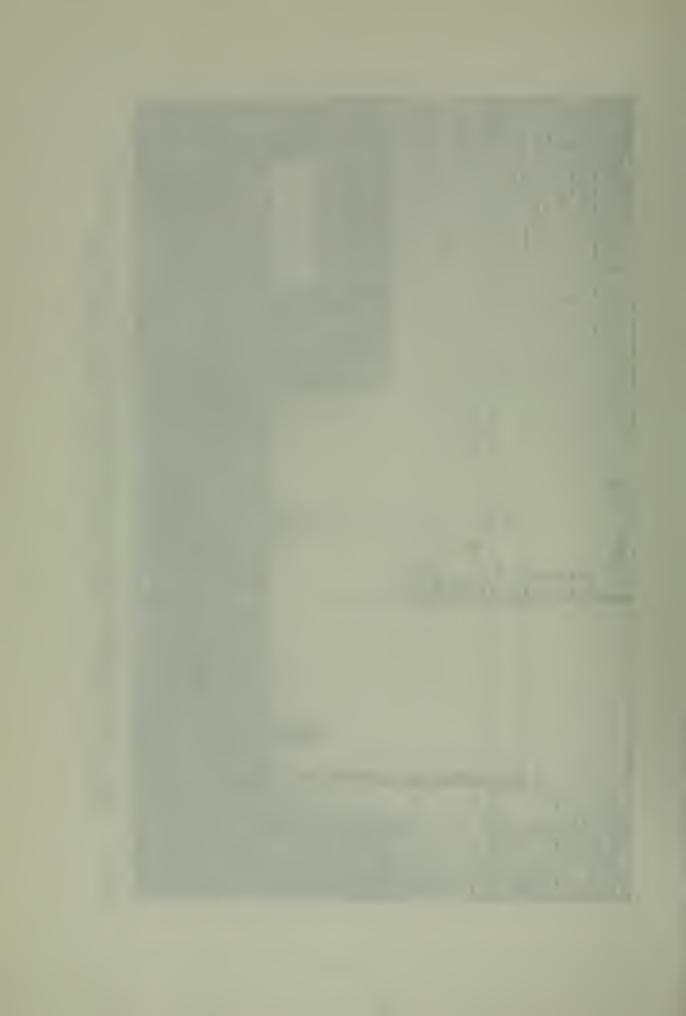
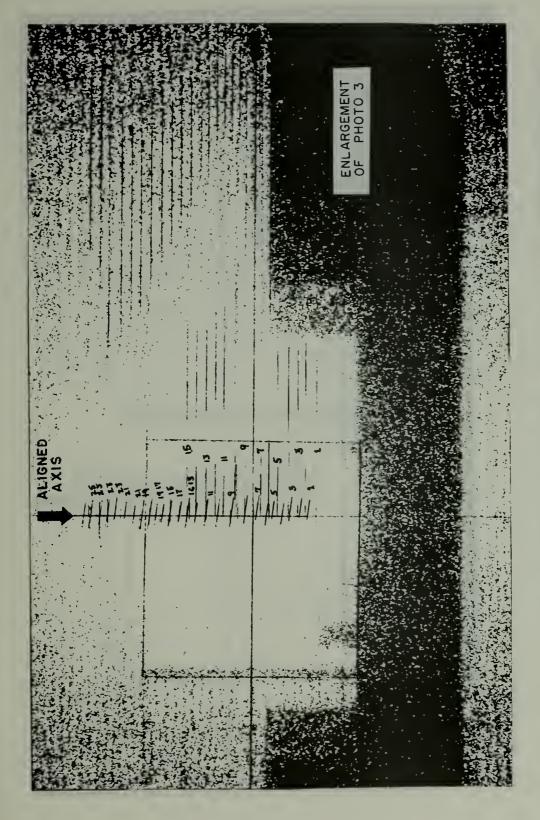
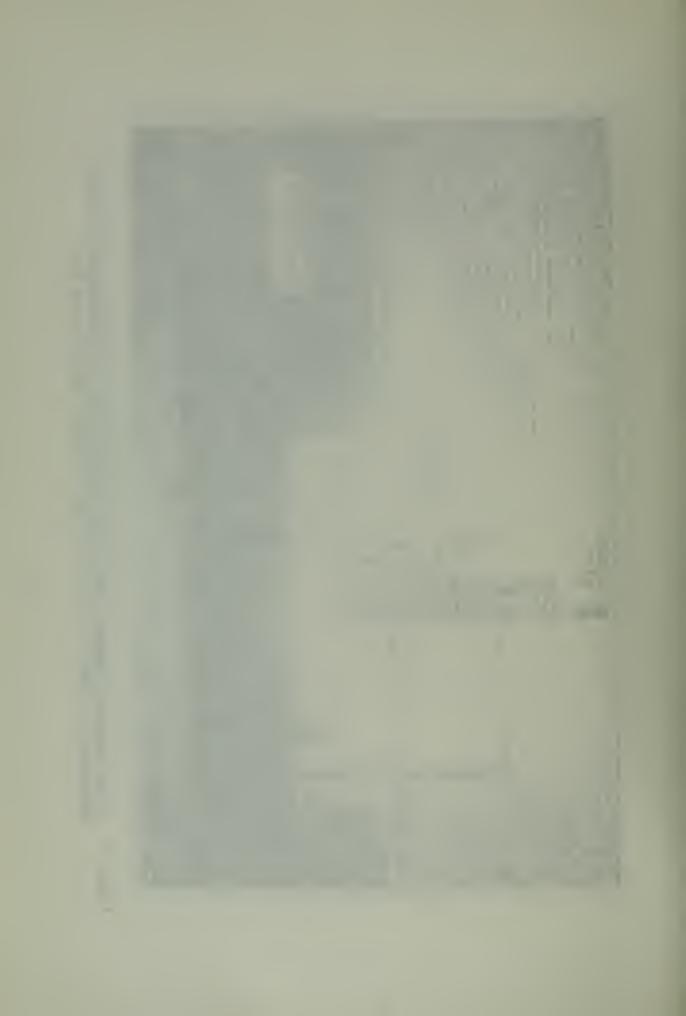


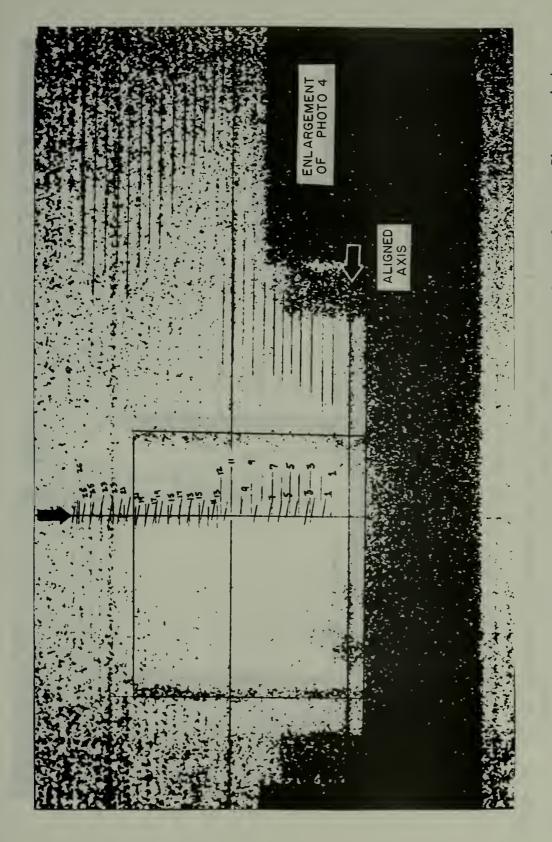
Figure 36. Actual Enlarged Photograph Used for the Data Reduction of Interferogram Photograph 2 Aligned at Z = 0.387, Y' = 0.847 for Mach 2.84 and 0° Model Rotation Angle





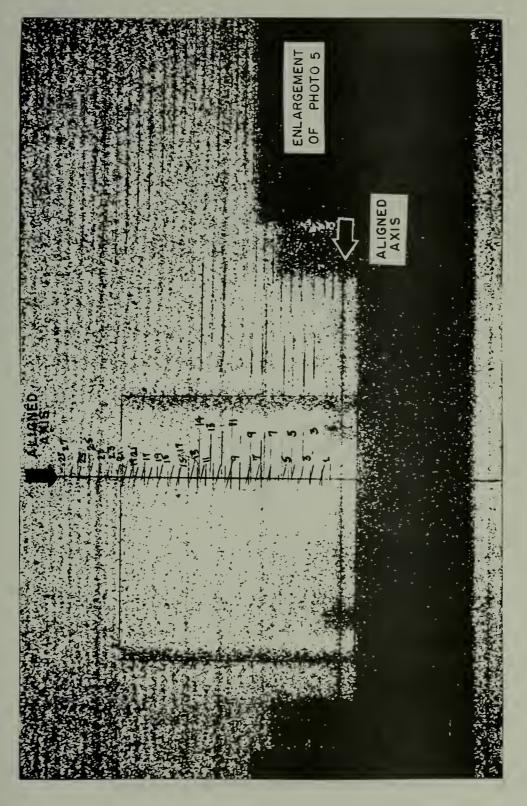
Actual Enlarged Photograph Used for the Data Reduction of Interferogram Photograph 3 Aligned at Z = 0.387, Y' = 0.451 for Mach 2.84 and 0 Model Potation Argle Figure 37.



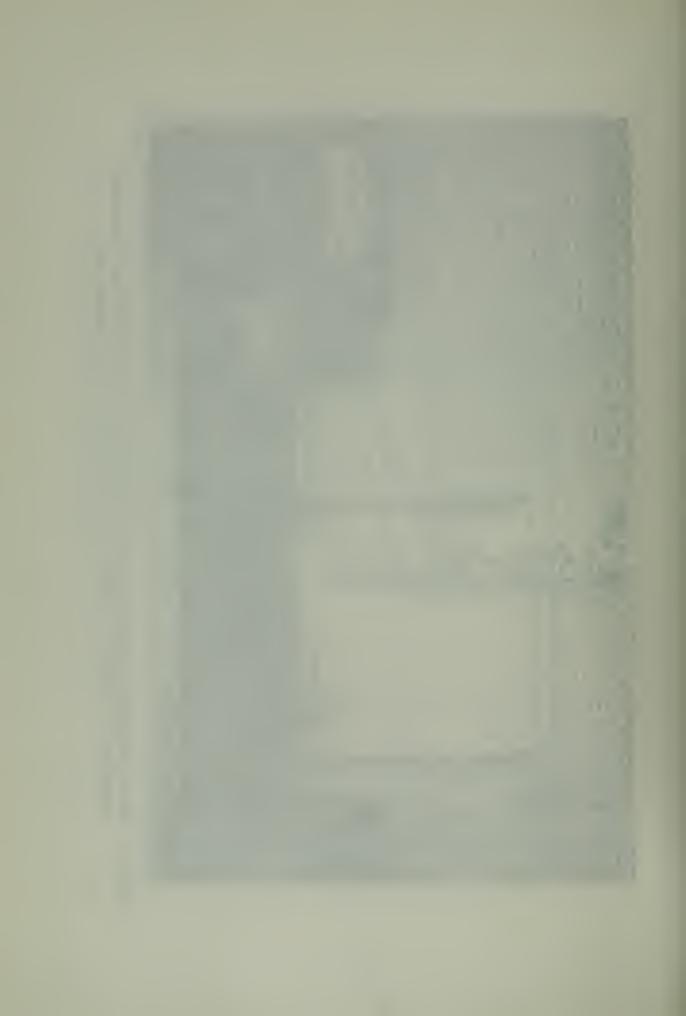


Actual Enlarged Photograph Used for the Data Reduction of Interferogram Photograph Aligned at Z = 0.387, Y' = 0.055 for Mach 2.84 and 0° Model Rotation Angle Figure 38.





Actual Enlarged Photograph Used for the Data Reduction of Interferopram Photograph Aligned at Z = 0.387, Y' = 0.055 for Mach 2.84 and 0° Model Rotation Apple Figure 39.



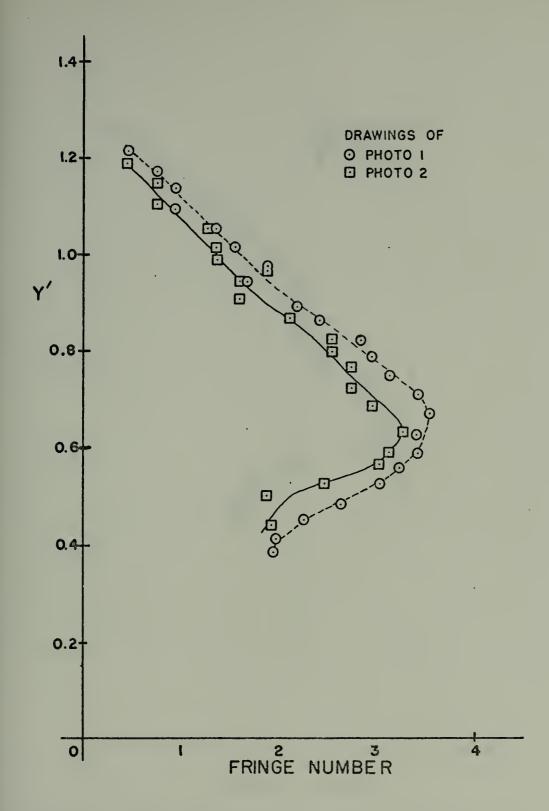
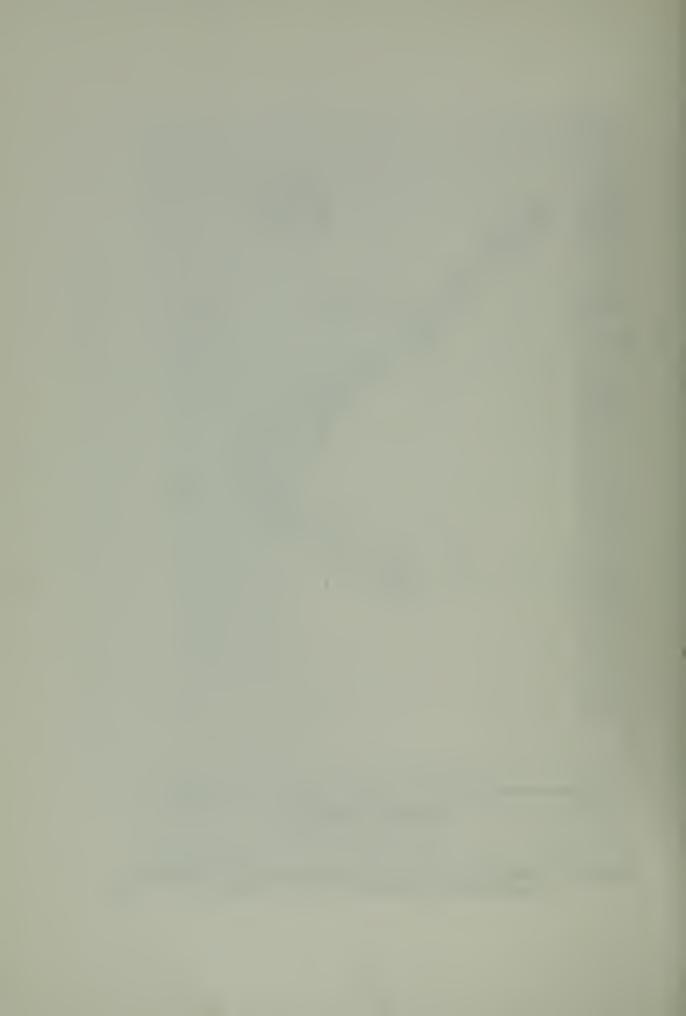


Figure 40. Comparison of Fringe Data Obtained from Drawings of Interferogram Photographs 1 and 2 Aligned at Y' = 0.847



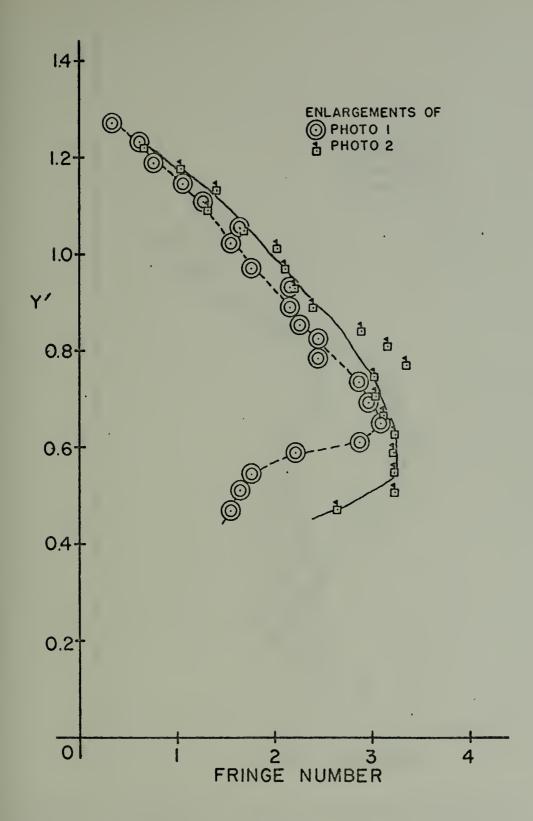


Figure 41. Comparison of Fringe Data Obtained from Photographic Enlargements of Interferogram Photographs 1 and 2 Aligned at Y' = 0.847



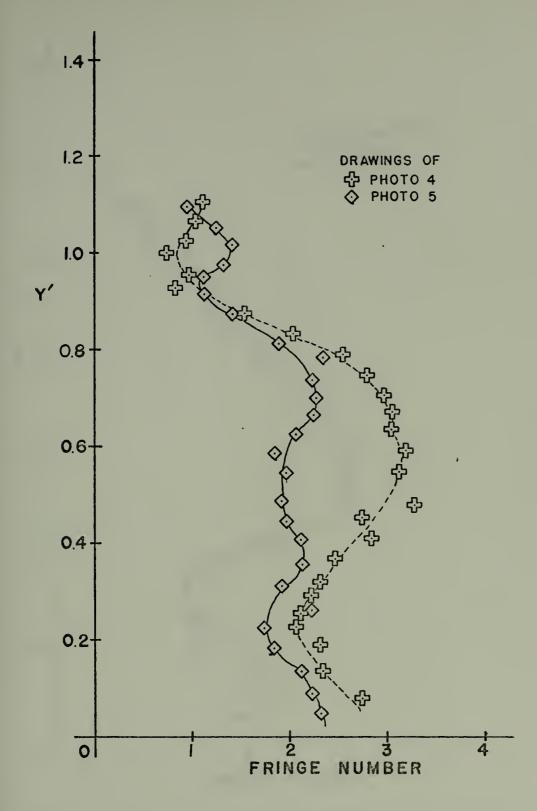


Figure 42. Comparison of Fringe Data Obtained from Drawings of Interferogram Photographs 4 and 5 Aligned at Y' = 0.055



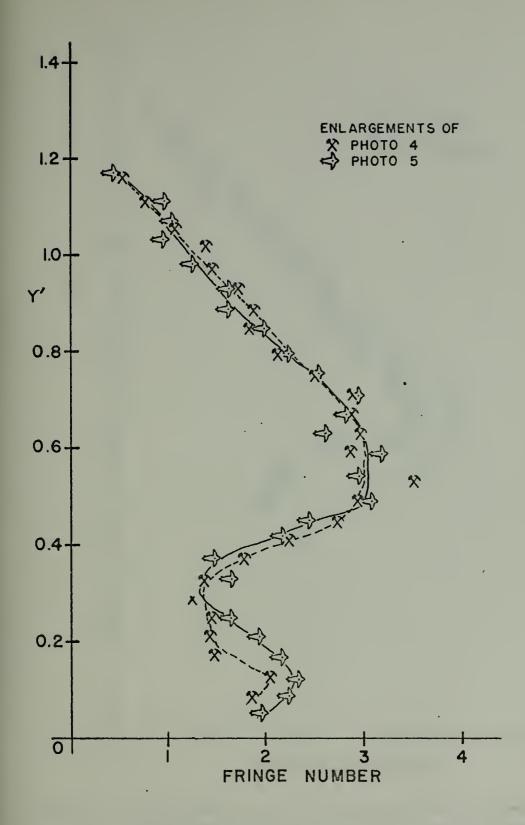
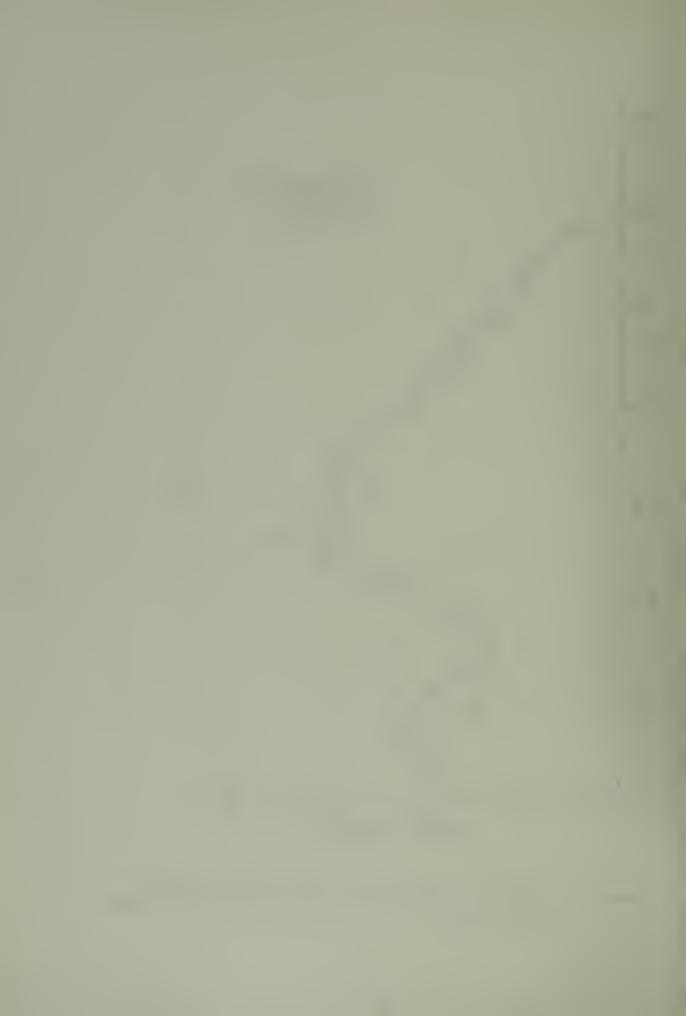


Figure 43. Comparison of Fringe Data Obtained from Photographic Enlargements of Interferogram Photographs 4 and 5 Aligned at Y' = 0.055



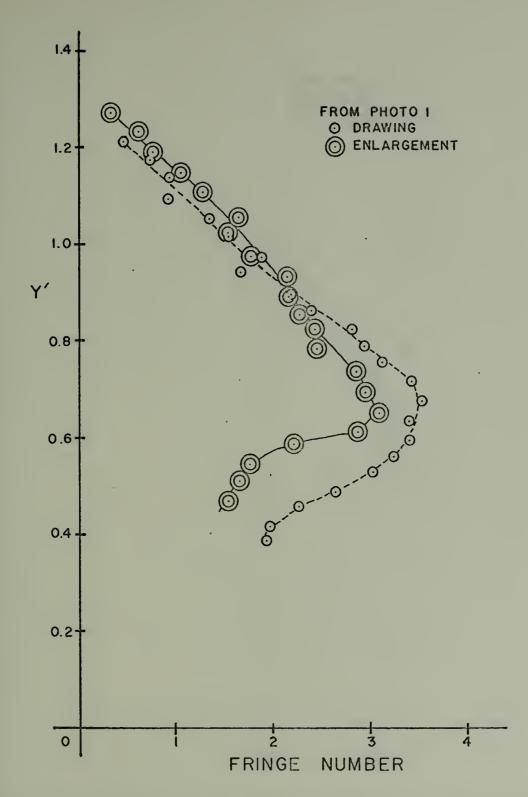
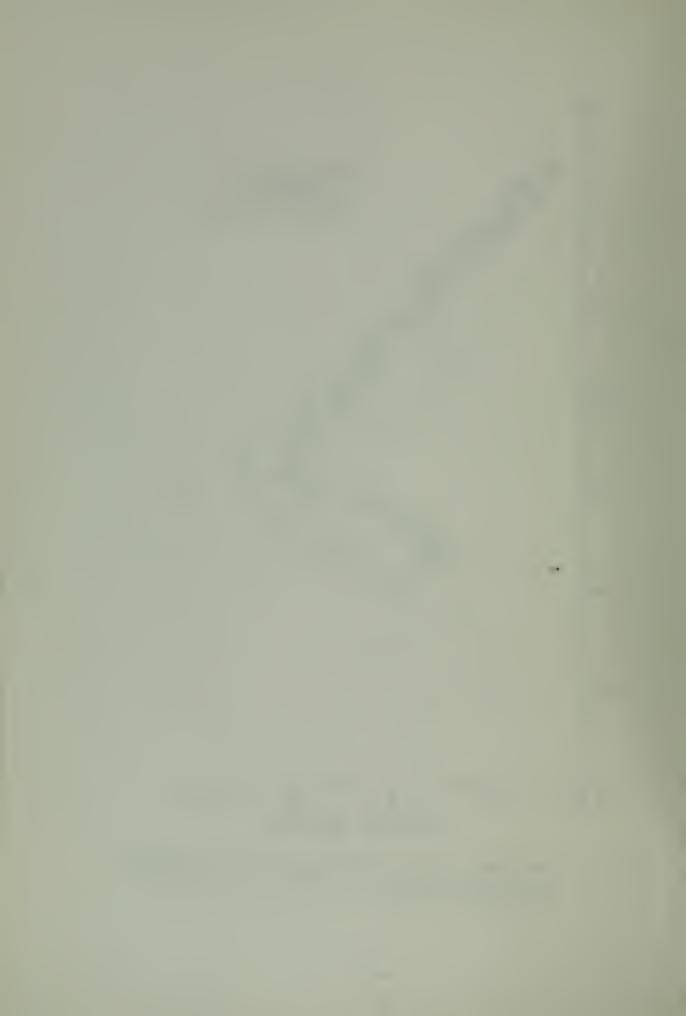


Figure 44. Comparison of Fringe Data Obtained from the Drawing and Photographic Enlargement of Interferogram Photograph 1 Aligned at Y \*\* 0.847



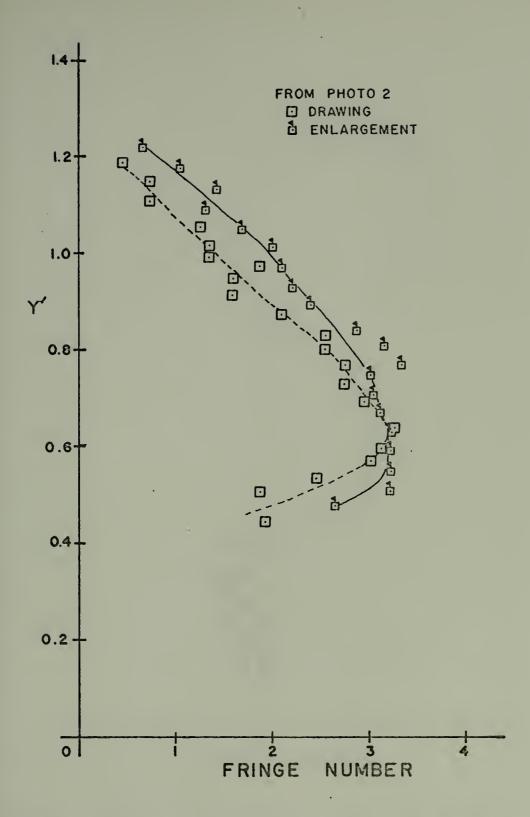


Figure 45. Comparison of Fringe Data Obtained from the Drawing and Photographic Enlargement of Interferogram Photograph 2

Aligned at Y' = 0.847



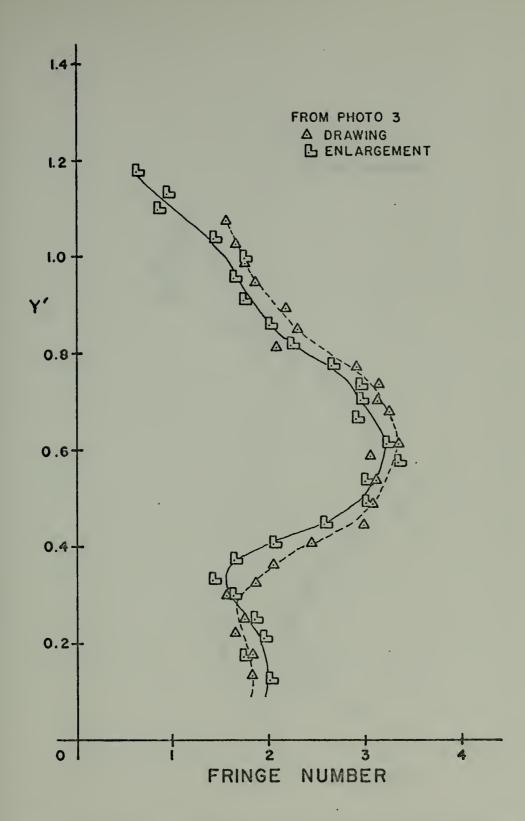
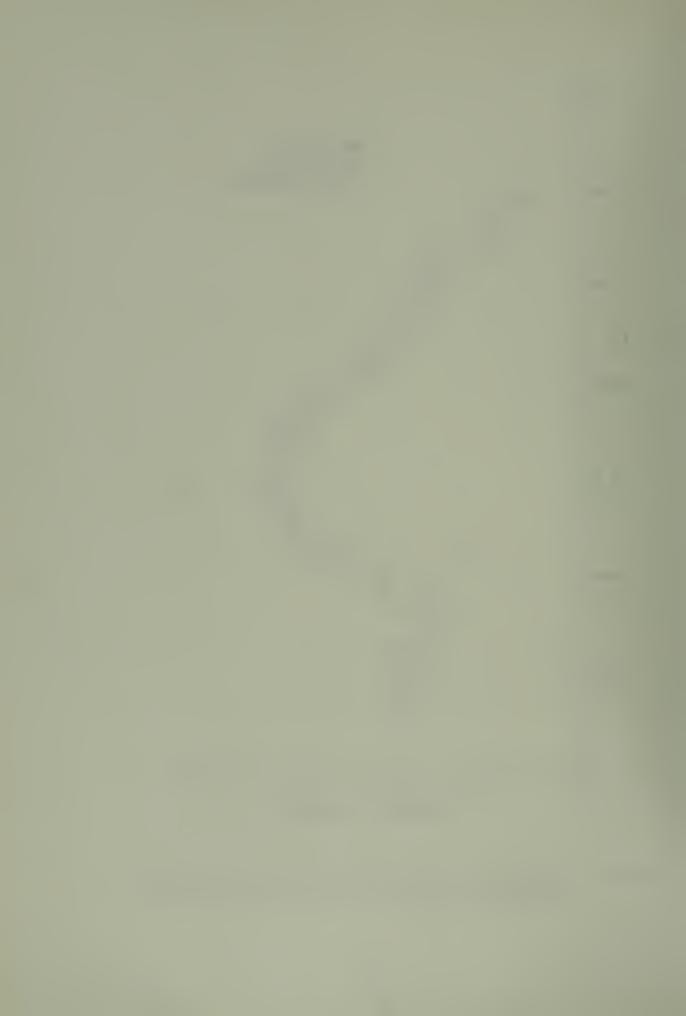


Figure 46. Comparison of Fringe Data Obtained from the Drawing and Photographic Enlargement of Interferogram Photograph 3 Aligned at Y' = 0.451



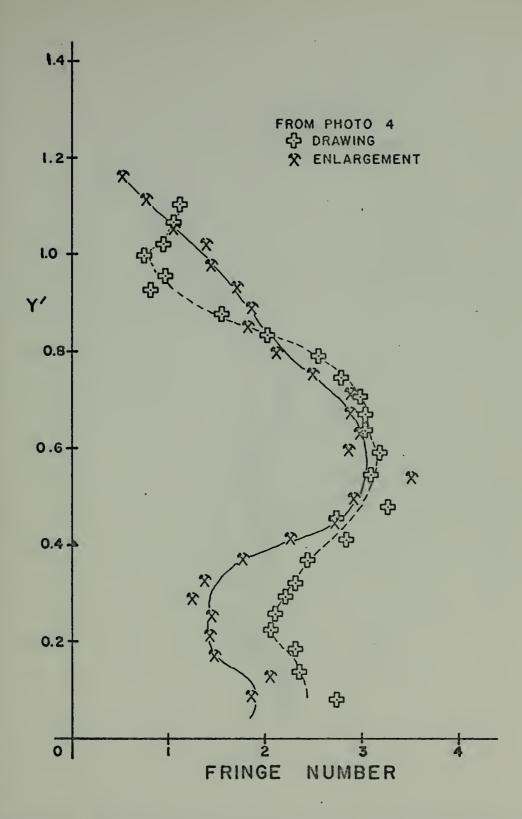
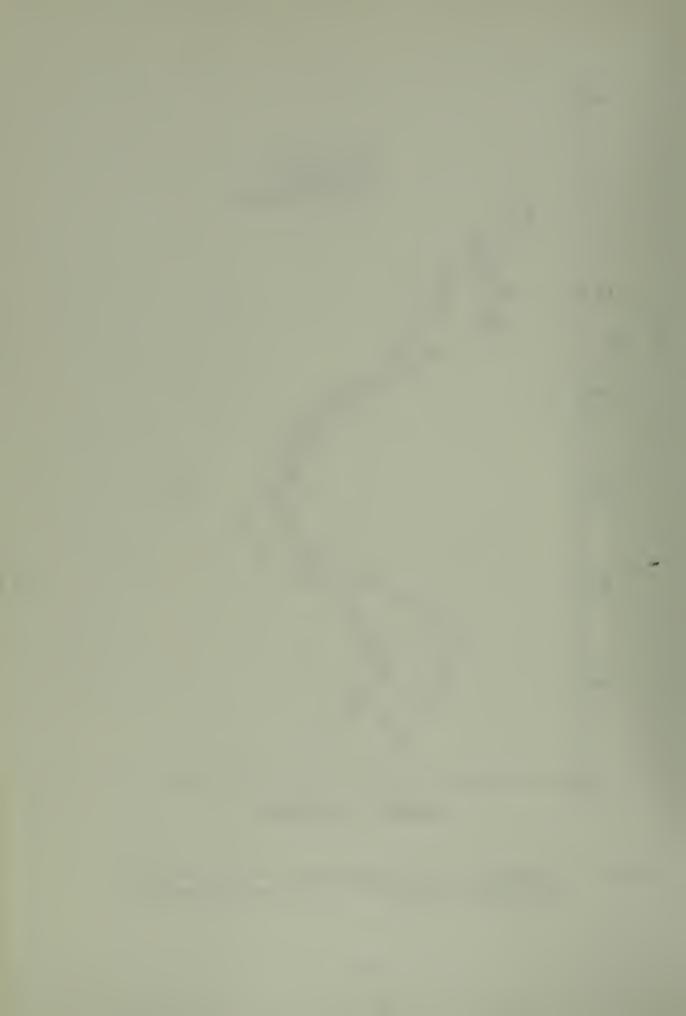


Figure 47. Comparison of Fringe Data Obtained from the Drawing and Photographic Enlargement of Interferogram Photograph 4
Aligned at Y' = 0.055



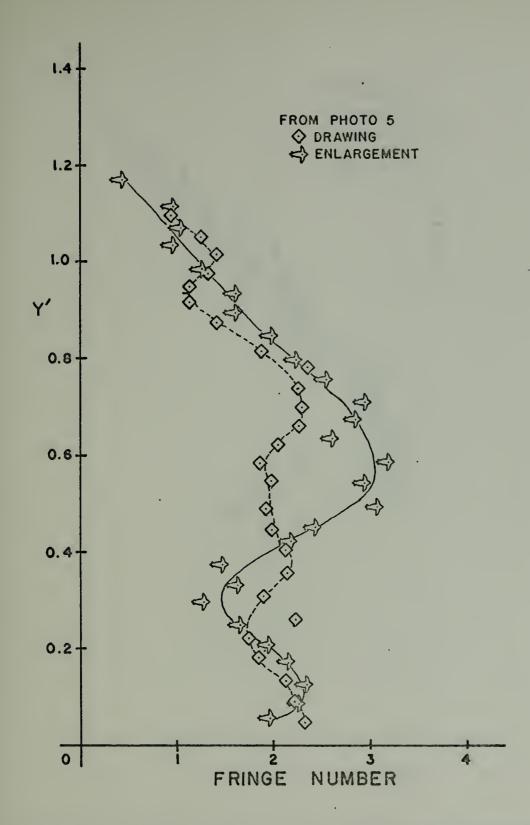


Figure 48. Comparison of Fringe Data Obtained from the Drawing and Photographic Enlargement of Interferogram Photograph 5
Aligned at Y' = 0.055



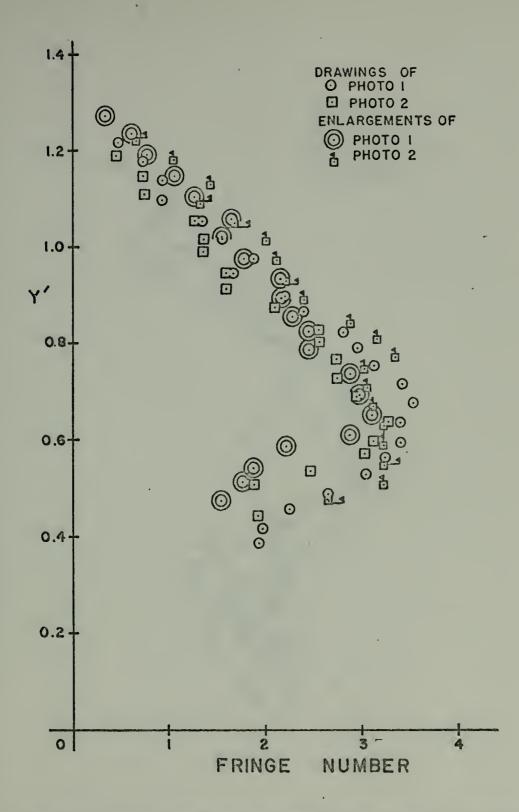
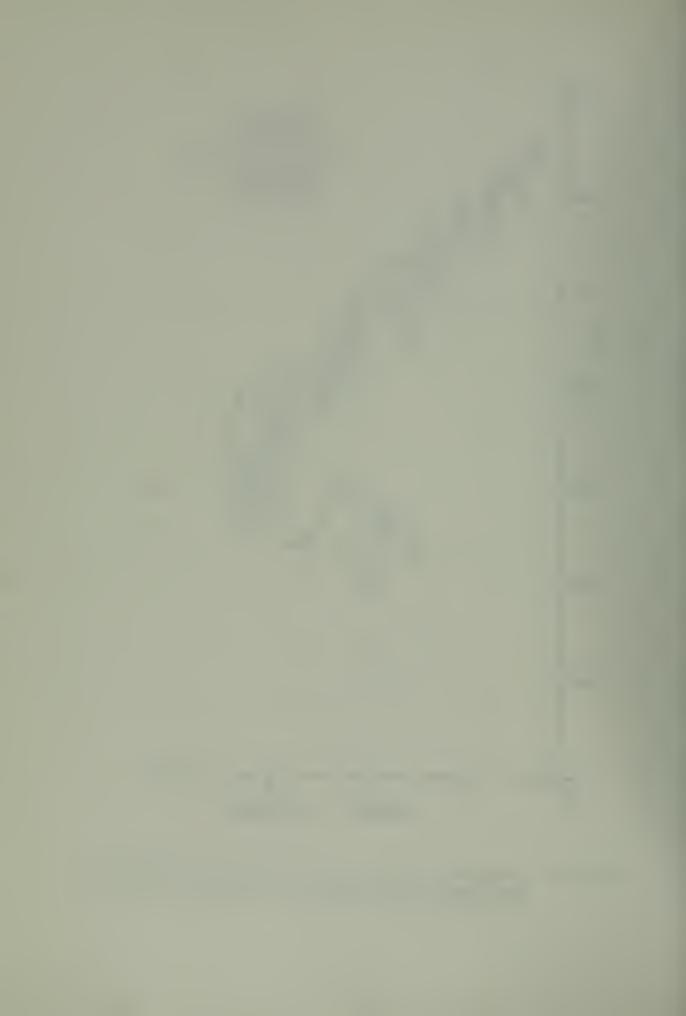


Figure 49. Comparison of Fringe Data Obtained from the Drawings and Photographic Enlargements of Interferogram Photographs 1 and 2 Aligned at Y' = 0.847



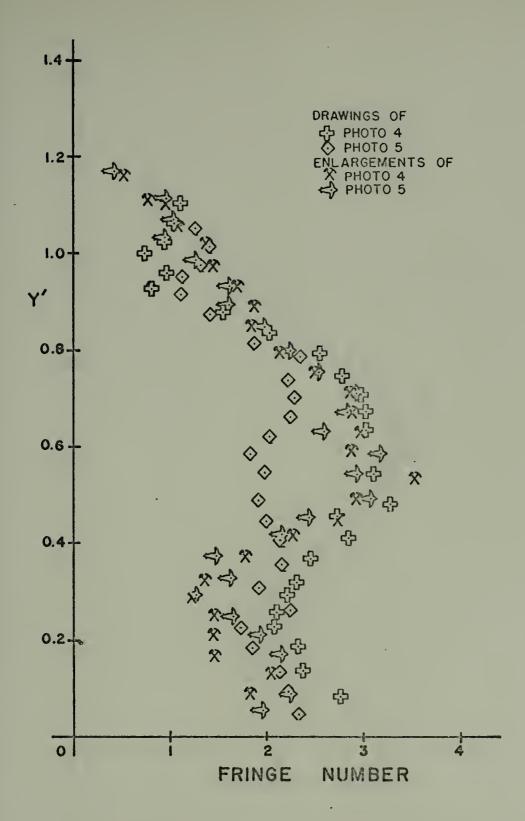
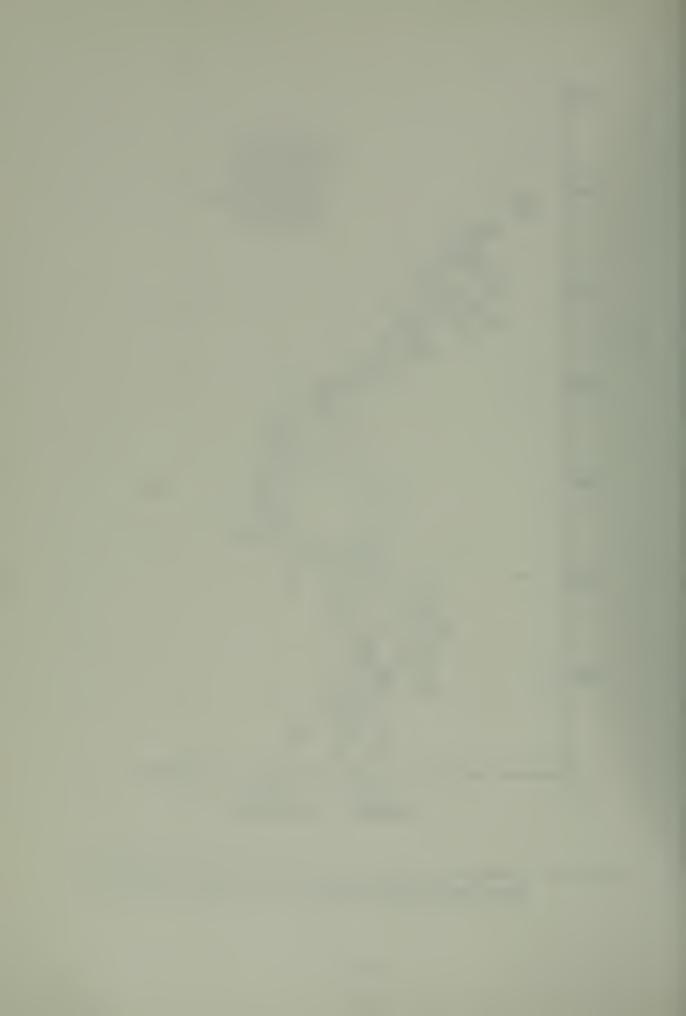


Figure 50. Comparison of Fringe Data Obtained from the Drawings and Photographic Enlargements of Interferogram Photographs 4 and 5 Aligned at Y' = 0.055



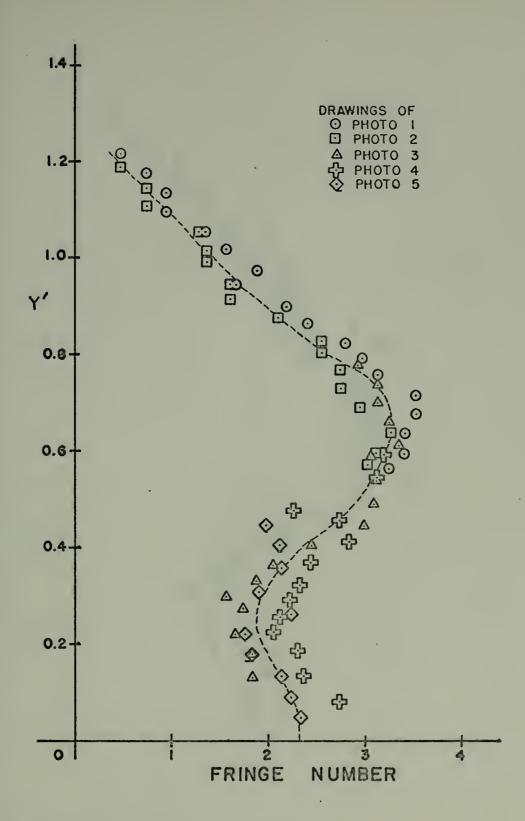
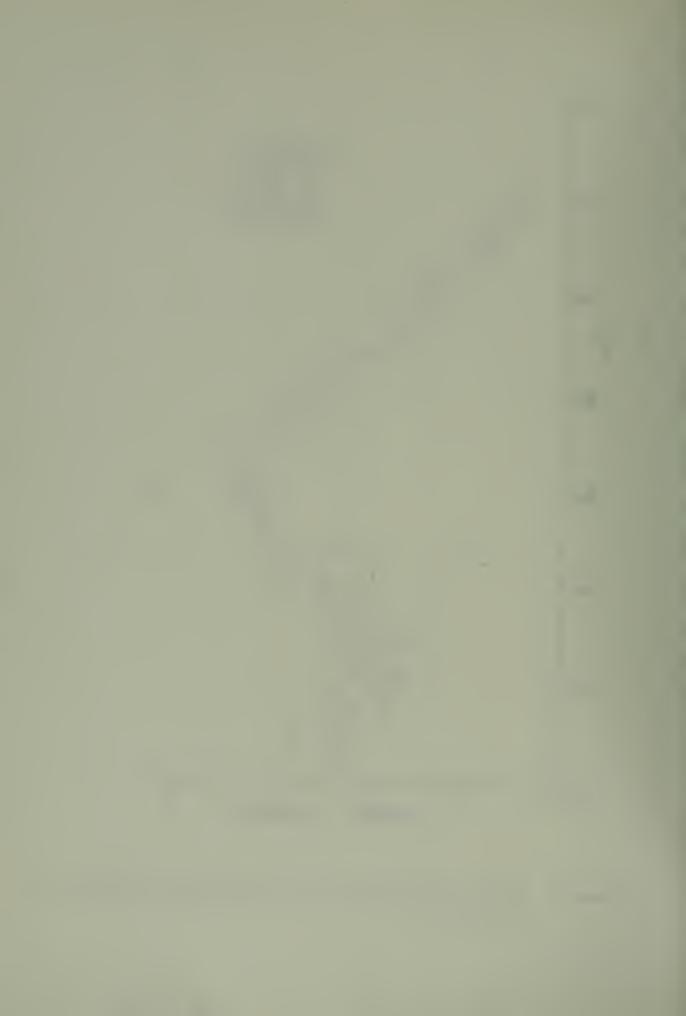


Figure 51. Fringe Number Across the Fin in the Z = 0.387 Plane for \$ = 0.387 Plane for



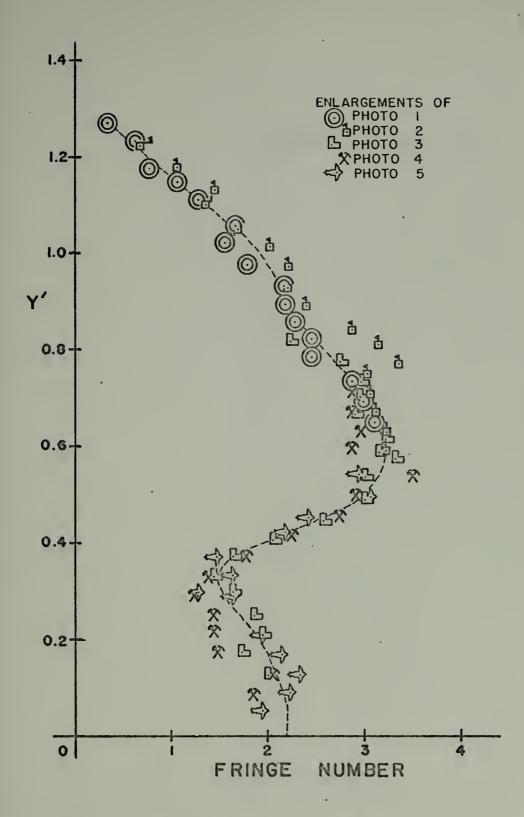
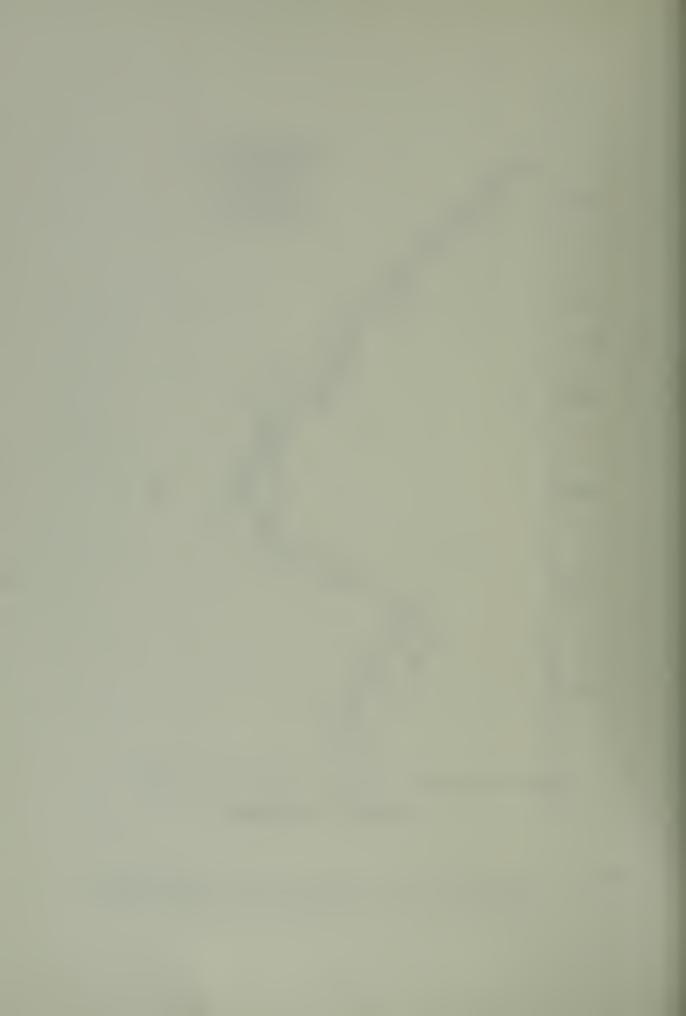


Figure 52. Fringe Number Across the Fin in the  $Z \approx 0.387$  Plane for  $\mathbf{\hat{s}} = 0^{\circ}$ , Mach 2.84 as Determined from the Photographic Enlargements of the Interferogram Photographs



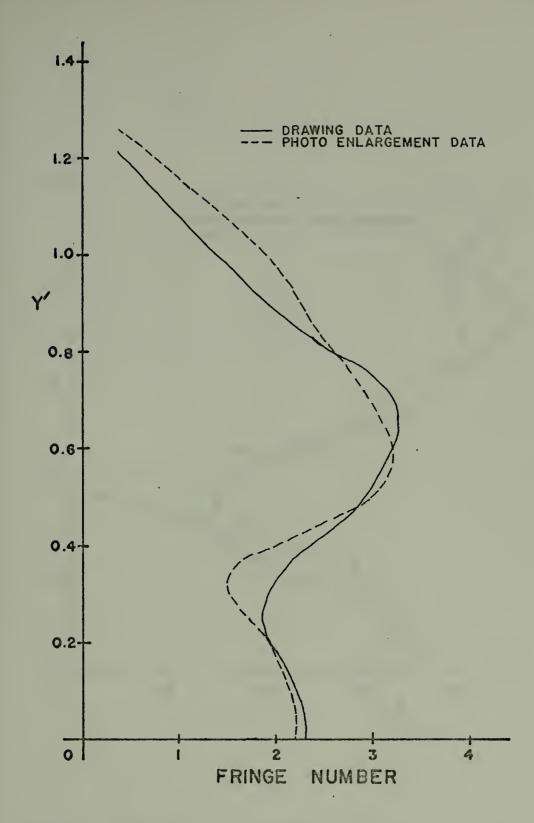
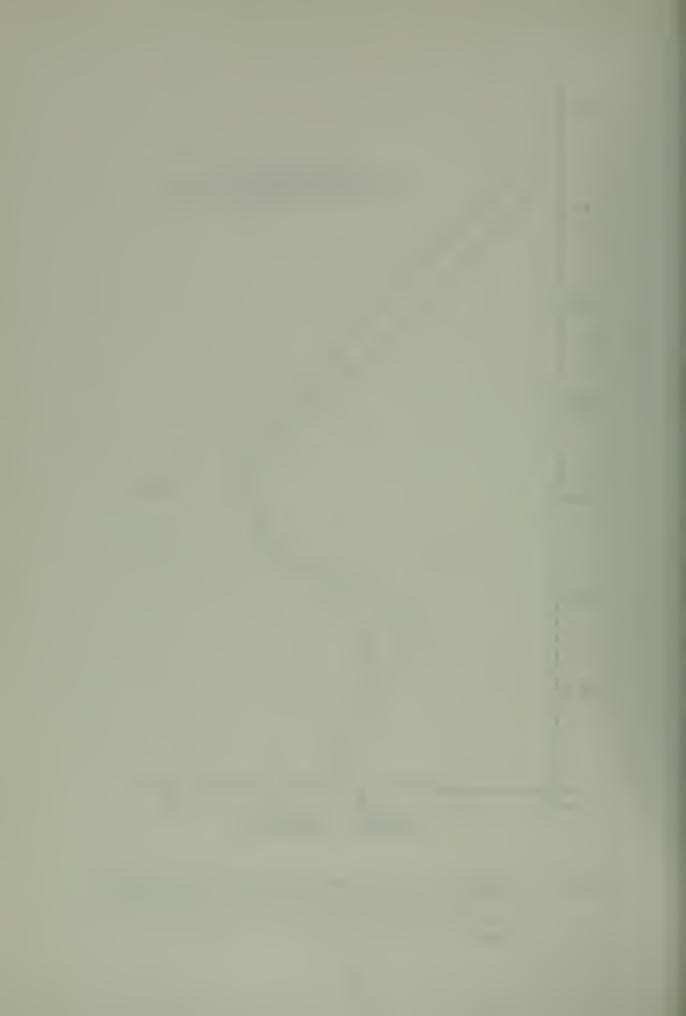


Figure 53. Comparison of the Fringe Numbers Across the Fin in the Z = 0.387 Plane for \$ = 0°, Mach 2.84 as Determined from the Drawings and Photographic Enlargements of the Interferogram Photographs



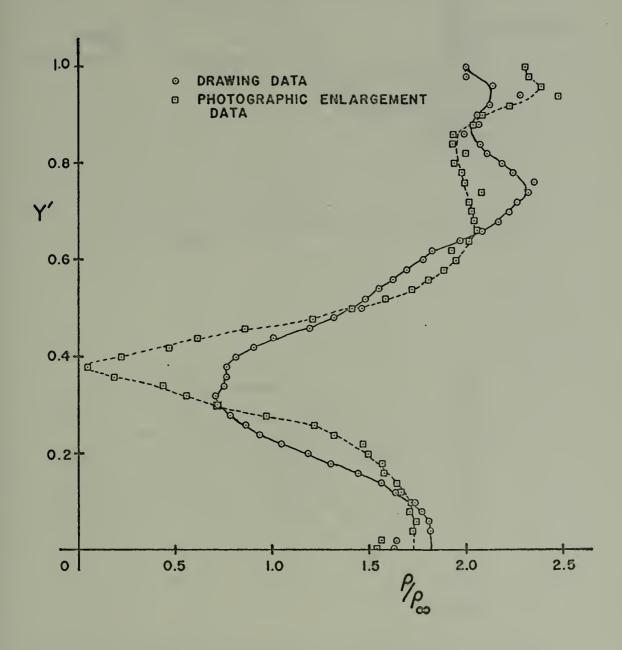
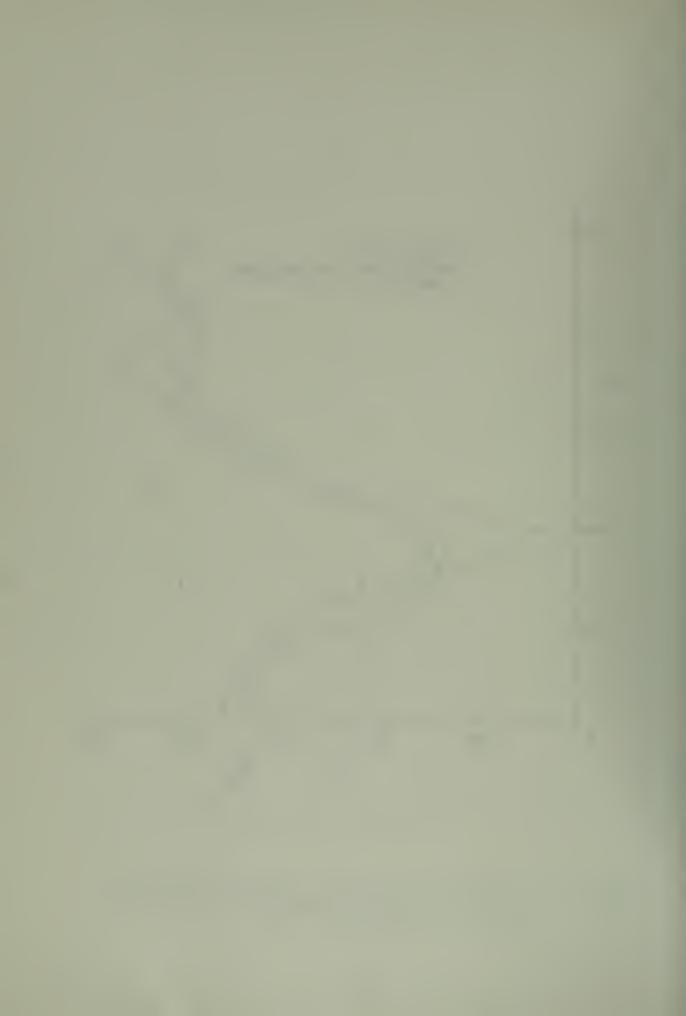


Figure 54. Comparison of the Density Distributions Calculated by HOLOVLR for an Axisymmetric Case



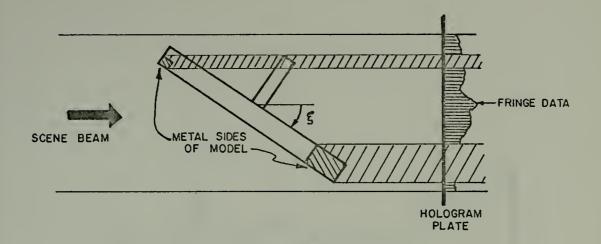


Figure 55. Schematic of the Model Center Section Rotated to Illustrate the Loss of Fringe Information Due to Model Shadows on the Hologram

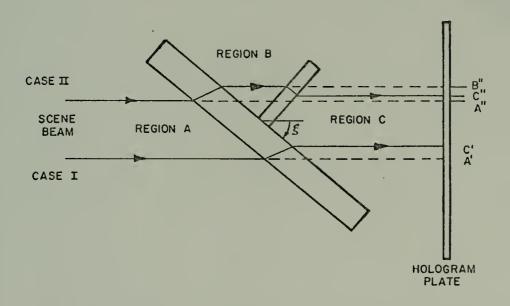
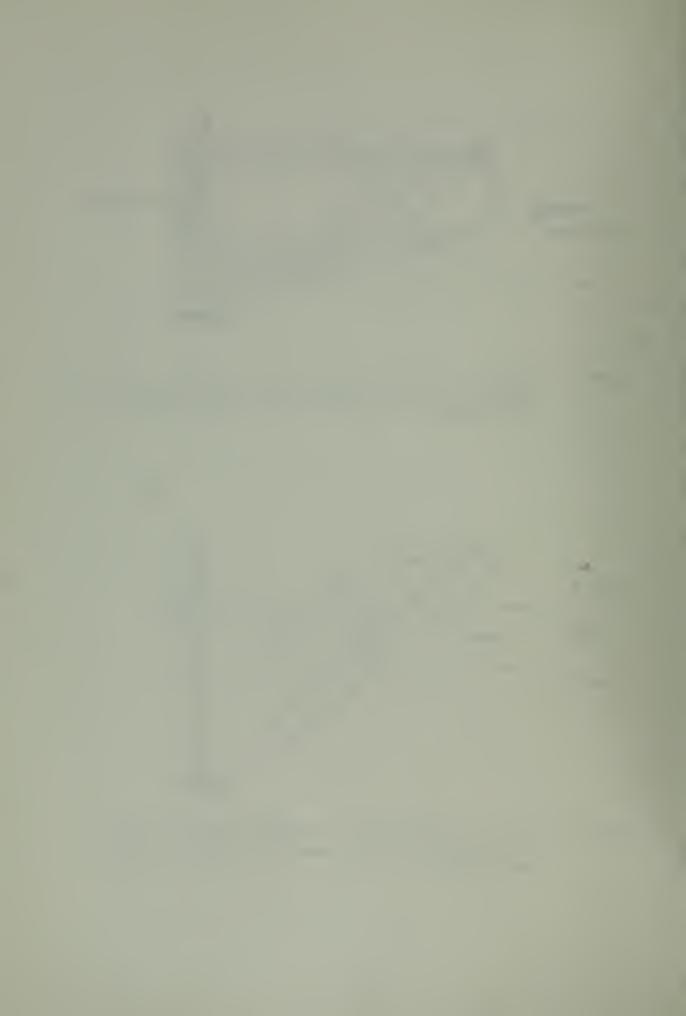


Figure 56. Schematic Illustrating the Problem of Different Fringe Information Being Superimposed on One Beam Caused by the Model Plastic



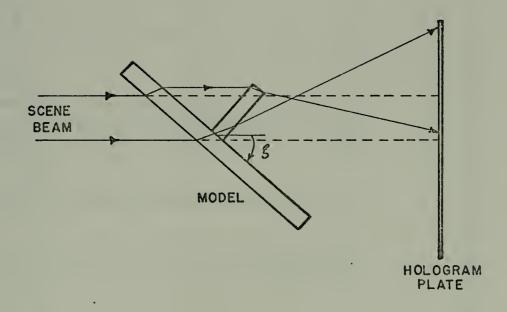


Figure 57. Schematic Illustrating the Scene Beam Refraction in the Fin Root and Tip Areas



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# Tabulation of Fringe Shift Data Taken From the Drawing of Photo Table I.

22.36 22.36 22.36 22.36 22.36 22.36 23.36

1.110 1.208 1.352 1.737 1.641 1.641 2.607 3.090 3.041 3.090 3.041 3.090 2.202 2.260 2.202 2.260 1.545 1.545 1.545 1.545 1.641 1.545 1.641

340 .460 .590 .765 .860 .1090 .11.240 .11.350 .11.500 .11.620 .11.620 .11.620 .11.620 .11.620 .11.620 .11.620 .11.620 .11.620 .11.620 .12.620 .2.290 .2.290 .2.290 .2.290 .2.290 .2.290 .2.290 .2.290 .2.290 .2.290 .2.290

.190 .330 .570 .660 .745 .745 .1745 .11.180 .11.180 .11.430 .11.430 .11.430 .11.430 .11.430 .11.625 .11.533 .11.625 .11.625 .11.630 .1

Corrected Values

Amount of 1.1 Fringe Correction Required

Fringe Number 8 8 8 8

Θ

Location (inches)

Fringe

350. + O= O

Corrected Distance from Tables

Distance Above Aligned Plane

Measured (inches)

Line No.

wing height Normalized

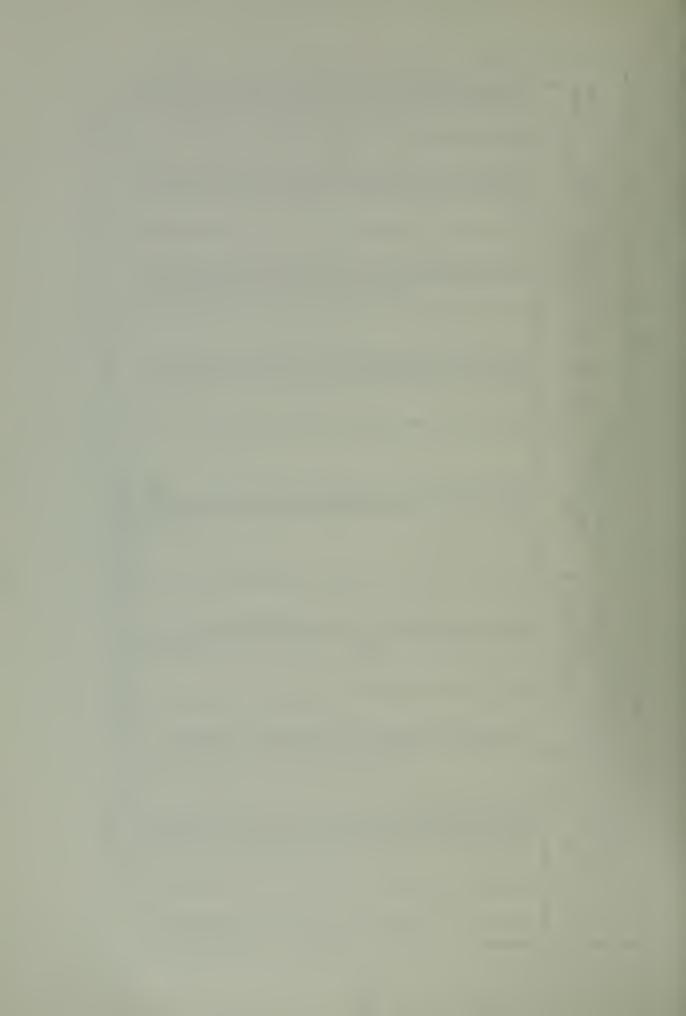
Fringe Change f = Q -

1.641 1.255 1.208 1.965 1.013



									_									_	_				_				
	Fringe Number g = f	1.867	2.050	1.482	1.434	1.434	1.242	1.386	1.771	2.252	2.738	2.926	3.503	2.887	2.984	2.887	2.887	2.502	2.117	1.829	1.877	1.704	1.444	1.396	1.059	٠22.	.529
	Fringe Change f = \$ - \$	.194	.214	.154	.149	.149	.129	.144	.184	.234	.284	.304	.364	.300	.310	.300	.300	.260	.220	.190	.195	.170	.150	.145	.110	080	.055
	Fringe Location (inches)	.270	.390	.450	.560	.650	.725	.830	086.	1.140	1.290	1.420	1.580	1.670	1.770	1.860	1.960	2.020	2.100	2.190	2.305	2.390	2.480	2.590	2.680	2.760	2.860
	γ' Φ = Φ + .055	.085	.124	.172	.217	.253	.290	.326	.369	.412	.451	.495	.535	.595	.631	029.	.710	.749	962.	.843	.887	.930	.974	1.019	1.068	1.111	1.161
	Corrected Distance from Tables ©	.030	690.	.117	.162	.198	.235	.271	.314	.357	.396	047.	.480	.540	.576	.615	.655	.694	.741	.788	.832	.875	.919	796.	1.013	1.056	1.106
Plane	Normalized     (wing height)	.032	.073	.123	.171	.208	.247	.285	.330	.376	.417	.463	.505	.568	909.	.647	.689	.730	.780	.830	.876	.921	.967	1.015	1.066	1.112	1.164
Above Aligned	Corrected to Free Stream (inches)	.076	.176	.296	.411	.501	.596	989.	962.	906.	1.006	1.116	1.216	1.370	1.460	1.560	1.660	1.760	1.880	2.000	2.110	2.220		2.445	2.570	2.680	2.805
Distance	Measured (inches)	*190*	*362.	.410*	.525*	.615*	.710*	*008.	*016.	1.020*	1.120*	1.230*	1.330*	1.370	1.460	1.560	1.660	1.760	1.880	2.000	2.110	2.220	2.330	•	2.570	2.680	2.805
	Line No.	1	2	3	7	2	9	7	<b>∞</b>	6	10		12	13	14	15	16	17	18	19	20	21	22	23	24	25	26

Tabulation of Fringe Shift Data Taken From the Photographic Enlargement of Photo 4 \* Locations to be corrected to free stream conditions Table II.



## APPENDIX A

# REDUCTION OF AN INTERFEROGRAM TO OBTAIN FRINGE SHIFT DATA

The fringe shift reduction process was accomplished using two techniques. The first involved projecting the interferogram negative onto a sheet of white paper using a photo-enlarger. The light fringes offered the best contrast and were therefore traced out in Figures 30-34. each drawing it was necessary to begin tracing the fringes above the fin and work towards the fin root since the transition across the fin tip determined the correct connection of the fringes across the fin leading edge shock. In order to determine the fringe change, one fringe line in the free stream region forward of the fin which appeared the straightest and paralleled the majority of other fringe lines was selected. straight line, called the fringe reference line, was drawn over its centerline and extended to cross the y' axis. The remaining reference lines were then drawn parallel to the first and along the centerlines of the remaining free stream fringes. In reducing the drawings it was not realized until later that the free stream fringe patterns before and after the plate leading edge Prandtl-Meyer expansion differed considerably. This effect was taken into account later.

A ruler scaled to 0.01 inches was then placed along the y' axis and the distances of the fringes and reference lines above and below the aligned y' plane were then recorded in Table I for Photograph 4. The fin width and length were then measured and the average values were recorded. The average fringe interval was determined by measuring the distance between the first and last fringes used and dividing by the number of intervals. The fringe change was found by subtracting the fringe crossing



point. The fringe number was calculated by dividing the change by the average fringe interval. The reference line location was then normalized with respect to the measured fin height. Since the actual fin height was very close to one inch, the above number was considered to be the number of inches above or below the aligned plane. Thus the table computing the tunnel wall and grid refraction displacements in Appendix B could be entered to determine the actual reference line location with respect to the aligned plane of the drawing (see Figure 29). The locations were then converted to the y' axis system by adding the normalized location of the aligned plane. After realizing that not all the reference lines were referenced to the free stream density forward of the plate leading edge, the enlarged photographs (Figures 35-39) were checked against the drawings. It was found that the Prandtl-Meyer expansion caused a fringe number change of approximately 1.1. Consequently each reference line in the drawing was compared with the enlarged photograph and an appropriate percentage of the 1.1 fringe number was used as a correction (see Table I). The calculated fringe numbers had the correction fringe number added to them while the y' locations were corrected bу

The second reduction technique was to use enlarged photographs made from the interferogram negatives to obtain the fringe change. The fringe lines were first traced over lightly with a pencil and then verified against the other photographs to ensure correct tracing. A datum fringe reference line was chosen as before and drawn. The remaining reference lines for the upper free stream fringe lines forward of the Prandtl-Meyer expansion were then drawn parallel to the datum. When it became impossible



to use fringe lines forward of the expansion, the reference lines were then drawn along the centerline of the fringes between the expansion and the fin. The same ruler was used to obtain the fringe crossing points, the reference crossing points, fin measurements, and the average fringe interval and they were recorded in Table 2. The reference line crossing points were then corrected to free stream conditions, using a 1.1 Fringe number correction for those referenced to the pattern between the expansion and fin leading edge. The fringe number and reference fin locations were calculated as before. This method was considered more accurate because the reference lines and fringe lines could easily be rechecked for accuracy and corrected in the event that a fringe line was traced incorrectly or a reference line was misaligned. The accuracy was directly proportional to the hologram resolution which was not true for the drawings since the reference lines are drawn parallel to handdrawn fringe lines.



### APPENDIX B

# CALCULATION OF TUNNEL WALL AND GRID PLASTIC REFRACTION CORRECTION

In the interferogram photographs used to obtain the fringe change, every point off the alignment axis will be slightly distorted due to the plastic tunnel wall and grid. This effect is illustrated in Figure 19.

From Snells Law of Refraction, the angles of incidence and refraction are related by

$$\sin \alpha = n \sin \beta$$
 (B-1)

where n is the index of refraction between plastic and air. Then can be written

$$\beta = \sin^{-1} \left( \frac{\sin \alpha}{n} \right)$$
 (B-2)

Since the tunnel wall and grid have the same index of refraction, they can be considered on material with thickness t = ab in Figure 19. Then consider the height bd which can be written

$$bd = t tan \propto$$
 (B-3)

The beam displacement,  $\Delta y$ , is then

$$\Delta Y = bd - bc = t \tan \alpha - t \tan \beta$$
 (B-4)

but tan 

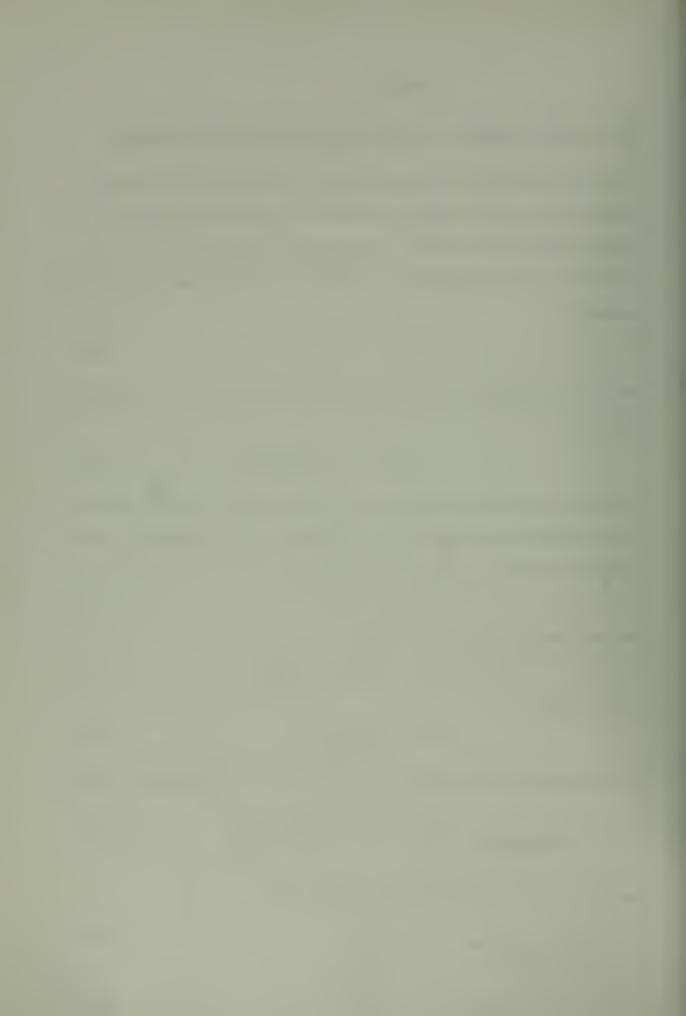
✓ is

Combining Equation (B-2), (B-4), and (B-5) the beam displacement becomes

$$\Delta Y = t \left[ \frac{Y \text{ observed}}{L} + \tan \left( \sin^{-1} \left( \frac{\sin \alpha}{n} \right) \right) \right]$$
 (B-6)

The true location of the observed point is then

$$Y_{\text{true}} = Y_{\text{observed}} - \Delta Y$$
 (B-7)



A FORTRAN computer program was written to generate a table giving the true locations versus the observed locations. In the program the constants and variables from the above equations were defined as

Y<sub>observed</sub> = Y

Y<sub>true</sub> = YTRUE

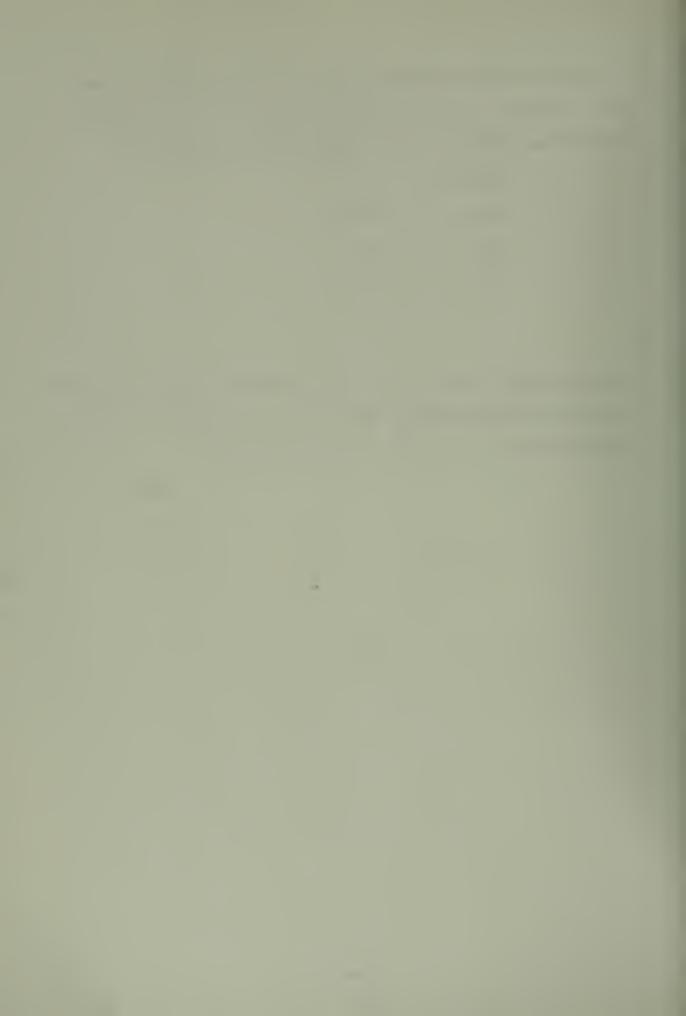
Δ Y = DY

Δ = ALFA

L = L

n = N

Since the computer cannot calculate Equations (B-2) and (B-6) as written, they were constructed by parts using such letters as AA, AB, etc. The program and tables are included in the next few pages.



```
COCCOCCOCCCCCCCC CCC
                **
                           THIS PROGRAM COMPUTES THE CORRECTION FOR THE OFF AXIS TUNNEL WALL PARALLAX. THE FOLLOWING INPUT PARAMETERS ARE REQUIRED:

L = DISTANCE FROM HOLOGRAM PLANE TO THE MODEL CENTER LINE
                **
                                                                                                                                                             **
                **
                                                                                                                                                             **
                ホネ
                                                                                                                                                             **
                **
                                                                                                                                                             **
                **
                                                                                                                                                             **
                                     INDEX OF REFRACTION FOR TUNNEL WALL
THICKNESS OF TUNNEL AND GRID PLEXIGLAS
OBSERVED DISTANCE OF THE POINT IN THE
PHOTOGRAPH FROM THE ALIGNED POINT
JE = TRUE DISTANCE FROM ALIGNED POINT T
THE POINT ON THE MODEL PLANE
( = MAXIMUM DISTANCE OF INTEREST FROM T
                **
                                                                                                                                                             **
                **
                                                                                                                                                             **
                **
                                                                                                                                                             **
                **
                                                                                                                                                             **
                **
                           YTRUE
                                                                                                                                                             **
                                                                                                                                        TO
                **
                                                                                                                                                             **
                                                                                                                           FROM THE
                **
                           YMAX =
                                                                                                                                                             **
                                                                  POINT
                **
                                              ALIGNED
                                                                                                                                                             **
                **
                                                                                                                                                             **
                REAL*4 L,N
                        INPUT PARAMETERS
                L=15.0
                N=1.5
T=2.25
                YMAX=2.5
PI=3.141593
                        OUTPUT FORMAT
     WRITE(6,100)L,N,T

100 FORMAT('1',T25,'THE FOLLOWING TABLE IS TO ACCOUNT FOR'

1' THE PARALLAX'/T25,'ERROR IN VIEWING THE MODEL'

2'THROUGH GLASS WALLS AT'/T25,'ANY OTHER POINT THAN '

3'AT THE ALIGNMENT POINT. INPUT'/T25,'PARAMETERS ARE:'

4/T35,'(1) L = ',F6.3,' INCHES'/T35,'(2) N = ',F6.3,'

5' INCHES'/T35,'(3) T = ',F6.3,' INCHES')

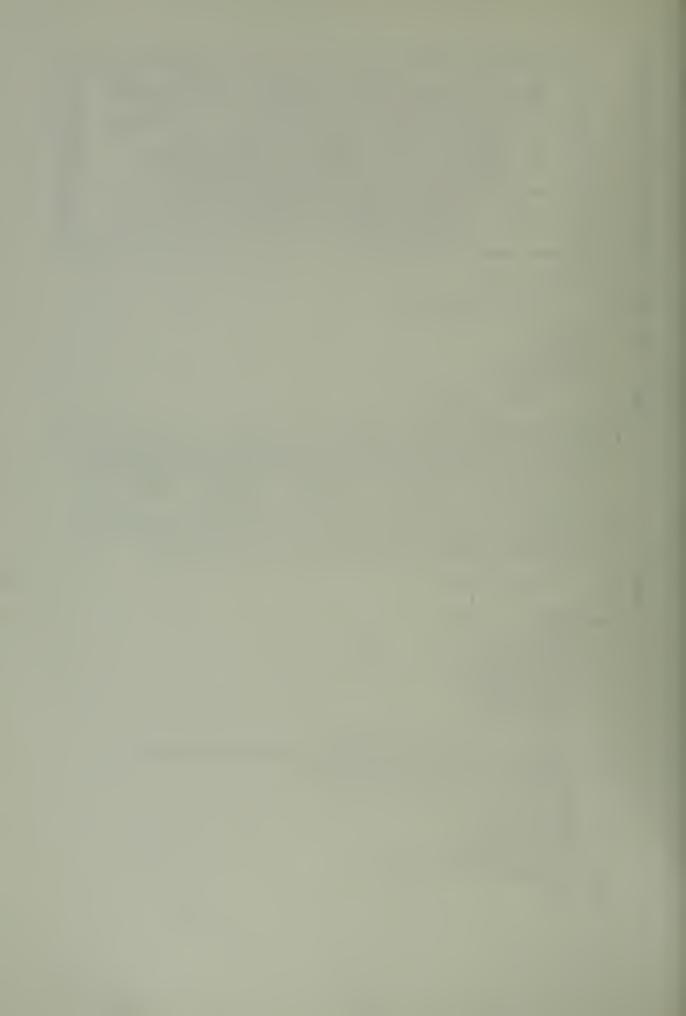
WRITE(6,110)

110 FORMAT(' ',T24,'OBSERVED',T40,'TRUE',T54,'ERROR',T70,

1'RAY'/T24,'DISTANCE',T38,'DISTANCE',T69,'ANGLE'/T24,

2'(INCHES)',T38,'(INCHES)',T53,'(INCHES)',T67,

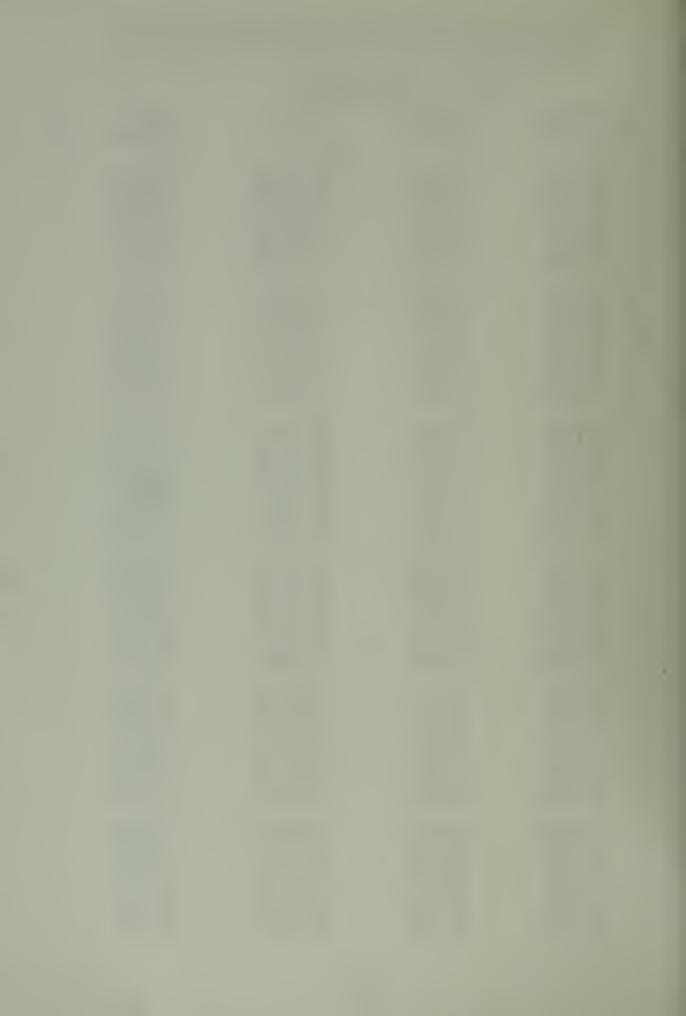
3'(DEGREES)'/)
                        CALCULATIONS
                Y=0.0
DO 15
AA=Y/L
                              I = 1,2000
        10
                 AB=ATAN(AA)
                ALFA=AB*180.0/PI
AC=SIN(AB)/N
BETA=ARSIN(AC)
AD=TAN(BETA)
                 DY=T*(AA-AD)
YTRUE=Y-DY
C
                WRITE(6,200)Y,YTRUE,DY,ALFA
FORMAT('',T24,F7.4,T38,F7.4,T53,F8.6,T67,F8.4)
IF(MOD(I,10).EQ.0)WRITE(6,201)
FORMAT('')
      200
      201
                 IF (MOD(1,60).EQ.0)GO TO 11
                WRITE(6,202)
FORMAT('1')
WRITE(6,110)
Y=.002*FLOAT(I)
IF(Y.GE.YMAX)GO TO 20
CONTINUE
STOP
                 GO TO 12
      11
202
         15
                 END
```



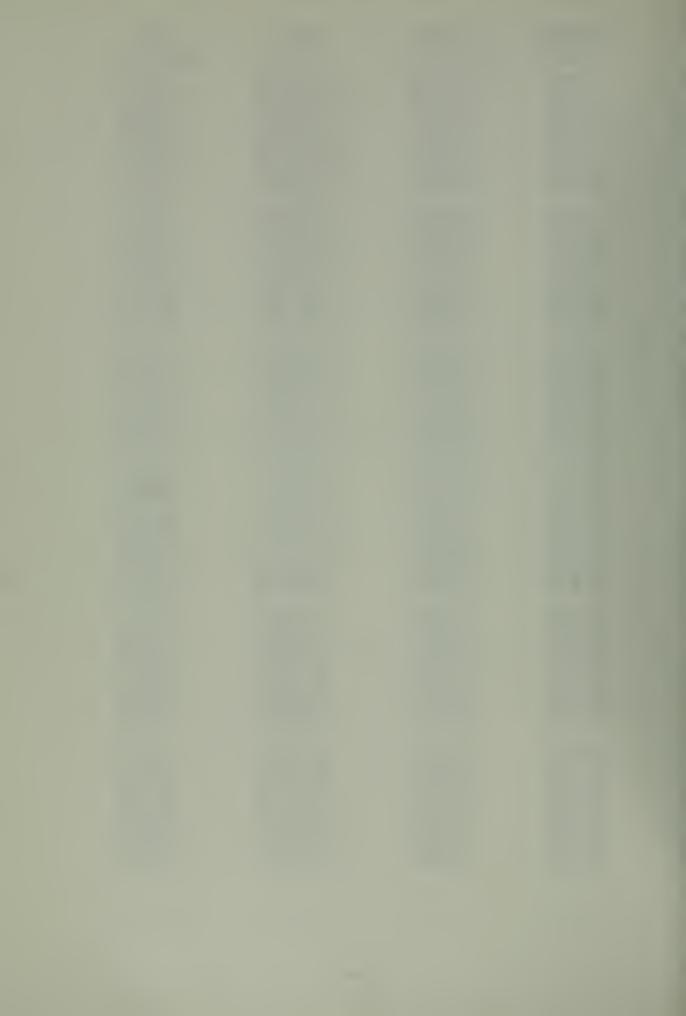
THE FOLLOWING TABLE IS TO ACCOUNT FOR THE PARALLAX ERROR IN VIEWING THE MODEL THROUGH GLASS WALLS AT ANY OTHER POINT THAN AT THE ALIGNMENT POINT. INPUT PARAMETERS ARE:

(1) L = 15.000 INCHES

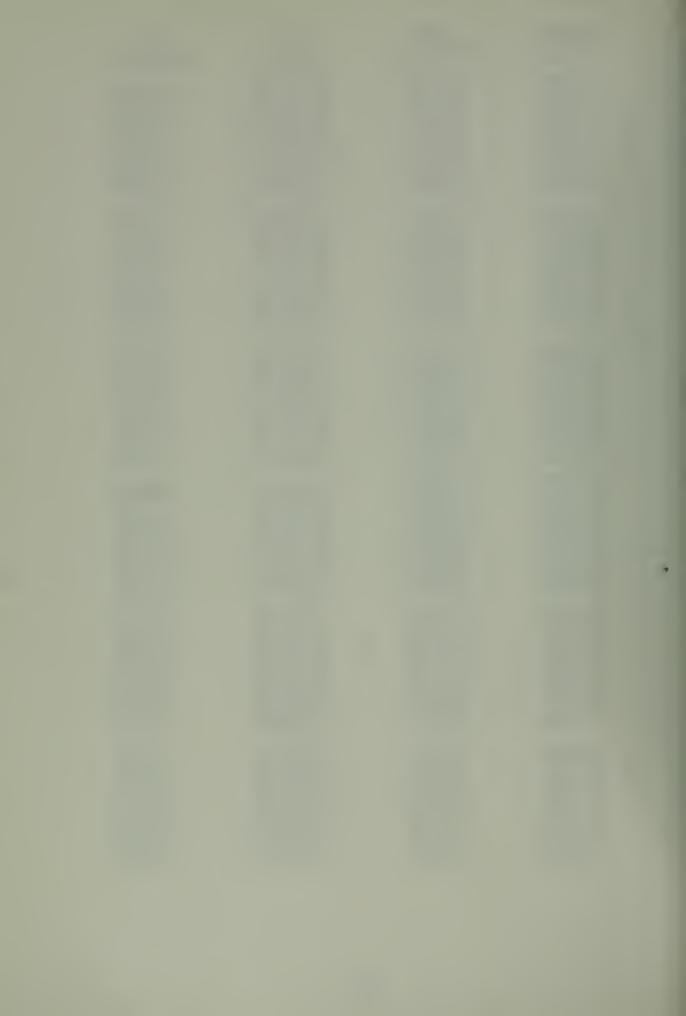
T ANAMETER	(1) L = 15.000 (2) N = 1.500 (3) T = 2.250	INCHES INCHES INCHES ERROR	
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
0.0 0.0020 0.0040 0.0060 0.0080 0.0120 0.0140 0.0160 0.0180	0.0 0.0019 0.0038 0.0057 0.0076 0.0095 0.0114 0.0133 0.0152 0.0171	0.0 0.000100 0.000200 0.000300 0.000400 0.000500 0.000600 0.000700 0.000800 0.000900	0.0 0.0076 0.0153 0.0229 0.0306 0.0382 0.0458 0.0535 0.0681
0.0200 0.0220 0.0240 0.0260 0.0280 0.0300 0.0320 0.0340 0.0360 0.0380	0.0190 0.0209 0.0228 0.0247 0.0266 0.0285 0.0304 0.0323 0.0342	0.001000 0.601100 0.001200 0.001300 0.001400 0.001500 0.001600 0.001700 0.001800 0.001900	0.0764 0.0840 0.0917 0.0993 0.1070 0.1146 0.1222 0.1299 0.1375 0.1451
0.0400 0.0420 0.0440 0.0460 0.0480 0.0500 0.0520 0.0540 0.0560 0.0580	0.0380 0.0399 0.0418 0.0437 0.0456 0.0475 0.0494 0.0513 0.0532 0.0551	0.002000 0.002100 0.002200 0.002300 0.002400 0.002500 0.002600 0.002700 0.002800 0.002900	0.1528 0.1604 0.1681 0.1757 0.1833 0.1910 0.1986 0.2063 0.2139 0.2215
0.0600 0.0620 0.0640 0.0660 0.0680 0.0700 0.0720 0.0740 0.0740 0.0780	0.0570 0.0589 0.0608 0.0627 0.0646 0.0665 0.0684 0.0703 0.0722	0.003000 0.003100 0.003200 0.003300 0.003400 0.003500 0.003600 0.003700 0.003800 0.003900	0.2292 0.2368 0.2445 0.2521 0.2597 0.2674 0.2750 0.2827 0.2903 0.2979
0.0800 0.0820 0.0840 0.0860 0.0880 0.0900 0.0920 0.0940 0.0960 0.0980	0.0760 0.0779 0.0798 0.0817 0.0836 0.0855 0.0874 0.0893 0.0912	0.004000 0.004100 0.004200 0.004300 0.004400 0.004500 0.004600 0.004700 0.004800 0.004900	0.3056 0.3132 0.3209 0.3285 0.3361 0.3438 0.3514 0.3590 0.3667 0.3743
0.1000 0.1020 0.1040 0.1060 0.1080 0.1100 0.1120 0.1140 0.1160 0.1180	0.0950 0.0969 0.0988 0.1007 0.1026 0.1045 0.1064 0.1083 0.1102 0.1121	0.005000 0.005100 0.005200 0.005300 0.005400 0.005500 0.005600 0.005700 0.005800 0.005900	0.3820 0.3896 0.3972 0.4049 0.4125 0.4202 0.4278 0.4354 0.4431 0.4507



OBSERVED	TRUE	ERROR	RAY
DISTANCE (INCHES)	DISTANCE (INCHES)	(INCHES)	(DEGREES)
0.1200	0.1140	0.006000	0.4584
0.1220	0.1159	0.006100	0.4660
0.1240	0.1178	0.006200	0.4736
0.1260	0.1197	0.006300	0.4813
0.1280	0.1216	0.006400	0.4889
0.1300	0.1235	0.006500	0.4966
0.1320	0.1254	0.006600	0.5042
0.1340	0.1273	0.006700	0.5118
0.1360	0.1292	0.006800	0.5195
0.1380	0.1311	0.006900	0.5271
0.1400 0.1420 0.1440 0.1460 0.1480 0.1500 0.1520 0.1540 0.1560 0.1580	0.1330 0.1349 0.1368 0.1387 0.1406 0.1425 0.1444 0.1463 0.1482 0.1501	0.007000 0.007100 0.007200 0.007300 0.007400 0.007500 0.007600 0.007700 0.007800 0.007901	0.5347 0.5424 0.5500 0.5577 0.5653 0.5729 0.5882 0.5959 0.6035
0.1600	0.1520	0.008001	0.6111
0.1620	0.1539	0.008101	0.6188
0.1640	0.1558	0.008201	0.6264
0.1660	0.1577	0.008301	0.6340
0.1680	0.1596	0.008401	0.6417
0.1700	0.1615	0.008501	0.6493
0.1720	0.1634	0.008601	0.6570
0.1740	0.1653	0.008701	0.6646
0.1760	0.1672	0.008801	0.6722
0.1780	0.1691	0.008901	0.6799
0.1800	0.1710	0.009001	0.6875
0.1820	0.1729	0.009101	0.6952
0.1840	0.1748	0.009201	0.7028
0.1860	0.1767	0.009301	0.7104
0.1880	0.1786	0.009401	0.7181
0.1900	0.1805	0.009501	0.7257
0.1920	0.1824	0.009601	0.7333
0.1940	0.1843	0.009701	0.7410
0.1960	0.1862	0.009801	0.7486
0.1980	0.1881	0.009901	0.7563
0.2000	0.1900	0.010001	0.7639
0.2020	0.1919	0.010101	0.7715
0.2040	0.1938	0.010201	0.7792
0.2060	0.1957	0.010301	0.7868
0.2080	0.1976	0.010401	0.7944
0.2100	0.1995	0.010501	0.8021
0.2120	0.2014	0.010601	0.8097
0.2140	0.2033	0.010701	0.8174
0.2160	0.2052	0.010801	0.8250
0.2180	0.2071	0.010901	0.8326
0.2200	0.2090	0.011001	0.8403
0.2220	0.2109	0.011101	0.8479
0.2240	0.2128	0.011201	0.8556
0.2260	0.2147	0.011301	0.8632
0.2280	0.2166	0.011401	0.8708
0.2300	0.2185	0.011502	0.8785
0.2320	0.2204	0.011602	0.8861
0.2340	0.2223	0.011702	0.8937
0.2360	0.2242	0.011802	0.9014
0.2380	0.2261	0.011902	0.9090



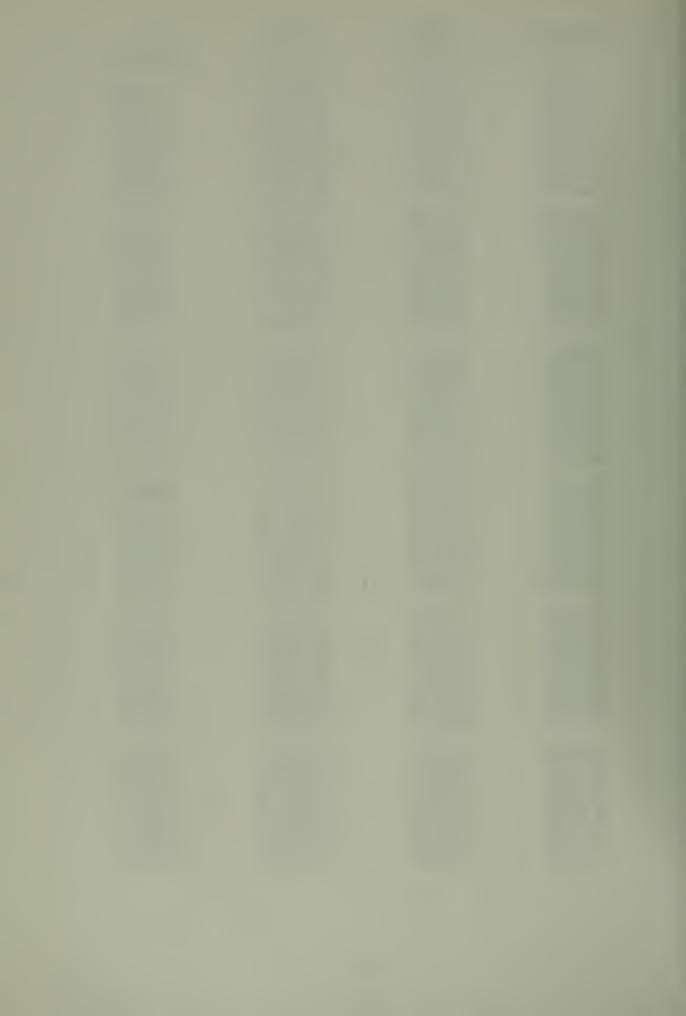
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
0. 2400	0.2280	0.012002	0.9167
0 • 2420 0 • 2440	n.2299	0.012102 0.012202 0.012302	0.9243 0.9319
0.2460 0.2480 0.2500	0.2318 0.2337 0.2356 0.2375	0.012302 0.012402	0.9396 0.9472
0.2500 0.2520 0.2540	0.2375 0.2394 0.2413	0.012402 0.012502 0.012602 0.012702 0.012802	0.9548 0.9625
0.2560 0.2580	0.2413 0.2432 0.2451	0.012702 0.012802 0.012902	0.9701 0.9778 0.9854
0.2600	0.2470	0.013002	0.9930
0.2620 0.2640	0 • 2489 0 • 2508	0.013102 0.013202 0.013302 0.013402	1.0007 1.0083
0.2660 0.2680	0 • 2527 0 • 2546 0 • 2565	0.013302 0.013402 0.013502	1.0159 1.0236 1.0312
0.2700 0.2720 0.2740	0.2565 0.2584 0.2603	0.013502 0.013602 0.013703	1.0312 1.0388 1.0465
0.2760 0.2780	0.2622 0.2641	0.013803 0.013903	1.0541 1.0618
0.2800 0.2820	0.2660 0.2679	0.014003 0.014103	1.0694 1.0770
0.2840	0 2608	0.014203 0.014303	1.0847 1.0923
0.2860 0.2880 0.2900	0.2736 0.2755	0.014403 0.014503	1.0999 1.1076
0.2920 0.2940	0.2717 0.2736 0.2755 0.2774 0.2793 0.2812	0.014603 0.014703	1.1152
0.2960 0.2980	0.2812	0.014803 0.014903	1.1305 1.1381
0.3000 0.3020	0.2850 0.2869	0.015003 0.015103	1.1458 1.1534
0.3040 0.3060 0.3080	0.2888 0.2907	0.015203 0.015304	1.1610 1.1687
0.3100	0.2907 0.2926 0.2945 0.2964	0.015404 0.015504 0.015604	1.1763 1.1839 1.1916
0.3100 0.3120 0.3140 0.3160 0.3180	0.2983 0.3002	0.015704 0.015804	1.1992 1.2069
	0.3021	0.015904	1.2145
0.3200 0.3220 0.3240 0.3260 0.3280 0.3300 0.3320 0.3340 0.3360 0.3380	0.3040 0.3059	0.016004 0.016104 0.016204	1.2221 1.2298 1.2374 1.2450 1.2527 1.2603 1.2679 1.2756 1.2832
0.3260	0.3076 0.3097 0.3116	0.016304 0.016404	1.2450
0.3300 0.3320	0.3135 0.3154	0.016504 0.016605	1.2603 1.2679
0.3340 0.3360	0.3078 0.3078 0.3097 0.3116 0.3135 0.3154 0.3173 0.3192 0.3211	0.016705 0.016805	1.2756 1.2832
	0.3211	0.016905	102,00
0.3400 0.3420 0.3440	0.3230 0.3249 0.3268 0.3287 0.3306 0.3325 0.3344 0.3363 0.3382 0.3401	0.017005 0.017105 0.017205	1.2985 1.3061 1.3138 1.3214 1.3290 1.3367 1.3443 1.3519 1.3596 1.3672
0 3460	0.3287 0.3306	0.017305 0.017405	1.3214 1.3290
0.3500 0.3520	0 • 3325 0 • 3344	0.017505 0.017605	1.3367 1.3443
0.3480 0.3500 0.3520 0.3540 0.3560 0.3580	0.3363 0.3382 0.3401	0.017705 0.017806 0.017906	1.3519 1.3596
J	0.0101	0.011700	1,0012



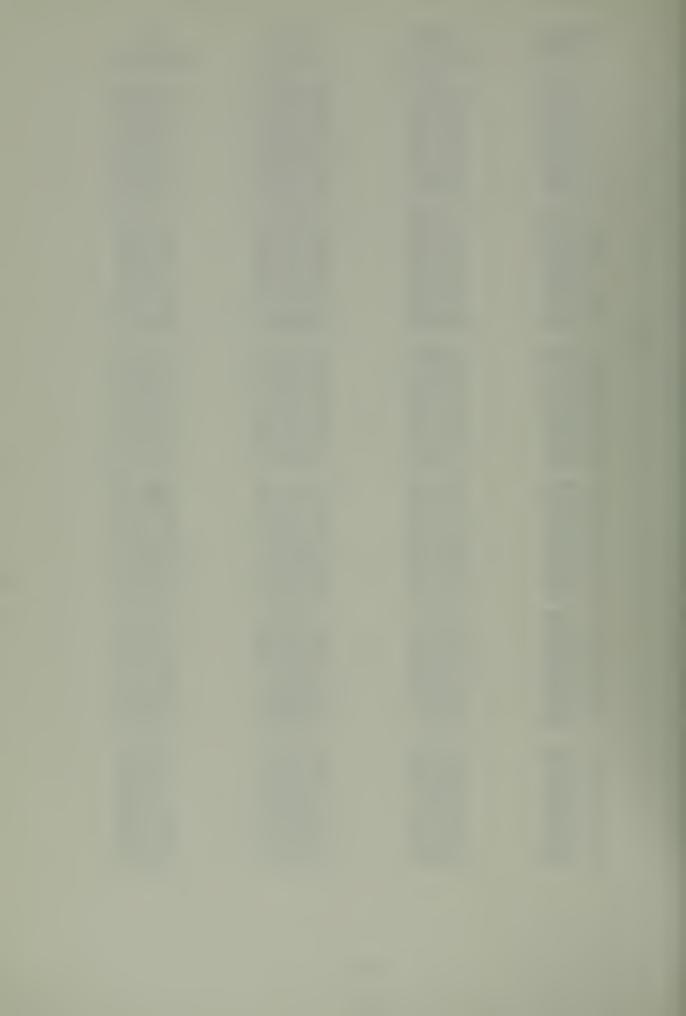
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
0.3600	0.3420	0.018006	1.3748
0.3620	0.3439	0.018106	1.3825
0.3640	0.3458	0.018206	1.3901
0.3660	0.34577	0.018306	1.3977
0.3680	0.3496	0.018406	1.4054
0.3700	0.3515	0.018506	1.4130
0.3720	0.3534	0.018606	1.4206
0.3740	0.3553	0.018706	1.4283
0.3760	0.3572	0.018807	1.4359
0.3780	0.3591	0.018907	1.4435
0.3800 0.3820 0.3840 0.3860 0.3880 0.3900 0.3920 0.3940 0.3960 0.3980	0.3610 0.3629 0.3648 0.3667 0.3686 0.3705 0.3724 0.3743 0.3762	0.019007 0.019107 0.019207 0.019307 0.019407 0.019507 0.019607 0.019708 0.019808 0.019908	1.4512 1.4588 1.4665 1.4741 1.4817 1.4894 1.4970 1.5046 1.5123 1.5199
0.4000	0.3800	0.020008	1.5275
0.4020	0.3819	0.020108	1.5352
0.4040	0.3838	0.020208	1.5352
0.4060	0.3857	0.020308	1.5504
0.4080	0.3876	0.020408	1.5581
0.4100	0.3895	0.020509	1.5657
0.4120	0.3914	0.020609	1.5733
0.4140	0.3933	0.020709	1.5810
0.4160	0.3952	0.020809	1.5886
0.4180	0.3971	0.020909	1.5962
0.4200 0.4220 0.4240 0.4260 0.4280 0.4300 0.4320 0.4340 0.4360 0.4380	0.3990 0.4009 0.4028 0.4047 0.4066 0.4085 0.4104 0.4123 0.4142	0.021009 0.021109 0.021209 0.021310 0.021410 0.021510 0.021610 0.021710 0.021810 0.021910	1.6039 1.6115 1.6191 1.6268 1.6344 1.6420 1.6497 1.6573 1.6649 1.6726
0.4400	0.4180	0.022011	1.6802
0.4420	0.4199	0.022111	1.6878
0.4440	0.4218	0.022211	1.6955
0.4460	0.4237	0.022311	1.7031
0.4480	0.4256	0.022411	1.7107
0.4500	0.4275	0.022511	1.7184
0.4520	0.4294	0.022611	1.7260
0.4540	0.4313	0.022712	1.7336
0.4560	0.4332	0.022812	1.7413
0.4580	0.4351	0.022912	1.7489
0.4600	0.4370	0.023012	1.7565
0.4620	0.4389	0.023112	1.7642
0.4640	0.4408	0.023212	1.7718
0.4660	0.4427	0.023312	1.7794
0.4680	0.4446	0.023413	1.7870
0.4700	0.4465	0.023513	1.7947
0.4720	0.4484	0.023613	1.8023
0.4740	0.4503	0.023713	1.8099
0.4760	0.4522	0.023813	1.8176
0.4780	0.4541	0.023913	1.8252



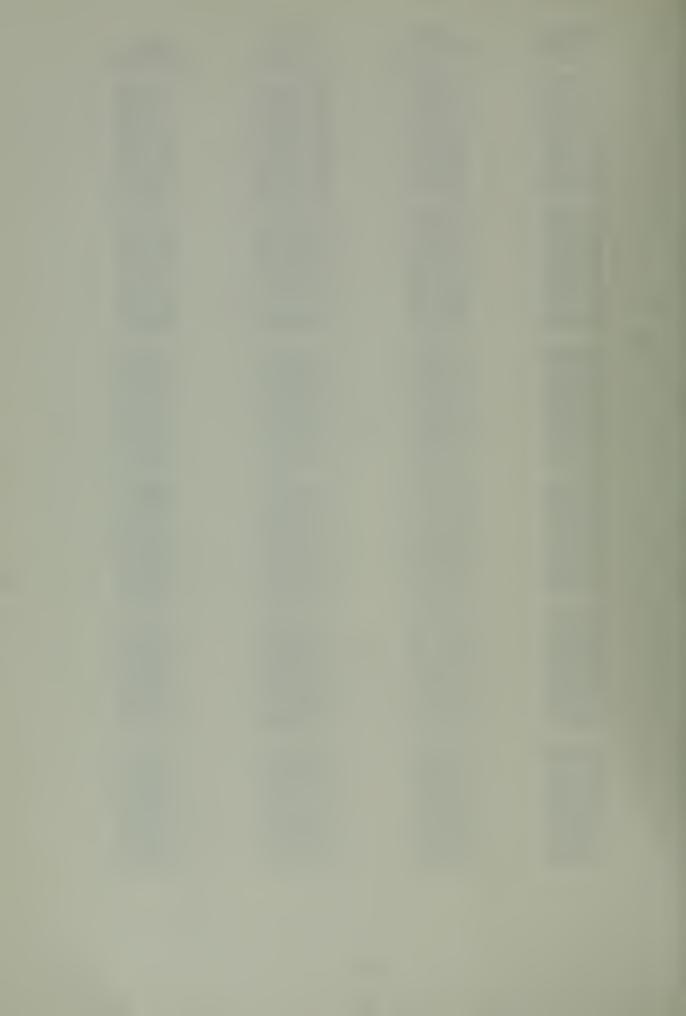
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
0.4800 0.4820 0.4840 0.4860 0.4880 0.4900 0.4920 0.4920 0.4940 0.4960 0.4980	0.4560 0.4579 0.4579 0.4617 0.4636 0.4655 0.4674 0.4693 0.4712 0.4731	0.024014 0.024114 0.024214 0.024314 0.024414 0.024515 0.024615 0.024715 0.024815 0.024915	1.8328 1.8405 1.8481 1.8557 1.8634 1.8710 1.8786 1.8863 1.8939 1.9015
0.5000 0.5020 0.5040 0.5060 0.5080 0.5100 0.5120 0.5140 0.5160 0.5180	0.4750 0.4769 0.4788 0.4807 0.4826 0.4845 0.4864 0.4883 0.4902 0.4921	0.025015 0.025116 0.025216 0.025316 0.025416 0.025516 0.025617 0.025717 0.025817 0.025917	1.9092 1.9168 1.9244 1.9320 1.9397 1.9473 1.9549 1.9626 1.9702
0.5200 0.5220 0.5240 0.5260 0.5280 0.53300 0.53320 0.53340 0.53360 0.53380	0.4940 0.4959 0.4978 0.4997 0.5016 0.5035 0.5054 0.5073 0.5073 0.5092	0.026017 0.026118 0.026218 0.026318 0.026418 0.026518 0.026619 0.026619 0.026719 0.026819 0.026919	1.9855 1.9931 2.0007 2.0083 2.0160 2.0236 2.0312 2.0389 2.0465 2.0541
0.5400 0.5420 0.5440 0.5480 0.5500 0.55540 0.55560 0.55580	0.5130 0.5149 0.5168 0.5187 0.5206 0.5225 0.5244 0.5263 0.5282 0.5301	0.027019 0.027120 0.027220 0.027320 0.027420 0.027521 0.027621 0.027721 0.027821 0.027921	2.0618 2.0694 2.0770 2.0846 2.0923 2.0999 2.1075 2.1152 2.1228 2.1304
0.5600 0.5620 0.5640 0.5660 0.5680 0.5700 0.5720 0.5720 0.5740 0.5760 0.5780	0.5320 0.5339 0.5358 0.5377 0.5396 0.5415 0.5434 0.5453 0.5472	0.028022 0.028122 0.028222 0.028322 0.028423 0.028523 0.028623 0.028723 0.028824 0.028924	2.1380 2.1457 2.1533 2.1609 2.1686 2.1762 2.1838 2.1914 2.1991 2.2067
0.5800 0.5820 0.5840 0.5860 0.5880 0.5900 0.5920 0.5940 0.5940 0.5980	0.5510 0.5529 0.5548 0.55667 0.5586 0.5605 0.56624 0.56643 0.56681	0.029024 0.029124 0.029225 0.029325 0.029425 0.029525 0.029626 0.029726 0.029826 0.029926	2.2143 2.2220 2.2296 2.2372 2.2448 2.2525 2.2601 2.2677 2.2754 2.2830



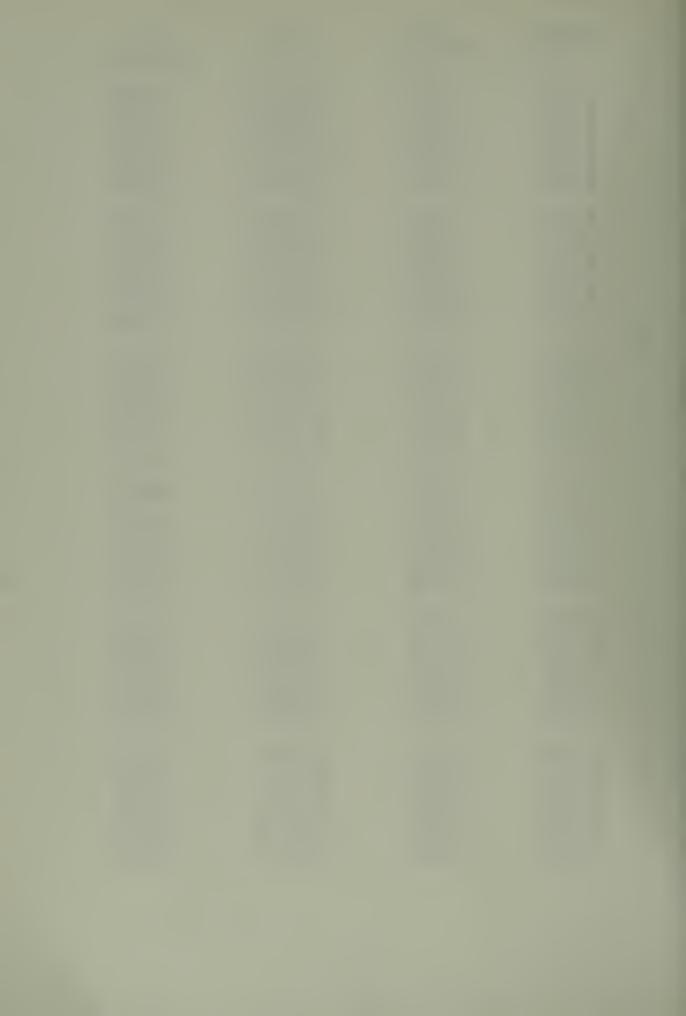
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
0.6000 0.6020 0.6040 0.6060 0.6080 0.6100 0.6120 0.6140 0.6160 0.6180	0.5700 0.5719 0.5738 0.5757 0.5776 0.5795 0.5814 0.5833 0.5852 0.5871	0.030027 0.030127 0.030227 0.030327 0.030428 0.030528 0.030628 0.030729 0.030829 0.030929	2 • 2906 2 • 2982 2 • 3059 2 • 3135 2 • 3211 2 • 3287 2 • 3364 2 • 3516 2 • 3593
0.6200 0.6220 0.6240 0.6260 0.6280 0.6300 0.6320 0.6340 0.6360 0.6380	0.5890 0.5909 0.5928 0.5947 0.5966 0.5985 0.6004 0.6023 0.6042	0.031029 0.031130 0.031230 0.031330 0.031431 0.031531 0.031631 0.031731 0.031832 0.031932	2.3669 2.3745 2.3821 2.3898 2.3974 2.4050 2.4126 2.4203 2.4279 2.4355
0.6400 0.6420 0.6440 0.6480 0.6500 0.6520 0.6540 0.6560 0.6580	0.6080 0.6099 0.6118 0.6137 0.6156 0.6175 0.6194 0.6213 0.6232 0.6251	0.032032 0.032133 0.032233 0.032333 0.032434 0.032534 0.032534 0.032634 0.032735 0.032835 0.032935	2.4431 2.4508 2.4584 2.4660 2.4736 2.4813 2.4889 2.4965 2.5041 2.5118
0.6600 0.6620 0.6640 0.6660 0.6680 0.6700 0.6720 0.6740 0.6760 0.6780	0.6270 0.6289 0.6308 0.6327 0.6346 0.6365 0.6384 0.6403 0.6422 0.6441	0.033035 0.033136 0.033236 0.0333336 0.033437 0.033537 0.033637 0.033738 0.033838	2.5194 2.5270 2.5346 2.53423 2.55499 2.5575 2.5651 2.5728 2.5804 2.5880
0.6800 0.6820 0.6840 0.6860 0.6880 0.6900 0.6920 0.6940 0.6960 0.6980	0.6460 0.6479 0.6498 0.6517 0.6536 0.6555 0.6574 0.6593 0.6612 0.6631	0.034039 0.034139 0.034239 0.034340 0.034440 0.034541 0.034641 0.034741 0.034842 0.034942	2.5956 2.6033 2.6109 2.6185 2.6261 2.6337 2.6414 2.6490 2.6566 2.6642
0.7000 0.7020 0.7040 0.7060 0.7080 0.7100 0.7120 0.7140 0.7160 0.7180	0.6650 0.6669 0.6688 0.6707 0.6726 0.6745 0.6764 0.6783 0.6802 0.6821	0.035042 0.035143 0.035243 0.035343 0.0355444 0.035544 0.035645 0.035745 0.035845 0.035946	2.6719 2.6795 2.6871 2.6947 2.7024 2.7100 2.7176 2.7252 2.7328 2.7405



CBSERVED	TRUE	ERROR	RAY
DISTANCE (INCHES)	DISTANCE (INCHES)	(INCHES)	AÑĜLE (DEGREES)
0.7200 0.7220 0.7240 0.7260 0.7280 0.7300 0.7320 0.7340 0.7360 0.7380	0.6840 0.6859 0.6878 0.6897 0.6916 0.6935 0.6954 0.6973 0.6992	0.036046 0.036146 0.036247 0.036347 0.036448 0.036548 0.036648 0.036749 0.036849 0.036950	2.7481 2.7557 2.7633 2.7710 2.7786 2.7862 2.7938 2.8014 2.8091 2.8167
0.7400 0.7420 0.7440 0.7460 0.7480 0.7500 0.7520 0.7540 0.7560 0.7580	0.7029 0.7048 0.7067 0.7086 0.7105 0.7124 0.7143 0.7162 0.7181 0.7200	0.037050 0.037150 0.037251 0.037351 0.037452 0.037552 0.037652 0.037753 0.037853 0.037954	2.8243 2.8319 2.8395 2.8472 2.8548 2.8624 2.8700 2.8776 2.8853 2.8929
0.7600 0.7620 0.7640 0.7660 0.7680 0.7700 0.7720 0.7740 0.7760 0.7780	0.7219 0.7238 0.7257 0.7276 0.7295 0.7314 0.7333 0.7352 0.7371	0.038054 0.038155 0.038255 0.038255 0.038355 0.038456 0.038657 0.038657 0.038858 0.038958	2.9005 2.9081 2.9157 2.9234 2.9310 2.9386 2.9462 2.9538 2.9615 2.9691
0.7800 0.7820 0.7840 0.7860 0.7880 0.7900 0.7920 0.7940 0.7960 0.7980	0.7409 0.7428 0.7447 0.7466 0.7485 0.7504 0.7523 0.7542 0.7561	0.039059 0.039159 0.039259 0.039360 0.039460 0.039661 0.039661 0.039762 0.039862 0.039963	2.9767 2.9843 2.9919 2.9996 3.0072 3.0148 3.0224 3.0300 3.0376 3.0453
0.8000 0.8020 0.8040 0.8060 0.8080 0.8100 0.8120 0.8140 0.8160 0.8180	0.7599 0.7618 0.7637 0.7656 0.7675 0.7694 0.7713 0.7732 0.7751	0.040063 0.040164 0.040264 0.040365 0.040465 0.040566 0.040666 0.040767 0.040867	3.0529 3.0605 3.0681 3.0757 3.0834 3.0910 3.0986 3.1062 3.1138 3.1214
0.8200 0.8220 0.8240 0.8260 0.8280 0.8390 0.8320 0.8340 0.8360 0.8380	0.7789 0.7808 0.7827 0.7846 0.7865 0.7884 0.7903 0.7922 0.7941	0.041068 0.041169 0.041269 0.041369 0.041470 0.041570 0.041671 0.041772 0.041872 0.041973	3.1291 3.1367 3.1443 3.1519 3.1595 3.1671 3.1748 3.1824 3.1900 3.1976



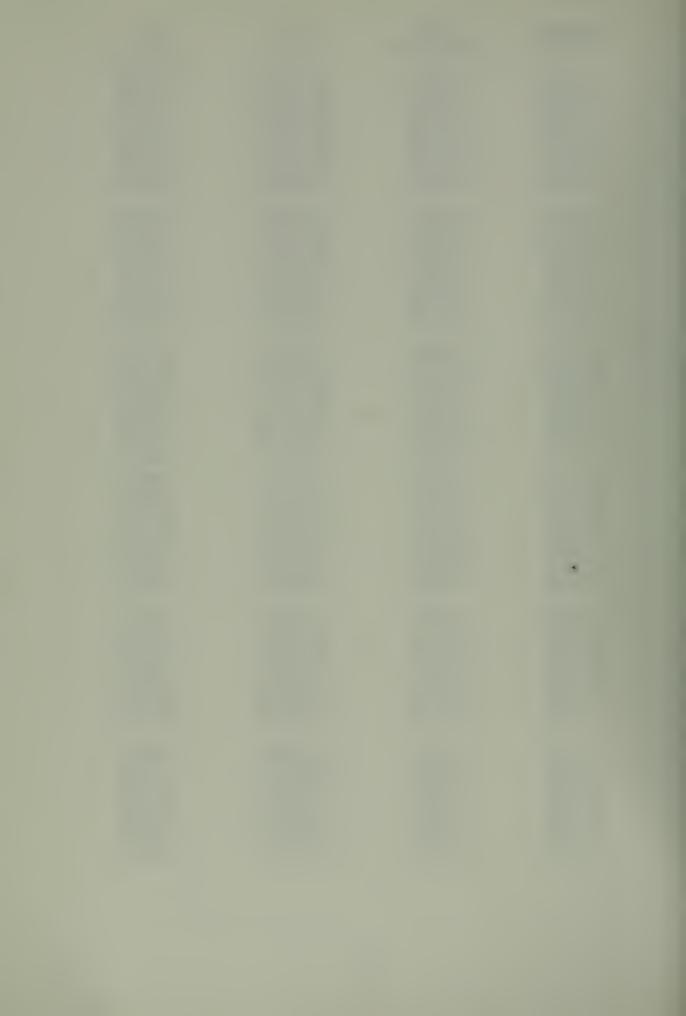
CBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
0 • 8400 0 • 8420 0 • 8440 0 • 8460 0 • 8500 0 • 8520 0 • 8540 0 • 8560 0 • 8580	0.7979 0.7998 0.8017 0.8036 0.8055 0.8074 0.8093 0.8112 0.8131 0.8150	0.042073 0.042174 0.042274 0.042375 0.042475 0.042576 0.042676 0.042777 0.042877 0.042978	3 • 2052 3 • 2128 3 • 2204 3 • 2281 3 • 2357 3 • 2433 3 • 2509 3 • 2585 3 • 2661 3 • 2738
0.8600 0.8620 0.8640 0.8660 0.8700 0.8720 0.8740 0.8760 0.8780	0.8169 0.8188 0.8207 0.8226 0.8245 0.8264 0.8283 0.8302 0.8321 0.8340	0.043078 0.043179 0.043280 0.043380 0.043481 0.043581 0.043682 0.043782 0.043883 0.043983	3 · 2814 3 · 2890 3 · 2966 3 · 3042 3 · 3118 3 · 3194 3 · 3270 3 · 3347 3 · 3423 3 · 3499
0.8800 0.8820 0.8840 0.8860 0.8890 0.8900 0.8920 0.8940 0.8960 0.8980	0.8359 0.8378 0.8397 0.8416 0.8435 0.8454 0.8473 0.8473 0.8511 0.8530	0.044084 0.044185 0.044285 0.044386 0.044486 0.044587 0.044688 0.044788 0.044889 0.044989	3.3575 3.3651 3.3727 3.3803 3.3880 3.3956 3.4032 3.4108 3.4184 3.4260
0.9000 0.9020 0.9040 0.9060 0.9080 0.9100 0.9120 0.9140 0.9160 0.9180	0.8549 0.8568 0.8587 0.8606 0.8625 0.8644 0.8663 0.8682 0.8701 0.8720	0.045090 0.045190 0.045191 0.045392 0.045492 0.045593 0.045693 0.045794 0.045895	3.4336 3.4412 3.4489 3.4565 3.4641 3.4717 3.4793 3.4869 3.4945 3.5021
0.9200 0.9220 0.9240 0.9260 0.9280 0.9300 0.9320 0.9340 0.9360 0.9380	0.8739 0.8758 0.8777 0.8796 0.8815 0.8834 0.8853 0.8872 0.8891	0.046096 0.046197 0.046297 0.046398 0.046499 0.046599 0.046700 0.046800 0.046901 0.047002	3.5097 3.5174 3.5250 3.5326 3.55402 3.55478 3.55554 3.55554 3.5706 3.57706
0.9400 0.9420 0.9440 0.9460 0.9480 0.9500 0.9520 0.9540 0.9560 0.9580	0.8929 0.8948 0.8967 0.8986 0.9005 0.9024 0.9043 0.9062 0.9081 0.9100	0.047102 0.047203 0.047304 0.047404 0.047505 0.047606 0.047706 0.047807 0.047908 0.048009	3.5858 3.5935 3.6011 3.6087 3.6163 3.6239 3.6315 3.6391 3.6467 3.6543



OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
0.9600 0.9620 0.9640 0.9660 0.9680 0.9700 0.9720 0.9740 0.9760 0.9780	0.9119 0.9138 0.9157 0.9157 0.9195 0.9214 0.9233 0.9252 0.9271	0.048109 0.048210 0.048310 0.048411 0.048512 0.048613 0.048713 0.048814 0.048915 0.049015	3.6619 3.6695 3.6771 3.6848 3.6924 3.7000 3.7076 3.7152 3.7228 3.7304
0.9800 0.9820 0.9840 0.9860 0.9880 0.9900 0.9920 0.9940 0.9960 0.9980	0.9309 0.9328 0.9347 0.9366 0.9385 0.9404 0.9423 0.9442 0.9461 0.9480	0.049116 0.049217 0.049318 0.049418 0.049519 0.049620 0.049720 0.049821 0.049922 0.050023	3.7380 3.7456 3.7532 3.7608 3.7684 3.7760 3.7836 3.7913 3.7989 3.8065
1.0000 1.0020 1.0040 1.0060 1.0080 1.0100 1.0120 1.0140 1.0160 1.0180	0.9499 0.9518 0.9537 0.9556 0.9575 0.9594 0.9613 0.9632 0.9651	0.050123 0.050224 0.050325 0.050426 0.050526 0.050627 0.050728 0.050829 0.050929 0.051030	3.8141 3.8217 3.8293 3.8369 3.8445 3.8521 3.8597 3.8673 3.8749 3.8825
1.0200 1.0220 1.0240 1.0260 1.0280 1.0300 1.0320 1.0340 1.0360 1.0380	0.9689 0.9708 0.9727 0.9746 0.9765 0.9784 0.9803 0.9822 0.9841 0.9860	0.051131 0.051232 0.051332 0.051433 0.051534 0.051635 0.051735 0.051836 0.051937 0.052038	3.8901 3.89077 3.9053 3.9129 3.9205 3.9281 3.9357 3.9433 3.9509 3.9586
1.0400 1.0420 1.0440 1.0460 1.0480 1.0500 1.0520 1.0540 1.0560 1.0580	0.9879 0.9898 0.9917 0.9936 0.9955 0.9974 0.9993 1.0012 1.0031	0.052139 0.052240 0.052340 0.052441 0.052542 0.052643 0.052743 0.052743 0.052945 0.053046	3.9662 3.9738 3.9814 3.9890 3.9966 4.0042 4.0118 4.0194 4.0270 4.0346
1.0600 1.0620 1.0640 1.0660 1.0680 1.0700 1.0720 1.0740 1.0760 1.0780	1.0069 1.0088 1.0107 1.0125 1.0144 1.0163 1.0182 1.0201 1.0220 1.0239	0.053147 0.053248 0.053348 0.0533449 0.053550 0.053651 0.053752 0.053853 0.053953 0.054054	4.0422 4.0498 4.0574 4.0650 4.0726 4.0802 4.0878 4.0954 4.1030 4.1106



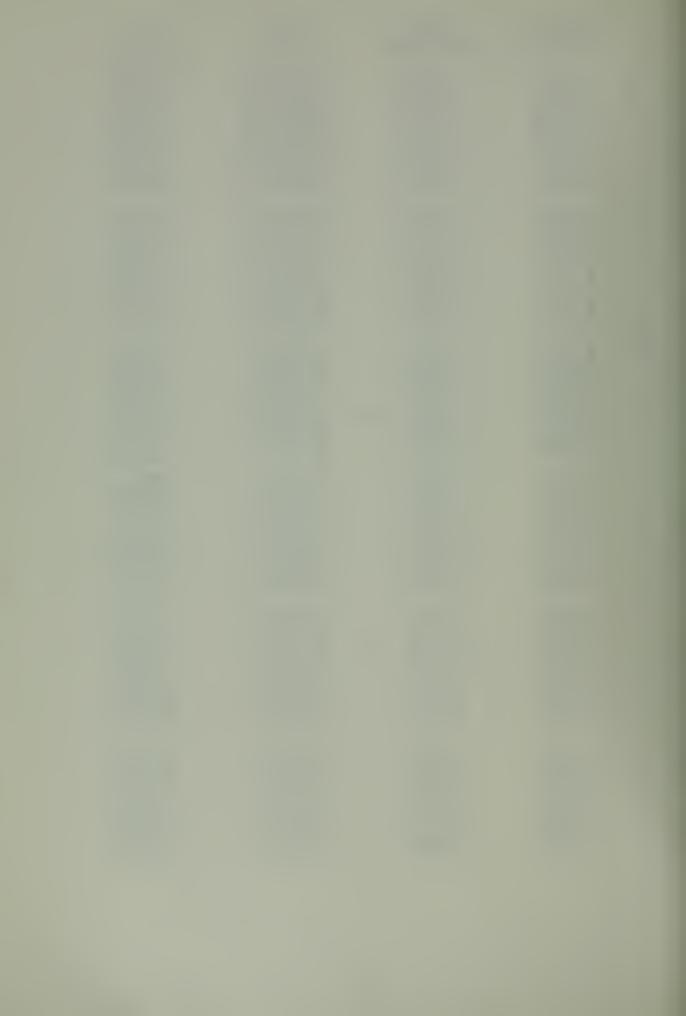
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
1.0800 1.0820 1.0840 1.0860 1.0880 1.0900 1.0920 1.0940 1.0960 1.0980	1.0258 1.0277 1.0296 1.0315 1.0334 1.0353 1.0372 1.0391 1.0410 1.0429	0.054155 0.054256 0.054357 0.054458 0.0544559 0.054660 0.054760 0.054861 0.054962 0.055063	4.1182 4.1258 4.1334 4.1410 4.1486 4.1562 4.1638 4.1714 4.1790 4.1866
1.1000 1.1020 1.1040 1.1060 1.1080 1.1100 1.1120 1.1140 1.1160 1.1180	1.0448 1.0467 1.0486 1.0505 1.0524 1.0543 1.0562 1.0581 1.0600 1.0619	0.055164 0.055265 0.055366 0.055467 0.055568 0.055668 0.055769 0.055870 0.055971	4.1942 4.2018 4.2094 4.2170 4.2246 4.2322 4.2398 4.2474 4.2550 4.2626
1.1200 1.1220 1.1240 1.1260 1.1280 1.1300 1.1320 1.1340 1.1360 1.1380	1.0638 1.0657 1.0676 1.0695 1.0714 1.0733 1.0752 1.0771 1.0790 1.0809	0.056173 0.056274 0.056375 0.056476 0.056577 0.056678 0.056779 0.056880 0.056981 0.057082	4.2702 4.2778 4.2854 4.2929 4.3005 4.3081 4.3157 4.3233 4.3309 4.3385
1 • 1 400 1 • 1 420 1 • 1 440 1 • 1 460 1 • 1 480 1 • 1 520 1 • 1 520 1 • 1 540 1 • 1 560 1 • 1 580	1.0828 1.0847 1.0866 1.0885 1.0904 1.0923 1.0942 1.0961 1.0980 1.0999	0.057183 0.057284 0.057385 0.057486 0.057586 0.057687 0.057788 0.0577889 0.057990 0.058091	4.3461 4.3537 4.3613 4.3689 4.3765 4.3841 4.3917 4.3993 4.4069 4.4145
1.1600 1.1620 1.1640 1.1660 1.1680 1.1700 1.1720 1.1740 1.1760 1.1780	1.1018 1.1037 1.1056 1.1075 1.1094 1.1113 1.1132 1.1151 1.1151	0.058192 0.058293 0.058293 0.058394 0.058495 0.058596 0.058697 0.058798 0.058899 0.059000 0.059101	4.4221 4.4297 4.4373 4.4448 4.4524 4.4600 4.4676 4.4752 4.4828 4.4904
1.1800 1.1820 1.1840 1.1860 1.1880 1.1900 1.1920 1.1940 1.1960 1.1980	1.1208 1.1227 1.1246 1.1265 1.1284 1.1303 1.1322 1.1341 1.1360 1.1379	0.059202 0.059303 0.059404 0.059506 0.059607 0.059708 0.059809 0.059810 0.060011	4.4980 4.5056 4.5132 4.5208 4.5284 4.5360 4.5436 4.5511 4.5587 4.5663



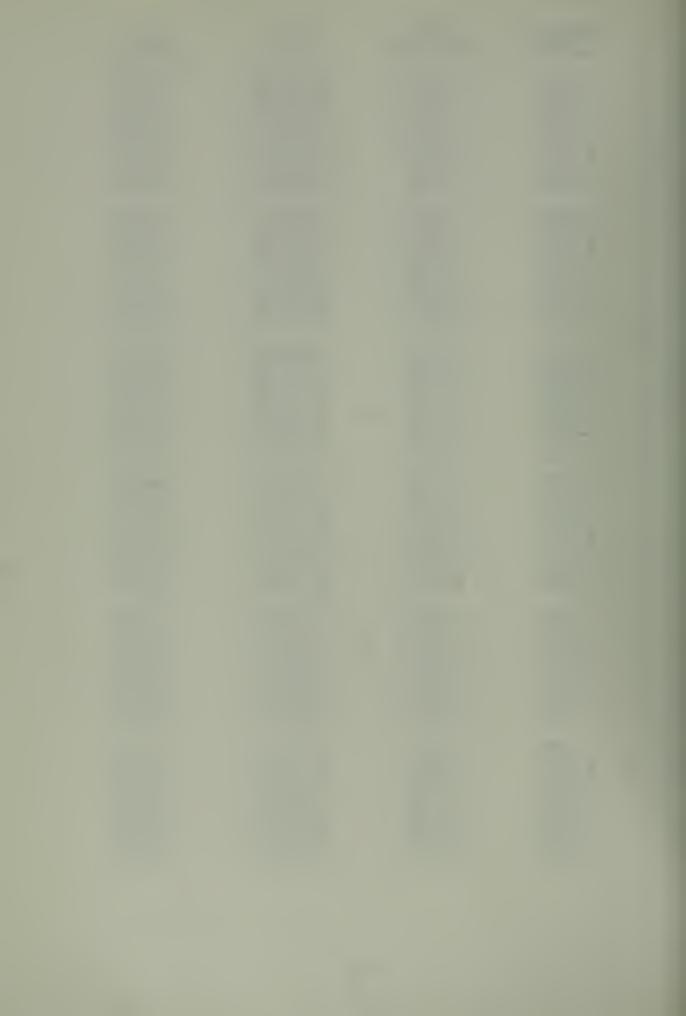
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
1. 2000 1. 2020 1. 2040 1. 2060 1. 2080 1. 2100 1. 2120 1. 2140 1. 2160 1. 2180	1.1398 1.1417 1.1436 1.1455 1.1474 1.1493 1.1512 1.1531 1.1550 1.1569	0.060213 0.060314 0.060415 0.060516 0.060617 0.060718 0.060819 0.060920 0.061021 0.061123	4.5739 4.5815 4.5891 4.5967 4.6043 4.6119 4.6270 4.6346 4.6422
1.2200 1.2220 1.2240 1.2260 1.2280 1.2300 1.2320 1.2340 1.2360 1.2380	1.1588 1.1607 1.1626 1.1645 1.1664 1.1683 1.1702 1.1721 1.1740 1.1759	0.061224 0.061325 0.061426 0.061527 0.061628 0.061729 0.061830 0.061931 0.062032 0.062134	4.6498 4.6574 4.6650 4.6726 4.6802 4.6878 4.6954 4.7029 4.7105 4.7181
1.2400 1.2420 1.2440 1.2460 1.2480 1.2520 1.2520 1.2520 1.2560	1.1778 1.1797 1.1816 1.1835 1.1854 1.1873 1.1892 1.1911 1.1930	0.062235 0.062336 0.062437 0.062538 0.062639 0.062740 0.062842 0.062943 0.063044 0.063145	4.7257 4.7333 4.7409 4.7485 4.7560 4.7636 4.7712 4.7788 4.7864 4.7940
1.2600 1.2620 1.2640 1.2660 1.2660 1.2700 1.2720 1.2740 1.2760 1.2780	1.1968 1.1987 1.2006 1.2024 1.2043 1.2062 1.2081 1.2100 1.2119 1.2138	0.063246 0.063348 0.063448 0.063550 0.063651 0.063752 0.063854 0.063955 0.064056	4.8016 4.8092 4.8167 4.8243 4.8319 4.8395 4.8471 4.8547 4.8622 4.8698
1.2800 1.2820 1.2840 1.2860 1.2860 1.2900 1.2920 1.2940 1.2960 1.2980	1.2157 1.2176 1.2195 1.2214 1.2233 1.2252 1.2271 1.2290 1.2309 1.2328	0.064258 0.064360 0.064461 0.064562 0.064663 0.064764 0.064865 0.064967 0.065068 0.065169	4.8774 4.8850 4.8926 4.9002 4.9078 4.9153 4.9229 4.9305 4.9381 4.9457
1.3000 1.3020 1.3040 1.3060 1.3100 1.3120 1.3140 1.3160 1.3180	1.2347 1.2366 1.2385 1.2404 1.2423 1.2442 1.2461 1.2480 1.2499 1.2518	0.065270 0.065372 0.065473 0.065574 0.065675 0.065777 0.065878 0.065979 0.066081 0.066182	4.9533 4.9608 4.9684 4.9760 4.9836 4.9912 4.9987 5.0063 5.0139 5.0215



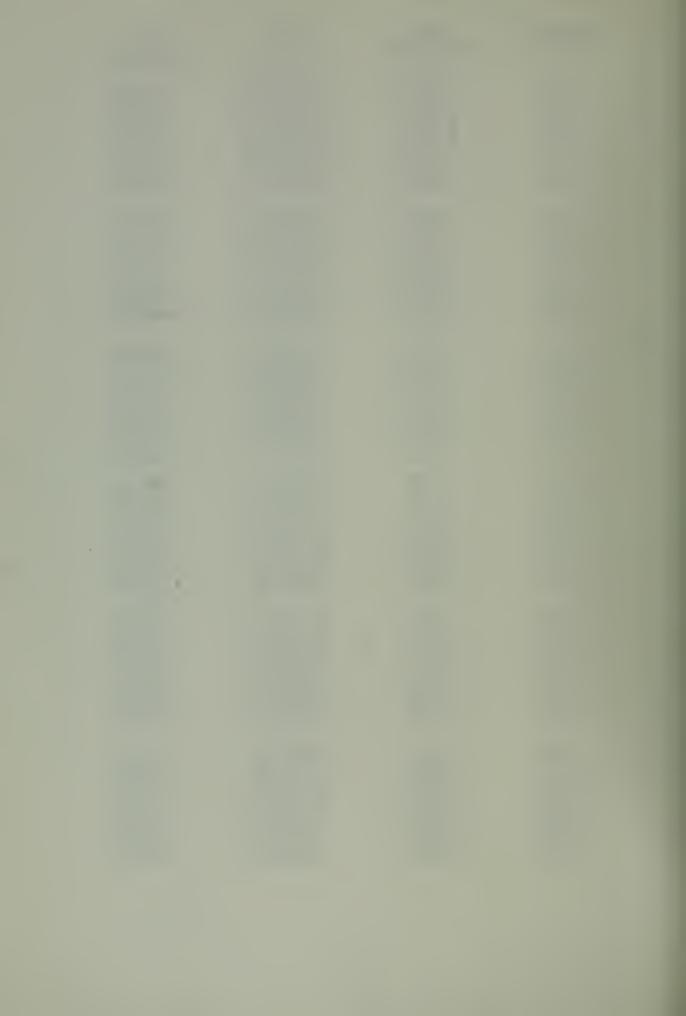
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
1.3200 1.3220 1.3240 1.3260 1.3280 1.3300 1.3320 1.3340 1.3360 1.3380	1.2537 1.2556 1.2575 1.2575 1.2613 1.2632 1.2651 1.2670 1.2689 1.2708	0.066283 0.066384 0.066486 0.066587 0.066688 0.066789 0.066891 0.066992 0.0667093 0.067195	5.0291 5.0366 5.0442 5.0518 5.0594 5.0670 5.0745 5.0897 5.0897
1.3400 1.3420 1.3440 1.3460 1.3500 1.3520 1.3540 1.3560 1.3580	1.2727 1.2746 1.2765 1.2784 1.2803 1.2822 1.2841 1.2860 1.2879 1.2898	0.067296 0.067398 0.067499 0.067600 0.067701 0.067803 0.067904 0.068C05 0.068107 0.068208	5.1049 5.1124 5.1200 5.1276 5.1352 5.1428 5.1579 5.1655 5.1731
1.3600 1.3620 1.3640 1.3660 1.3680 1.3700 1.3720 1.3740 1.3760 1.3780	1.2917 1.2936 1.2955 1.2974 1.2993 1.3012 1.3031 1.3050 1.3069 1.3088	0.068309 0.068411 0.068512 0.068614 0.068715 0.068816 0.068918 0.069019 0.069121 0.069222	5.1806 5.1882 5.1958 5.2010 5.2110 5.2185 5.2261 5.2337 5.2488
1.3800 1.3820 1.3840 1.3860 1.3880 1.3900 1.3920 1.3940 1.3960 1.3980	1.31.07 1.31.26 1.31.45 1.31.64 1.31.83 1.32.02 1.32.21 1.32.40 1.32.59 1.32.78	0.069323 0.069425 0.069526 0.069628 0.069729 0.069830 0.069932 0.070033 0.070135 0.070236	5.2564 5.2716 5.2716 5.2791 5.2867 5.2943 5.3019 5.3094 5.3170 5.3246
1.4000 1.4020 1.4040 1.4060 1.4080 1.4100 1.4120 1.4140 1.4160 1.4180	1.3297 1.3316 1.3335 1.3354 1.3373 1.3392 1.3411 1.3448 1.3467	0.070338 0.070439 0.070540 0.070642 0.070743 0.070845 0.070946 0.071048 0.071149 0.071251	5.3322 5.3327 5.3473 5.3549 5.3624 5.3700 5.3776 5.3852 5.3927 5.4003
1 • 4200 1 • 4220 1 • 4240 1 • 4260 1 • 4280 1 • 4300 1 • 4320 1 • 4340 1 • 4360 1 • 4380	1.3486 1.35505 1.3524 1.3543 1.3562 1.3581 1.3600 1.3619 1.3638 1.3657	0.071352 0.071454 0.071555 0.071657 0.071758 0.071860 0.071961 0.072063 0.072164 0.072266	5.4079 5.4154 5.4230 5.4306 5.4382 5.4457 5.4533 5.4609 5.4684 5.4760



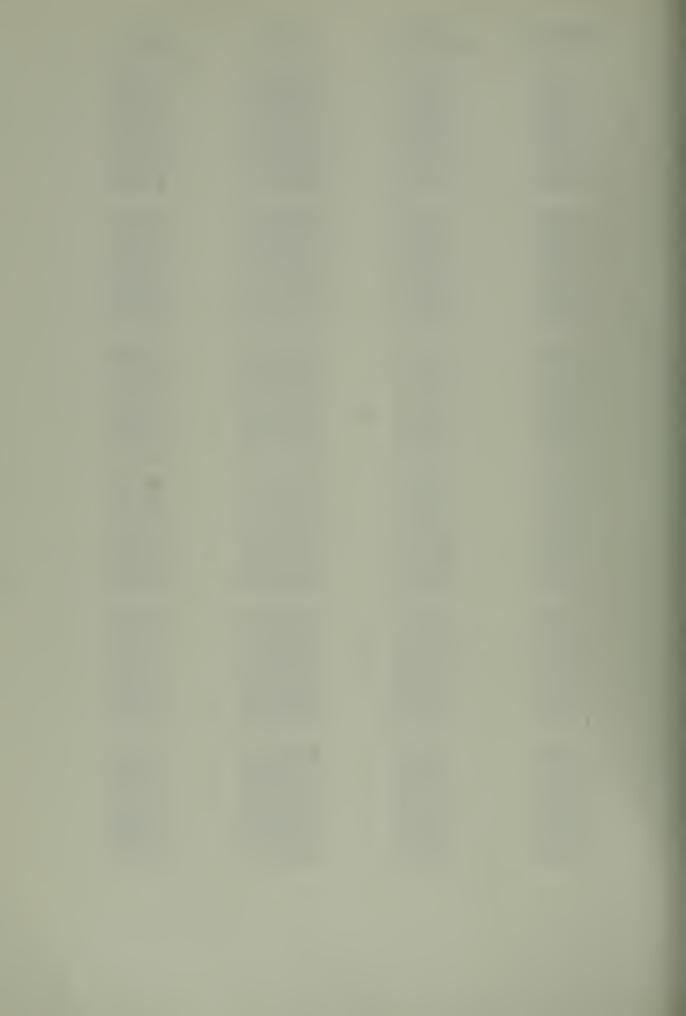
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
1.4400 1.4420 1.4440 1.4460 1.4480 1.4500 1.4520 1.4540 1.4560 1.4580	1.3676 1.3695 1.3714 1.3733 1.3752 1.3771 1.3790 1.3809 1.3828 1.3847	0.072367 0.072469 0.072571 0.072672 0.072774 0.072875 0.072977 0.073078 0.073180 0.073281	5.4836 5.4912 5.4987 5.5063 5.5139 5.5214 5.5290 5.5366 5.5441 5.5517
1.4600 1.4620 1.4640 1.4660 1.4680 1.4700 1.4720 1.4740 1.4760 1.4780	1.3866 1.3885 1.3904 1.3923 1.3942 1.3961 1.3980 1.3999 1.4018 1.4037	0.073383 0.073484 0.073586 0.073687 0.073789 0.073891 0.073992 0.074094 0.074196 0.074297	5.593 5.5668 5.5744 5.5820 5.5895 5.5971 5.6047 5.6128 5.6274
1.4800 1.4820 1.4840 1.4860 1.4880 1.4900 1.4920 1.4940 1.4960 1.4980	1.4056 1.4075 1.4075 1.413 1.4132 1.4151 1.4170 1.4189 1.4208 1.4227	0.074399 0.074500 0.074602 0.074704 0.074805 0.074907 0.075009 0.075110 0.075212 0.075314	5.6349 5.6425 5.6501 5.66576 5.6652 5.6879 5.6879 5.6879 5.7030
1.5000 1.5020 1.5040 1.5060 1.5080 1.5100 1.5120 1.5140 1.5160 1.5180	1. 42 46 1. 42 65 1. 42 84 1. 43 03 1. 43 22 1. 43 41 1. 43 60 1. 43 79 1. 43 98 1. 441 7	0.075415 0.075517 0.075618 0.075720 0.075822 0.075923 0.076025 0.076127 0.076229 0.076330	5.7106 5.7182 5.7257 5.7333 5.7408 5.7484 5.7560 5.7635 5.7711 5.7787
1.5200 1.5220 1.5240 1.5260 1.5280 1.53300 1.53320 1.53360 1.5380	1 • 4436 1 • 4455 1 • 4474 1 • 4493 1 • 4512 1 • 4531 1 • 4550 1 • 4569 1 • 4588 1 • 4607	0.076432 0.076534 0.076635 0.076737 0.076839 0.076940 0.077042 0.077144 0.077246 0.077347	5.7862 5.7938 5.8013 5.8089 5.8165 5.8240 5.8316 5.8391 5.8467 5.8543
1.5400 1.5420 1.5440 1.5460 1.5500 1.5520 1.55540 1.5560 1.5580	1.4625 1.4644 1.4663 1.4682 1.4701 1.4720 1.4739 1.4758 1.4777	0.077449 0.077551 0.077653 0.077754 0.077856 0.077958 0.078060 0.078161 0.078263 0.078365	5.8618 5.8694 5.8769 5.8845 5.8921 5.8996 5.9072 5.9147 5.9223 5.9299



OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
1 • 5600 1 • 5620 1 • 5640 1 • 5660 1 • 5680 1 • 5720 1 • 5720 1 • 5740 1 • 5760 1 • 5780	1 • 4815 1 • 4834 1 • 4853 1 • 4872 1 • 4891 1 • 4910 1 • 4929 1 • 4948 1 • 4967 1 • 4986	0.078467 0.078568 0.078670 0.078772 0.078874 0.078976 0.079078 0.079179 0.079281 0.079383	5.9374 5.9450 5.9525 5.9601 5.9676 5.9752 5.9827 5.9903 5.9979 6.0054
1.5800 1.5820 1.5840 1.5860 1.5880 1.5900 1.5920 1.5940 1.5960 1.5980	1.5005 1.5024 1.5043 1.5062 1.5081 1.5100 1.5119 1.5138 1.5157	0.079485 0.079587 0.079689 0.079790 0.079892 0.079994 0.080096 0.080198 0.080300 0.080402	6.0130 6.0205 6.0281 6.0356 6.0432 6.0508 6.0583 6.0659 6.0734 6.0810
1.6000 1.6020 1.6040 1.6060 1.6080 1.6100 1.6120 1.6140 1.6160	1.5195 1.5214 1.5233 1.5252 1.5271 1.5290 1.5309 1.5328 1.5347 1.5366	0.080503 0.080605 0.080707 0.080809 0.080911 0.081013 0.081115 0.081217 0.081319 0.081421	6.0885 6.0961 6.1036 6.1112 6.1187 6.1263 6.1338 6.1414 6.1489 6.1565
1.6200 1.6220 1.6240 1.6260 1.6280 1.6300 1.6320 1.6340 1.6360 1.6380	1.5385 1.5404 1.54423 1.5442 1.5461 1.5480 1.5499 1.5518 1.5537 1.5556	0.081523 0.081624 0.081726 0.081828 0.081930 0.082032 0.082134 0.082236 0.082338 0.082440	6.1640 6.1716 6.1791 6.1867 6.1942 6.2018 6.2093 6.2169 6.2244 6.2320
1.6400 1.6420 1.6440 1.6460 1.6480 1.6500 1.6520 1.6560 1.6560	1.5575 1.5594 1.5613 1.5632 1.5669 1.5669 1.5688 1.5707 1.5726	0.082542 0.082644 0.082746 0.082848 0.082950 0.083052 0.083154 0.083256 0.083358 0.083460	6.2395 6.2471 6.2546 6.2622 6.2697 6.2773 6.2848 6.2924 6.2999 6.3075
1.6600 1.6620 1.6640 1.6660 1.6680 1.6700 1.6720 1.6740 1.6760	1.5764 1.5783 1.5802 1.5821 1.5840 1.5859 1.5878 1.5897 1.5916	0.083562 0.083664 0.083766 0.083868 0.083970 0.084072 0.084174 0.084276 0.084378 0.084480	6.3150 6.3226 6.3301 6.3377 6.3452 6.3528 6.3603 6.3679 6.3754 6.3829



OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
1.6800 1.6820 1.6840 1.6860 1.6880 1.6900 1.6920 1.6940 1.6960 1.6980	1.5954 1.5973 1.5992 1.6011 1.6030 1.6049 1.6068 1.6(87 1.6106 1.6125	0.084583 0.084685 0.084787 0.084787 0.084889 0.084991 0.085093 0.085195 0.085399 0.085399	6.3905 6.3980 6.4056 6.4131 6.4207 6.4282 6.4358 6.4433 6.4508
1.7000 1.7020 1.7040 1.7060 1.7080 1.7100 1.7120 1.7140 1.7160 1.7180	1.6144 1.6163 1.6182 1.6201 1.6220 1.6239 1.6258 1.6277 1.6296 1.6315	0.085603 0.085706 0.085808 0.085910 0.086012 0.086114 0.086216 0.086318 0.086421 0.086523	6.4659 6.4735 6.4810 6.4886 6.4961 6.5036 6.5112 6.5187 6.5263 6.5338
1.7200 1.7220 1.7240 1.7260 1.7280 1.7300 1.7320 1.7340 1.7360 1.7380	1.6334 1.6353 1.6372 1.6391 1.6410 1.6429 1.6448 1.6467 1.6486	0.086625 0.086727 0.086829 0.086932 0.087034 0.087136 0.087238 0.087340 0.087443	6.5413 6.5489 6.5564 6.5640 6.5715 6.5790 6.5866 6.5941 6.6092
1.7400 1.7420 1.7440 1.7460 1.7480 1.7500 1.7520 1.7540 1.7560 1.7580	1.6524 1.6543 1.6561 1.6599 1.6618 1.6637 1.6656 1.6675	0.087647 0.087749 0.087851 0.087954 0.088056 0.088158 0.088260 0.088363 0.088465 0.088567	6.6167 6.6243 6.6318 6.6393 6.6469 6.6544 6.6620 6.6695 6.6770 6.6846
1.7600 1.7620 1.7640 1.7660 1.7680 1.7700 1.7720 1.7740 1.7760 1.7780	1.6713 1.6732 1.6751 1.6770 1.6789 1.6808 1.6827 1.6846 1.6865	0.088669 0.088772 0.088874 0.088976 0.089079 0.089181 0.089283 0.089385 0.089488 0.089590	6.6921 6.6996 6.7072 6.7147 6.7222 6.7298 6.7373 6.7448 6.7524 6.7599
1.7800 1.7820 1.7840 1.7860 1.7880 1.7900 1.7920 1.7940 1.7960 1.7980	1.6903 1.6922 1.6941 1.6960 1.6979 1.6998 1.7017 1.7036 1.7055	0.089692 0.089795 0.089897 0.089999 0.090102 0.090204 0.090306 0.090409 0.090511 0.090613	6.7674 6.7750 6.7825 6.7900 6.7976 6.8051 6.8126 6.8202 6.8277 6.8352



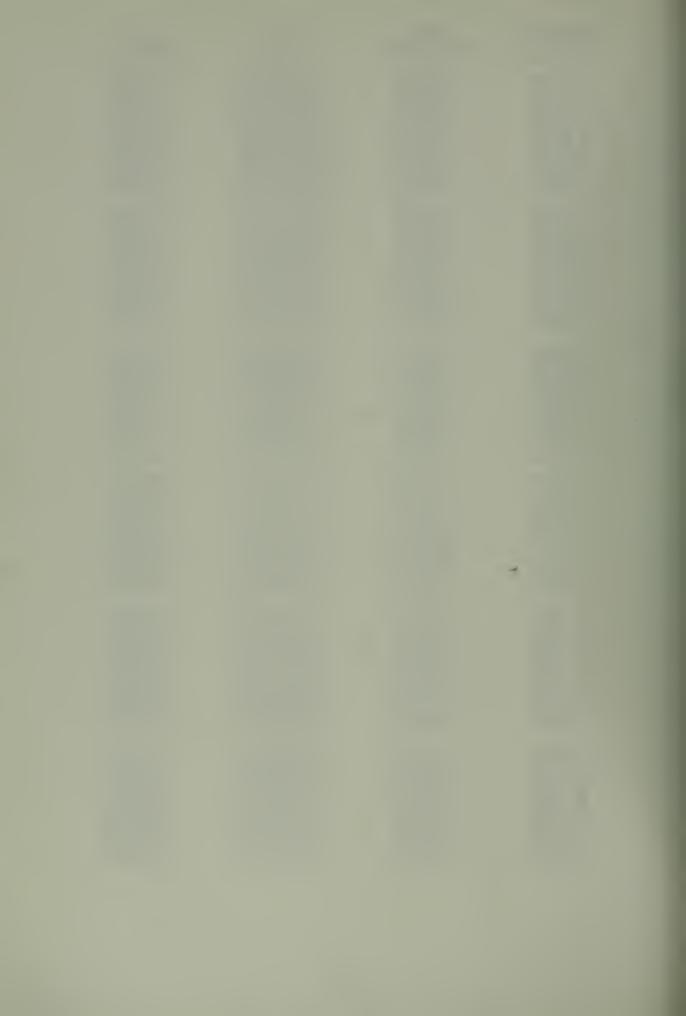
000550450	TOUE	50000	D.4.V
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
1.8000 1.8020 1.8040 1.8060 1.8780 1.8120 1.8120 1.8140 1.8160 1.8180	1.7093 1.7112 1.7131 1.7150 1.7169 1.7188 1.7207 1.7226 1.7226	0.090716 0.090818 0.090921 0.091023 0.091125 0.091228 0.091330 0.091433 0.091535 0.091637	6.8428 6.8503 6.8578 6.8654 6.86729 6.8804 6.8879 6.8955 6.9030 6.9105
1.8200 1.8220 1.8240 1.8260 1.8280 1.8300 1.8320 1.8340 1.8360 1.8380	1.7283 1.7302 1.7321 1.7340 1.7358 1.7377 1.7396 1.7415 1.7434	0.091740 0.091842 0.091945 0.092047 0.092150 0.092252 0.092355 0.092457 0.092559 0.092662	6.9181 6.9256 6.9331 6.9406 6.9482 6.9557 6.9632 6.9708 6.9783 6.9858
1.8400 1.8420 1.8440 1.8460 1.8480 1.8500 1.8520 1.8540 1.8560 1.8580	1.7472 1.7491 1.7510 1.7529 1.7548 1.7567 1.7586 1.7605 1.7624	0.092764 0.092867 0.092969 0.093072 0.093174 0.093277 0.093379 0.093482 0.093584 0.093687	6.9933 7.0009 7.0084 7.0159 7.0234 7.0310 7.0385 7.0460 7.0535 7.0611
1.8600 1.8620 1.8640 1.8660 1.8680 1.8700 1.8720 1.8740 1.8760 1.8760	1.7662 1.7681 1.7700 1.7719 1.7738 1.7757 1.7776 1.7776	0.093790 0.093892 0.093995 0.094097 0.094200 0.094302 0.094405 0.094507 0.094610 0.094713	7.0686 7.0761 7.0836 7.0912 7.0987 7.1062 7.1137 7.1213 7.1288 7.1363
1.8800 1.8820 1.8840 1.8860 1.8880 1.8900 1.8920 1.8940 1.8960 1.8980	1.7852 1.7871 1.7890 1.7909 1.7928 1.7947 1.7966 1.7985 1.8004 1.8023	0.094815 0.094918 0.095020 0.095123 0.095226 0.095328 0.095431 0.095533 0.095636 0.095739	7.1438 7.1513 7.1589 7.1664 7.1739 7.1814 7.1889 7.1965 7.2040 7.2115
1.9000 1.9020 1.9040 1.9060 1.9080 1.9100 1.9120 1.9140 1.9160 1.9180	1.8042 1.8061 1.8080 1.8098 1.8117 1.8136 1.8155 1.8174 1.8193 1.8212	0.095841 0.095944 0.096047 0.096149 0.096252 0.096355 0.096457 0.096560 0.096663	7 • 2190 7 • 2265 7 • 2340 7 • 2416 7 • 2491 7 • 2566 7 • 2641 7 • 2716 7 • 2792 7 • 2867



OBSERVED	TRUE	ERROR	RAY
DISTANCE (INCHES)	DISTANCE (INCHES)	(INCHES)	ANGLE (DEGREES)
1.9200 1.9220 1.9240 1.9260 1.9280 1.9300 1.9320 1.9340 1.9360 1.9380	1.8231 1.8250 1.8269 1.8288 1.8307 1.8326 1.8345 1.8364 1.8383 1.8402	0.096868 0.096971 0.097073 0.097176 0.097279 0.097382 0.097484 0.097587 0.097690 0.097793	7.2942 7.3017 7.3092 7.3167 7.3242 7.3318 7.3393 7.3468 7.3543 7.3618
1.9400 1.9420 1.9440 1.9460 1.9480 1.9500 1.9520 1.9540 1.9560 1.9580	1.8421 1.8440 1.8459 1.8478 1.8497 1.8516 1.8535 1.8554 1.8573	0.097895 0.097998 0.098101 0.098204 0.098306 0.098409 0.098512 0.098615 0.098718 0.098820	7.3693 7.3769 7.3844 7.3919 7.3994 7.4069 7.4144 7.4219 7.4294 7.4370
1.9600 1.9620 1.9640 1.9660 1.9680 1.9700 1.9720 1.9740 1.9760 1.9780	1.8611 1.8630 1.8649 1.8668 1.8687 1.8706 1.8725 1.8744 1.8763 1.8782	0.098923 0.099026 0.099129 0.099232 0.099335 0.099437 0.099540 0.099643 0.099746	7.4445 7.4520 7.4595 7.4670 7.4745 7.4820 7.4895 7.4970 7.5045 7.5121
1.9800 1.9820 1.9840 1.9860 1.9880 1.9900 1.9920 1.9940 1.9960 1.9980	1.8800 1.8819 1.8838 1.8857 1.8876 1.8895 1.8914 1.8933 1.8952 1.8971	0.099952 0.100054 0.100157 0.100260 0.100363 0.100466 0.100569 0.100672 0.100775 0.100878	7.5196 7.5271 7.5346 7.5346 7.5421 7.5571 7.5571 7.5646 7.5721 7.5796 7.5871
2.0000 2.0020 2.0040 2.0060 2.0080 2.0100 2.0120 2.0140 2.0160 2.0180	1.8990 1.9009 1.9028 1.9047 1.9066 1.9085 1.9104 1.9123 1.9142	0.100981 0.101084 0.101186 0.101289 0.101392 0.101495 0.101598 0.101701 0.101804 0.101907	7.5946 7.6021 7.6096 7.6172 7.6247 7.6322 7.6397 7.6472 7.6547 7.6622
2.0200 2.0220 2.0240 2.0260 2.0280 2.0300 2.0320 2.0340 2.0360 2.0380	1.9180 1.9199 1.9218 1.9237 1.9256 1.9275 1.9294 1.9313 1.9332	0.102010 0.102113 0.102216 0.102319 0.102422 0.102525 0.102628 0.102731 0.102834 0.102937	7.6697 7.6772 7.6847 7.6922 7.6997 7.7072 7.7147 7.7222 7.7297 7.7372



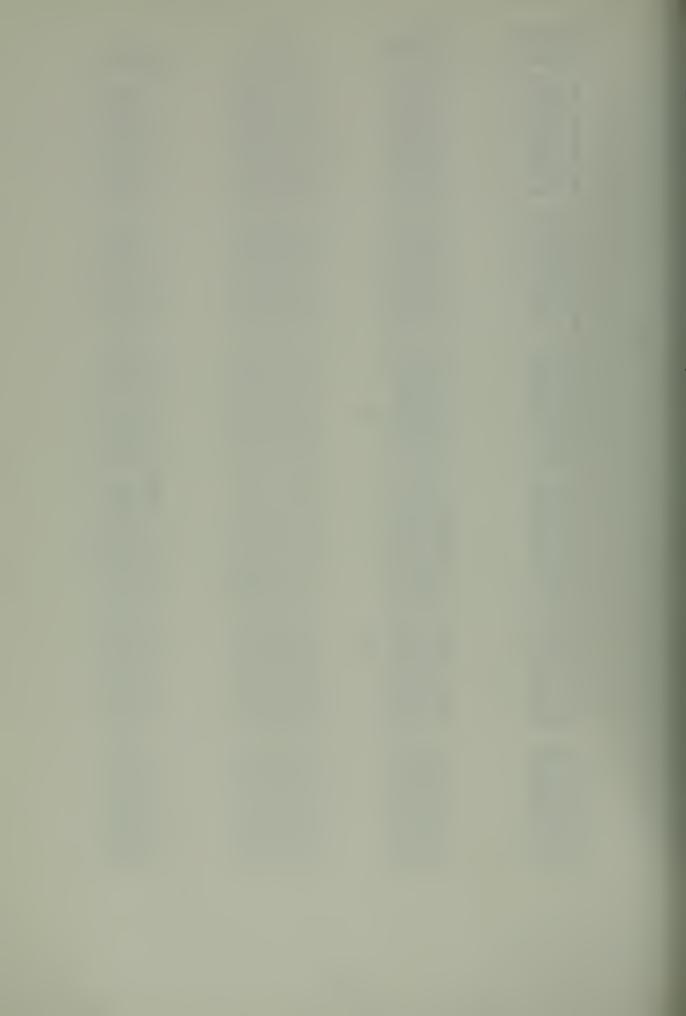
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
2.0400 2.0420 2.0440 2.0460 2.0480 2.0500 2.0520 2.0520 2.0560 2.0560 2.0580	1.9370 1.9389 1.9408 1.9426 1.9445 1.9464 1.9483 1.9502 1.9521	0.103040 0.103143 0.103246 0.103349 0.103453 0.103556 0.103659 0.103762 0.103865 0.103968	7.7447 7.7522 7.7597 7.7672 7.7747 7.7822 7.7897 7.7972 7.8047 7.8122
2.0600 2.0620 2.0640 2.0660 2.0680 2.0700 2.0720 2.0740 2.0760 2.0780	1.9559 1.9578 1.9597 1.9616 1.9635 1.9654 1.9673 1.9692 1.9711	0.104071 0.104174 0.104277 0.104380 0.104483 0.104587 0.104690 0.104793 0.104896 0.104999	7.8197 7.8272 7.8347 7.8422 7.8497 7.8572 7.8647 7.8722 7.8797 7.88797
2.0800 2.0820 2.0840 2.0860 2.0880 2.0900 2.0920 2.0940 2.0960 2.0980	1.9749 1.9768 1.9787 1.9806 1.9825 1.9844 1.9863 1.9882 1.9901	0.105102 0.105205 0.105309 0.105412 0.105515 0.105618 0.105721 0.105825 0.105928 0.106031	7.8947 7.9022 7.9097 7.9172 7.9246 7.9321 7.9396 7.9471 7.9546 7.9621
2.1000 2.1020 2.1040 2.1060 2.1080 2.1100 2.1120 2.1140 2.1160 2.1180	1.9939 1.9958 1.9977 1.9996 2.0015 2.0033 2.0052 2.0071 2.0090 2.0109	0.106134 0.106238 0.106341 0.106444 0.106547 0.106650 0.106754 0.106857 0.106960 0.107064	7.9696 7.9771 7.9846 7.9921 7.9996 8.0071 8.0146 8.0220 8.0295 8.0370
2. 1200 2. 1220 2. 1240 2. 1260 2. 1280 2. 1300 2. 1320 2. 1340 2. 1360 2. 1380	2.0128 2.0147 2.0166 2.0185 2.0204 2.0223 2.0242 2.0261 2.0280 2.0299	0.107167 0.107270 0.107374 0.107477 0.107580 0.107683 0.107787 0.107890 0.107993 0.108097	8.0445 8.0520 8.0595 8.0670 8.0745 8.0820 8.0894 8.0969 8.1044 8.1119
2. 1400 2. 1420 2. 1440 2. 1460 2. 1480 2. 1500 2. 1520 2. 1540 2. 1560 2. 1580	2.0318 2.0337 2.0356 2.0375 2.0394 2.0413 2.0432 2.0451 2.0470 2.0489	0.108200 0.108303 0.108407 0.108510 0.108613 0.108717 0.108820 0.108924 0.109027 0.109131	8.1194 8.1269 8.1344 8.1419 8.1493 8.1568 8.1643 8.1718 8.1793 8.1868



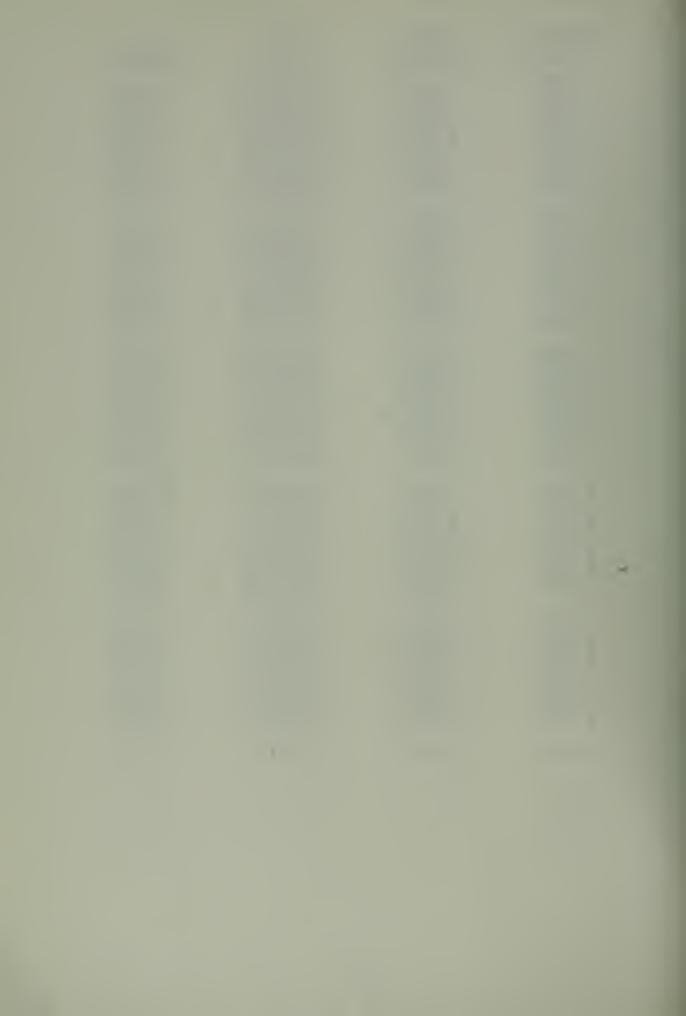
OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
2. 1600	2.0508	0.109234	8 • 1943
2. 1620	2.0527	0.109337	8 • 2017
2. 1640	2.0546	0.109441	8 • 2092
2. 1660	2.0565	0.109544	8 • 2167
2. 1680	2.0584	0.109647	8 • 2242
2. 1700	2.0602	0.109751	8 • 2317
2. 1720	2.0621	0.109854	8 • 2392
2. 1740	2.0640	0.109958	8 • 2466
2. 1760	2.0659	0.110061	8 • 2541
2. 1780	2.0678	0.110165	8 • 2616
2. 1800	2.0697	0.110268	8.2691
2. 1820	2.0716	0.110372	8.2766
2. 1840	2.0735	0.110475	8.2840
2. 1860	2.0754	0.110579	8.2915
2. 1880	2.0773	0.110682	8.2990
2. 1900	2.0792	0.110786	8.3065
2. 1920	2.0811	0.110889	8.3140
2. 1940	2.0830	0.110993	8.3214
2. 1960	2.0849	0.111096	8.3289
2. 1980	2.0868	0.111200	8.3364
2.2000	2.0887	0 • 11 13 03	8.3439
2.2020	2.0906	0 • 11 14 07	8.3514
2.2040	2.0925	0 • 11 15 10	8.3588
2.2060	2.0944	0 • 11 16 14	8.3663
2.2080	2.0963	0 • 11 17 17	8.3738
2.2100	2.0982	0 • 11 18 21	8.3813
2.2120	2.1001	0 • 11 19 24	8.3887
2.2140	2.1020	0 • 11 20 28	8.3962
2.2160	2.1039	0 • 11 21 31	8.4037
2.2180	2.1058	0 • 11 22 35	8.4112
2.2200	2.1077	0.112339	8 • 4187
2.2220	2.1096	0.112442	8 • 4261
2.2240	2.1115	0.112546	8 • 4336
2.2260	2.1133	0.112650	8 • 4411
2.2280	2.1152	0.112753	8 • 4486
2.2300	2.1171	0.112857	8 • 4560
2.2320	2.1190	0.112960	8 • 4635
2.2340	2.1209	0.113064	8 • 4710
2.2360	2.1228	0.113168	8 • 4784
2.2380	2.1247	0.113271	8 • 4859
2.2400 2.2420 2.2440 2.2460 2.2480 2.2500 2.2520 2.2540 2.2560 2.2580	2.1266 2.1285 2.1304 2.1323 2.1342 2.1361 2.1380 2.1399 2.1418 2.1437	0.113375 0.113479 0.113582 0.113686 0.113790 0.113893 0.113997 0.114101 0.114205 0.114308	8 • 4934 8 • 50083 8 • 55158 8 • 55338 8 • 53382 8 • 55457 8 • 55606
2.2600 2.2620 2.2640 2.2660 2.2680 2.2700 2.2720 2.2720 2.2740 2.2760 2.2780	2.1456 2.1475 2.1494 2.1513 2.1532 2.1551 2.1570 2.1589 2.1608 2.1627	0.114412 0.114516 0.114620 0.114723 0.114827 0.114931 0.115034 0.115138 0.115242 0.115346	8.5681 8.5756 8.5830 8.5995 8.5980 8.6055 8.6129 8.6204 8.6279



OBSERVED	TRUE_	ERROR	RAY
OBSERVED DISTANCE (INCHES)	DIŜTĂÑCE (INCHES)	(INCHES)	ANGLE (DEGREES)
2. 2800 2. 2820 2. 2840 2. 2860 2. 2880 2. 2900 2. 2920 2. 2940 2. 2960 2. 2980	2.1645 2.1664 2.1683 2.1702 2.1721 2.1740 2.1759 2.1778 2.1797 2.1816	0.115450 0.115553 0.115657 0.115761 0.115865 0.115969 0.116072 0.116176 0.116280 0.116384	8 • 6428 8 • 6503 8 • 6577 8 • 6652 8 • 6727 8 • 6801 8 • 6876 8 • 6951 8 • 7025 8 • 7100
2.3000 2.3020 2.3040 2.3060 2.3100 2.3120 2.3140 2.3160 2.3180	2.1835 2.1854 2.1873 2.1892 2.1911 2.1930 2.1949 2.1968 2.1987 2.2006	0.116488 0.116592 0.116695 0.116799 0.116903 0.117007 0.117111 0.117215 0.117319 0.117423	8.7174 8.7249 8.7324 8.7398 8.7473 8.7548 8.7622 8.7697 8.7772
2.3200 2.3220 2.3240 2.3260 2.3280 2.3300 2.3320 2.3340 2.3360 2.3380	2.2025 2.2044 2.2063 2.2082 2.2101 2.2120 2.2138 2.2157 2.2176 2.2195	0.117527 0.117631 0.117735 0.117839 0.117942 0.118046 0.118150 0.118254 0.118358 0.118462	8.7921 8.7995 8.8070 8.8145 8.8219 8.8294 8.8368 8.8443 8.8518
2.3400 2.3440 2.3460 2.3480 2.3500 2.3520 2.3540 2.3560 2.3580	2 · 2214 2 · 2233 2 · 2252 2 · 2271 2 · 2290 2 · 2309 2 · 2328 2 · 2347 2 · 2366 2 · 2385	0.118566 0.118670 0.118774 0.118878 0.118982 0.119086 0.119190 0.119294 0.119398 0.119502	8.8667 8.8741 8.8816 8.8890 8.8965 8.9040 8.9114 8.9189 8.9263 8.9338
2.3600 2.3620 2.3640 2.3660 2.3680 2.3700 2.3720 2.3740 2.3760 2.3780	2 · 2404 2 · 2423 2 · 2442 2 · 2461 2 · 2480 2 · 2499 2 · 2518 2 · 2537 2 · 2556 2 · 2575	0.119606 0.119711 0.119815 0.119919 0.120023 0.120127 0.120231 0.120335 0.120439 0.120543	8 • 9412 8 • 9487 8 • 9561 8 • 9636 8 • 9711 8 • 9785 8 • 9860 8 • 9934 9 • 0009 9 • 0083
2.3800 2.3820 2.3840 2.3860 2.3880 2.3900 2.3920 2.3940 2.3960 2.3980	2.2594 2.2612 2.2631 2.2650 2.2669 2.2688 2.2707 2.2726 2.2745 2.2764	0.120647 0.120751 0.120856 0.120960 0.121064 0.121168 0.121272 0.121376 0.121480 0.121585	9.0158 9.0232 9.0307 9.0381 9.0456 9.0530 9.0605 9.0679 9.0754 9.0828



OBSERVED DISTANCE (INCHES)	TRUE DISTANCE (INCHES)	ERROR (INCHES)	RAY ANGLE (DEGREES)
2.4000 2.4020 2.4040 2.4060 2.4080 2.4100 2.4120 2.4140 2.4160 2.4180	2.2783 2.2802 2.2821 2.2840 2.2859 2.2878 2.2897 2.2916 2.2935 2.2954	0.121689 0.121793 0.121897 0.122001 0.122106 0.122210 0.122314 0.122418 0.122523 0.122627	9.0903 9.0977 9.1052 9.1126 9.1201 9.1275 9.1350 9.1424 9.1498 9.1573
2.4200 2.4220 2.4240 2.4260 2.4280 2.4300 2.4320 2.4340 2.4360 2.4380	2.2973 2.2992 2.3011 2.3030 2.3049 2.3067 2.3086 2.3105 2.3124 2.3143	0 • 122731 0 • 122835 0 • 122940 0 • 123044 0 • 123148 0 • 123253 0 • 123357 0 • 123461 0 • 123565 0 • 123670	9.1647 9.1722 9.1796 9.1871 9.1945 9.2020 9.2094 9.2168 9.2243 9.2317
2.4400 2.4420 2.4440 2.4460 2.4500 2.4520 2.4520 2.4560 2.4580	2.3162 2.3181 2.3200 2.3219 2.3238 2.3257 2.3276 2.3295 2.3314 2.3333	0.123774 0.123879 0.123983 0.124087 0.124191 0.124296 0.124400 0.124505 0.124609 0.124713	9.2392 9.2466 9.2541 9.2615 9.2689 9.2764 9.2838 9.2913 9.2987 9.3061
2.4600 2.4620 2.4640 2.4660 2.4680 2.4700 2.4720 2.4740 2.4760 2.4780	2.3352 2.3371 2.3390 2.3409 2.3428 2.3447 2.3466 2.3484 2.3503 2.3522	0 • 124818 0 • 124922 0 • 125027 0 • 125131 0 • 125235 0 • 125340 0 • 125444 0 • 125549 0 • 125653 0 • 125758	9.3136 9.3210 9.3285 9.3359 9.3433 9.3508 9.3582 9.3657 9.3731 9.3805
2. 4800 2. 4820 2. 4840 2. 4860 2. 4880 2. 4900 2. 4920 2. 4940 2. 4960 2. 4980	2.3541 2.3560 2.3579 2.3598 2.3617 2.3636 2.3655 2.3674 2.3693 2.3712	0.125862 0.125967 0.126071 0.126176 0.126280 0.126385 0.126489 0.126594 0.126698 0.126803	9.3880 9.3954 9.4028 9.4103 9.4177 9.4251 9.4326 9.4400 9.4474
2.5000	2.3731	0.126907	9.4623



#### APPENDIX C

# APPLICATION OF COMPUTER PROGRAM "HOLOFER"

The computer program is an adptation of the inversion first proposed by C. D. Maldonado [9, 10, 11] and is designed to invert fringe numbers across a field to the density field. It can be operated in three different modes as described below:

# (a) Mode 1

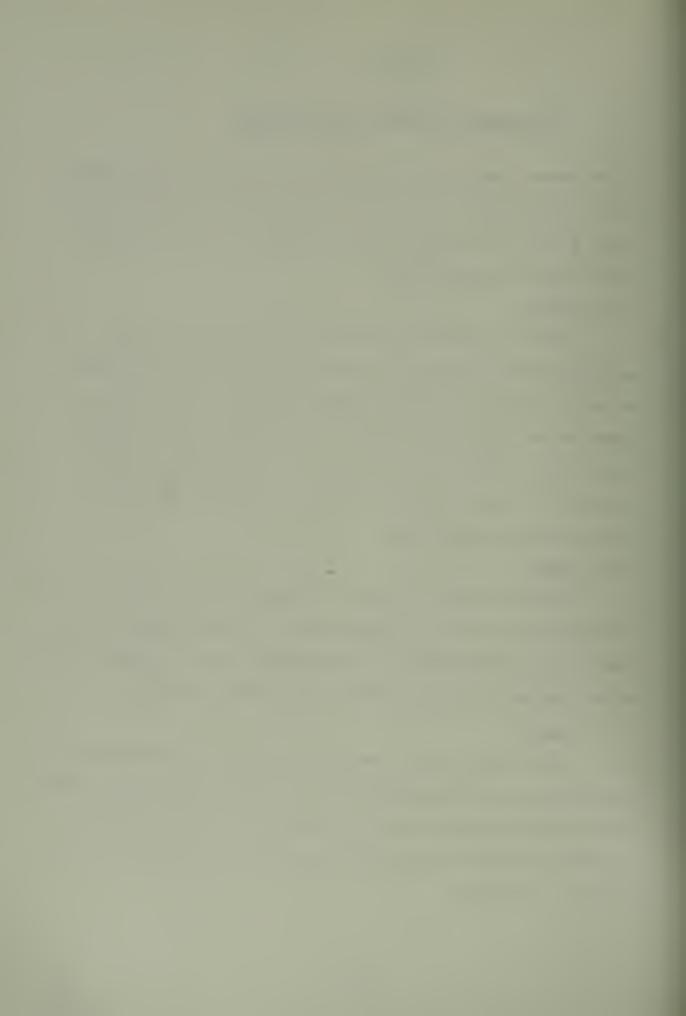
Mode 1 is utilized as a self-test of the computer program. It can either generate its own input density field using Subroutine FUNCT or read in a density field through Subroutine FREAD. The program then generates the fringe array and inverts the array back to the original density field. This mode was utilized in the present investigation to determine the value of the scale factor, < , required to obtain the correct density across the fin.

### (b) Mode 2

This mode reads in irregularly spaced fringe data and generates the fringe array at regular intervals across the field using Subroutine SHEET. By specifying NCODE = 1, the fringe array can be generated by one of the functions in Subroutine FUNCT. Mode 2 was not utilized.

## (c) Mode 3

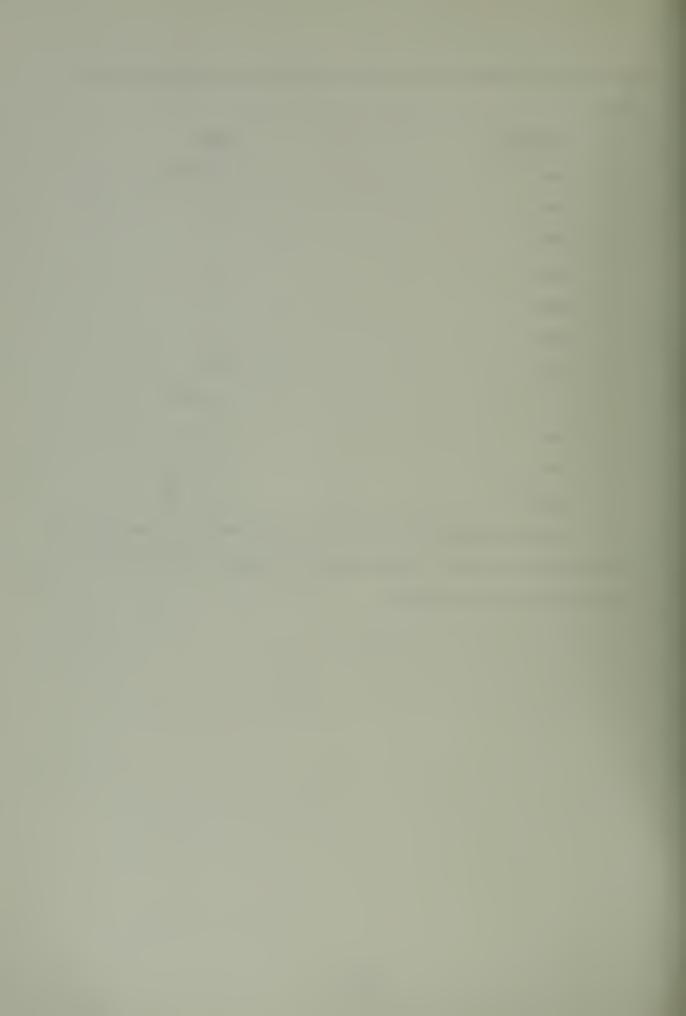
Mode 3 reads in the fringe data at regularly spaced intervals and inverts the array to density data across the field. The Subroutine GARRAY calls Subroutine READ to read in the fringe data. The first two cards preceding the fringe data provide the program with the fringe field size, location, and symmetry.



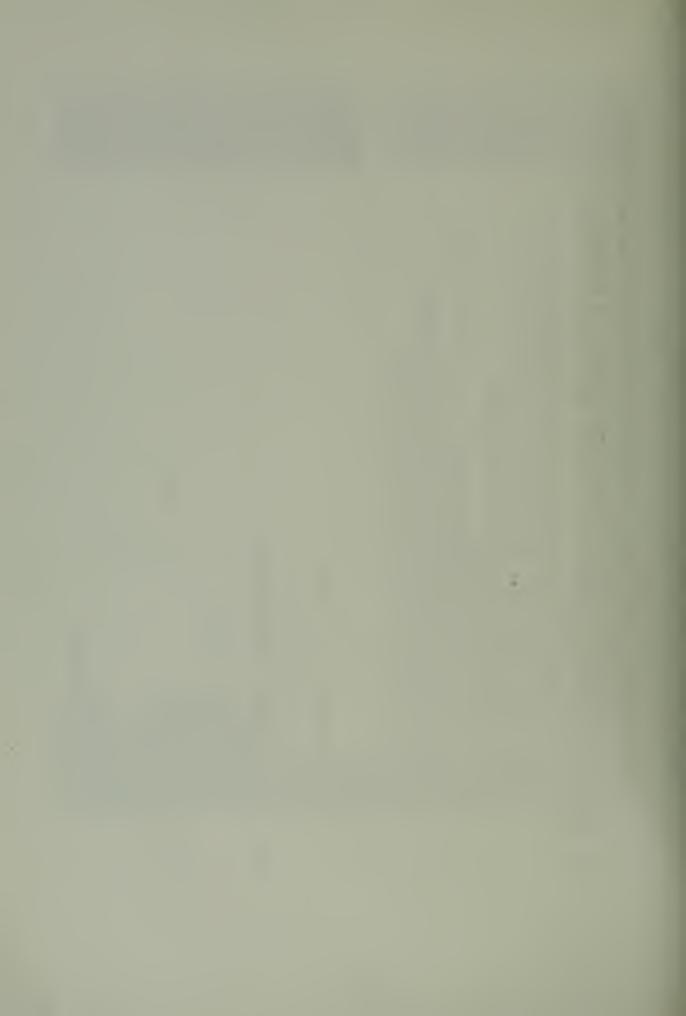
The following parameters were used in considering the symmetric field case:

PARAMETER	INPUT
NOF	Run Number
IMAX	201
JMAX	1
ISYM	101
JSYM	· 1
IMS .	2
JMS	100
Z	0.387
хо	0.0
YO	0.0
PHISYM	0.0

References [3] and [12] contain further details and applications of the computer program. A print-out of the program is included in the next few pages of this appendix.

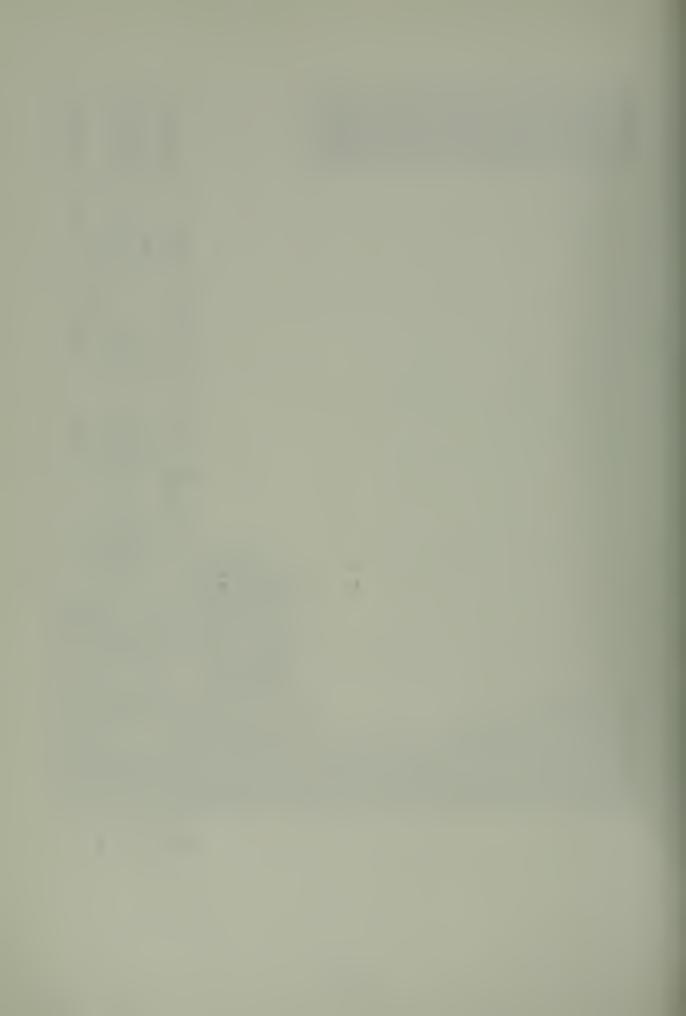


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IN1=5
IN2=5
IN4=5
IN4=5
IN 1=1
IN 4=6
IN 1=1
IN 4=6
IN NO = 1
IN AX = AR (1)
IN AX = AR (2)
IN AX = AR (4)
IN IN IT = AR (4)
IN IT = AR (1)
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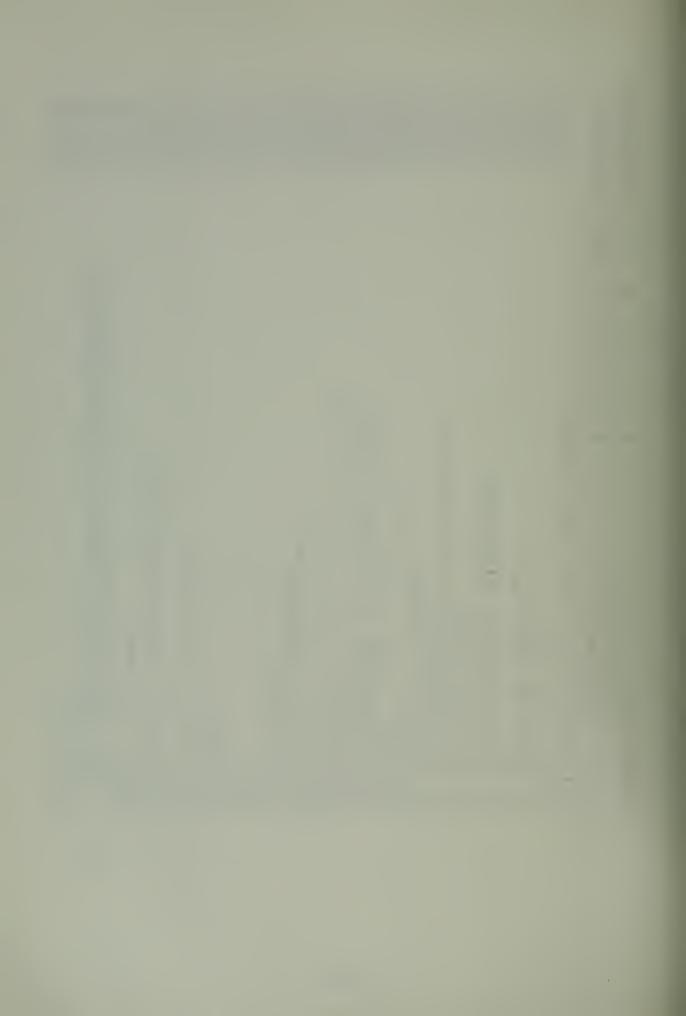


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ET00340
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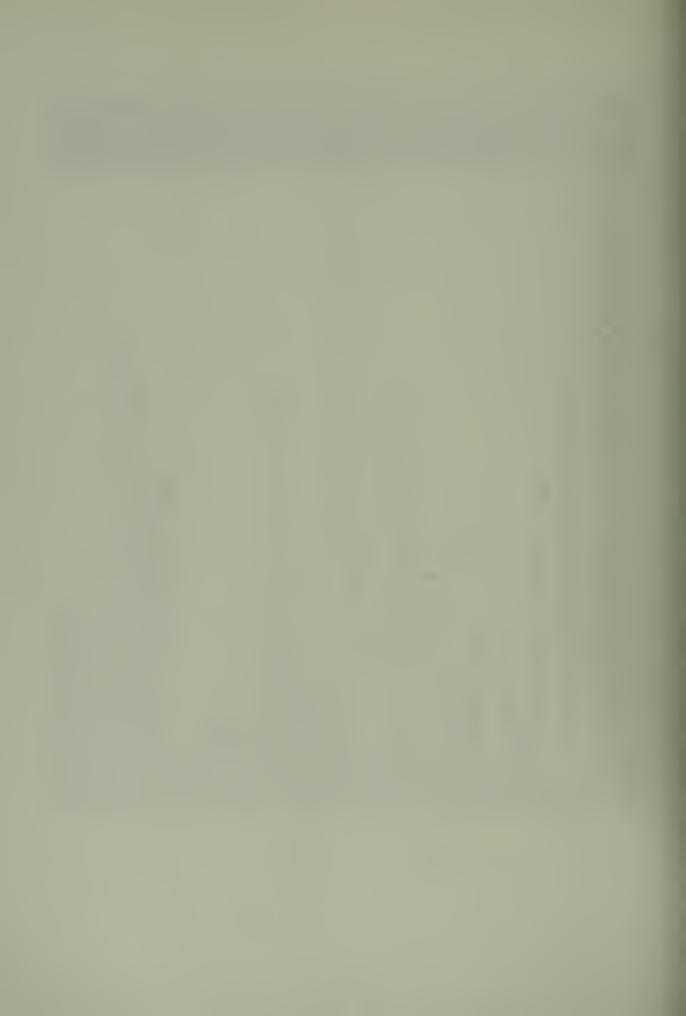
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TS.LE.20).AND.(I.NE.NPTS)) WRITE (6,79)
                                                                                                                                                                OR. (TTL.EQ.1.).OR. (I.EQ.NPTS)) TL(L)=PL
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L(I)=PL
                        = (CALC(I, J)-THEO(I, J))
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 )=RHOINF*(CALC(I, J)+1.)

=CA(I)

HODE.EQ.1) ERR(I)=(CALC(I)DE.GT.1) THEO(I, J)=FA(I)

PT.EQ.0) GO TO 3
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+1.)*20.+2.5
RHO(I) = RHOINF* (CALC(I, J))

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IF (MODE • EQ•1) THEO(I; J)

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IF (IC• EQ•0) TL(I) = PL

IF (IC• EQ•10) LC=0

IF (IC• EQ•10) LC=0

IF (LC• EQ•10) LC=0

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IF (LW• GT•62) LW=62

IF (LW• CA(I)+1.)*20.+2.5

IF (LW• CA(I)+1.)*20.+2.5

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IF (LV• GT•62) LY=62

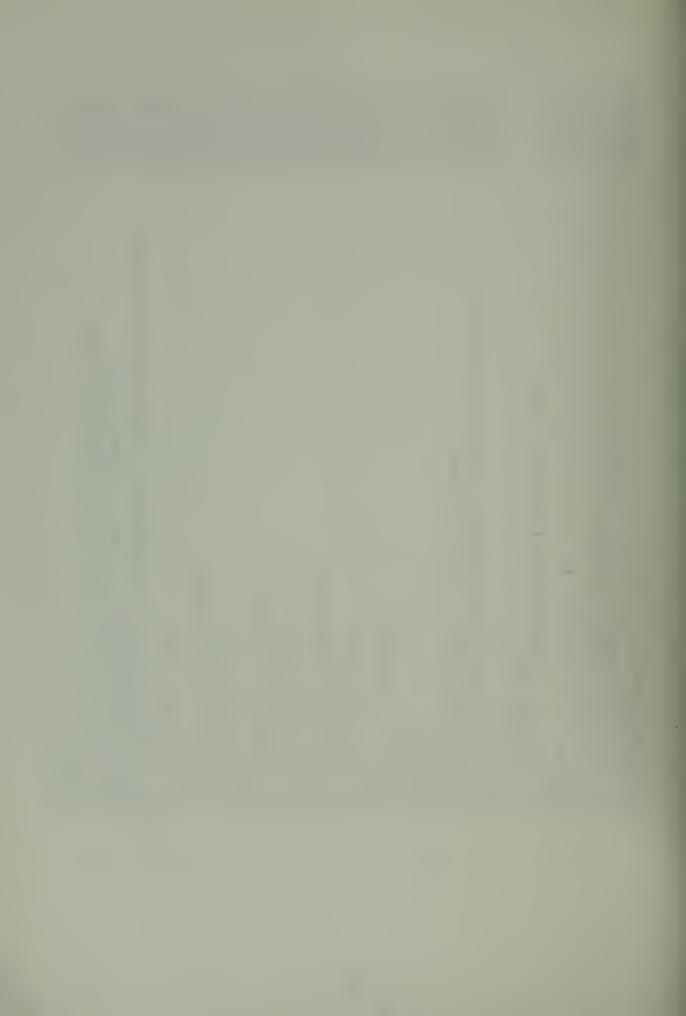
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IF (LV• GT•62) LY=62

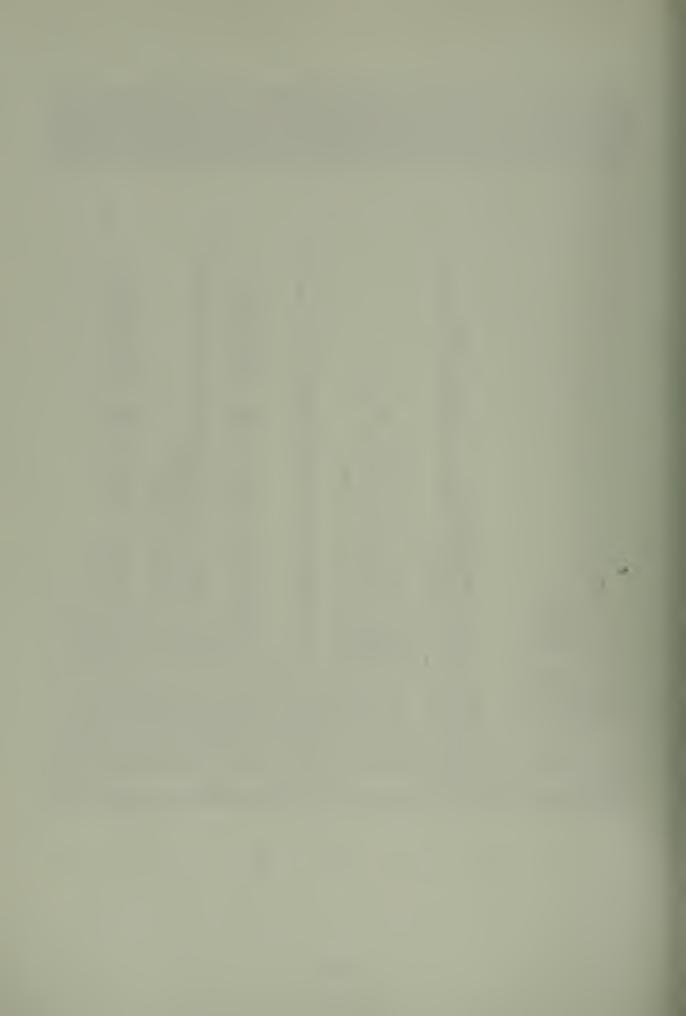
IF (LV• GT•62) LY=62

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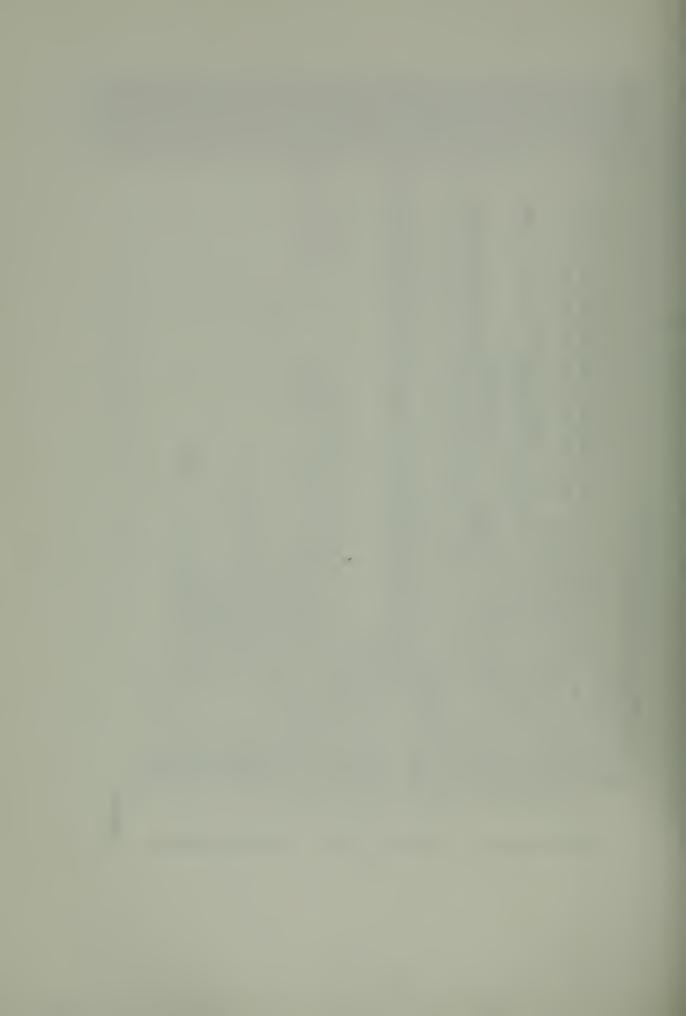
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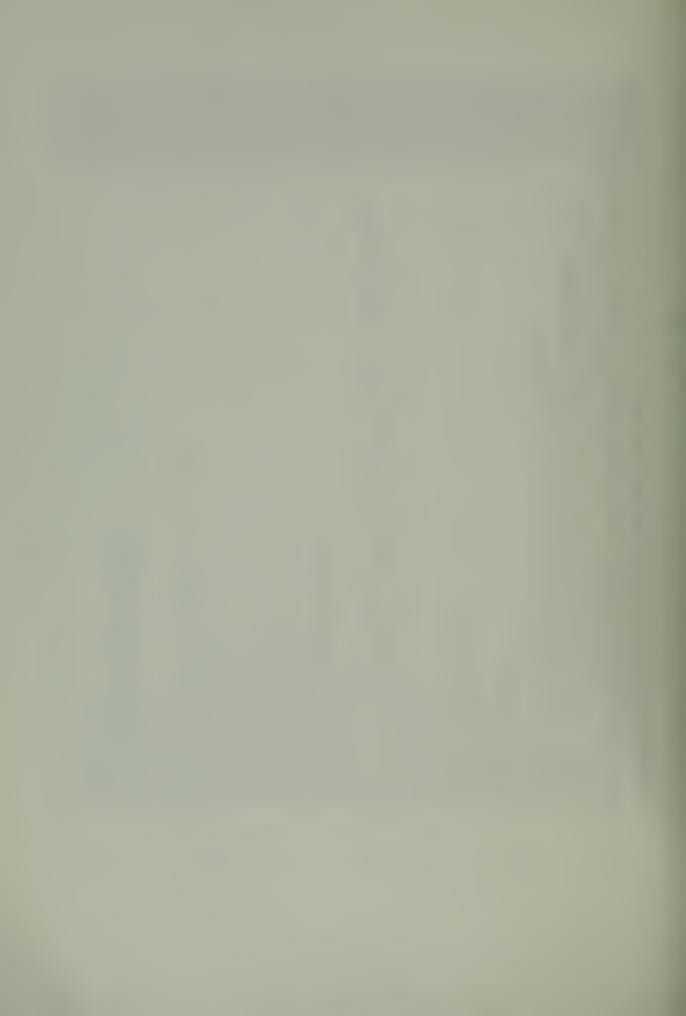
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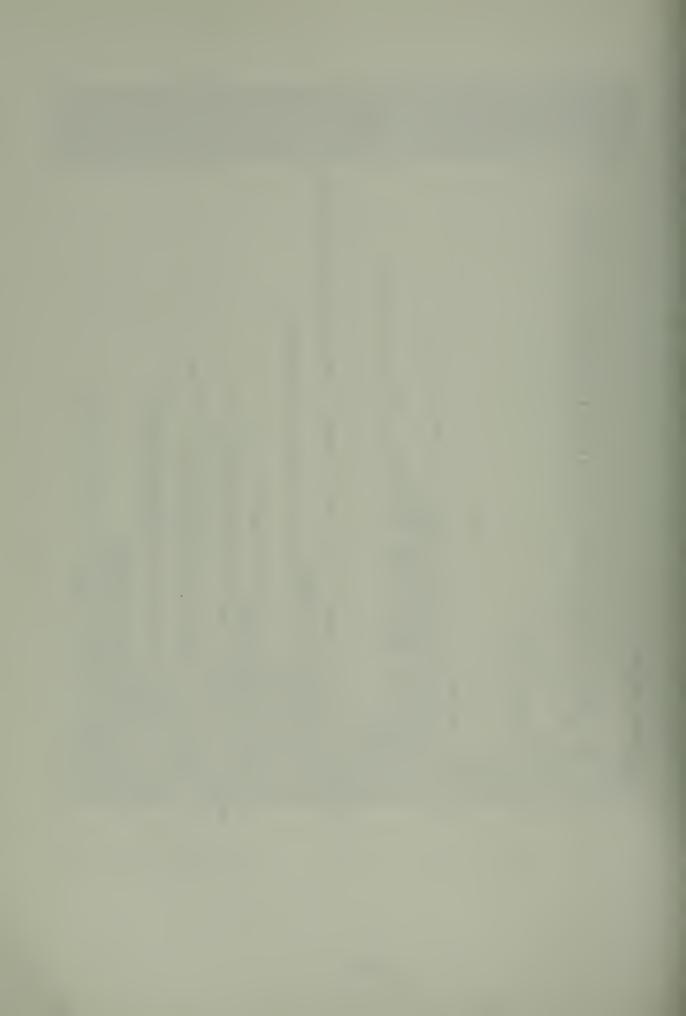
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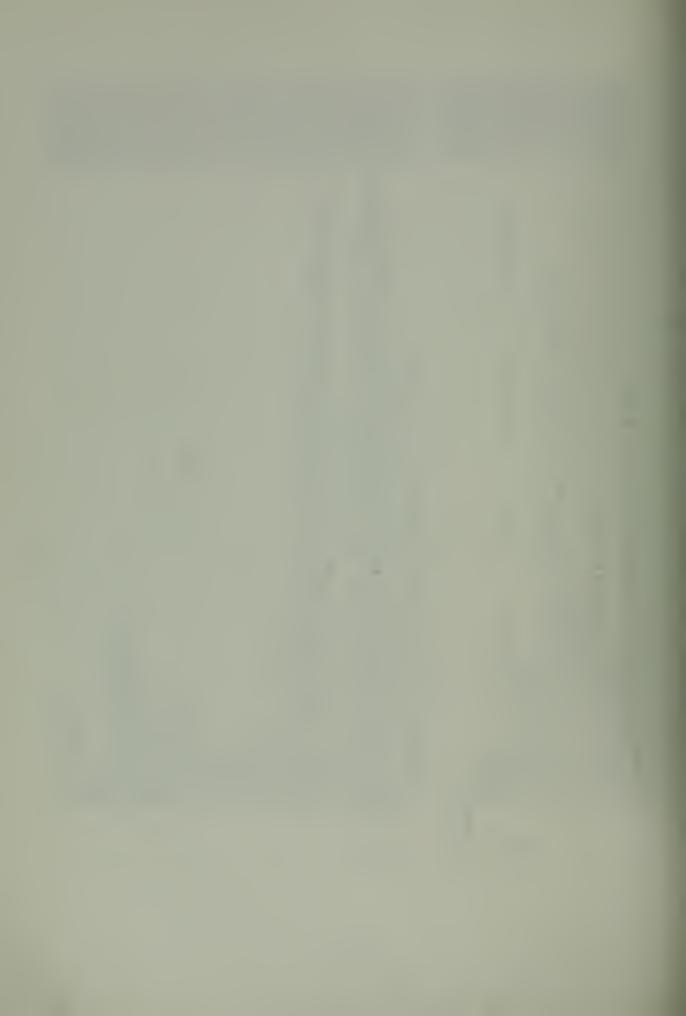
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COMMON / TAB/ INDEX, KEXTRA, MEXTRA, KLIMIT, ML
COMMON / SYM/ I SYM, JSYM, MSYM, FCU, IMS; JMS, QS
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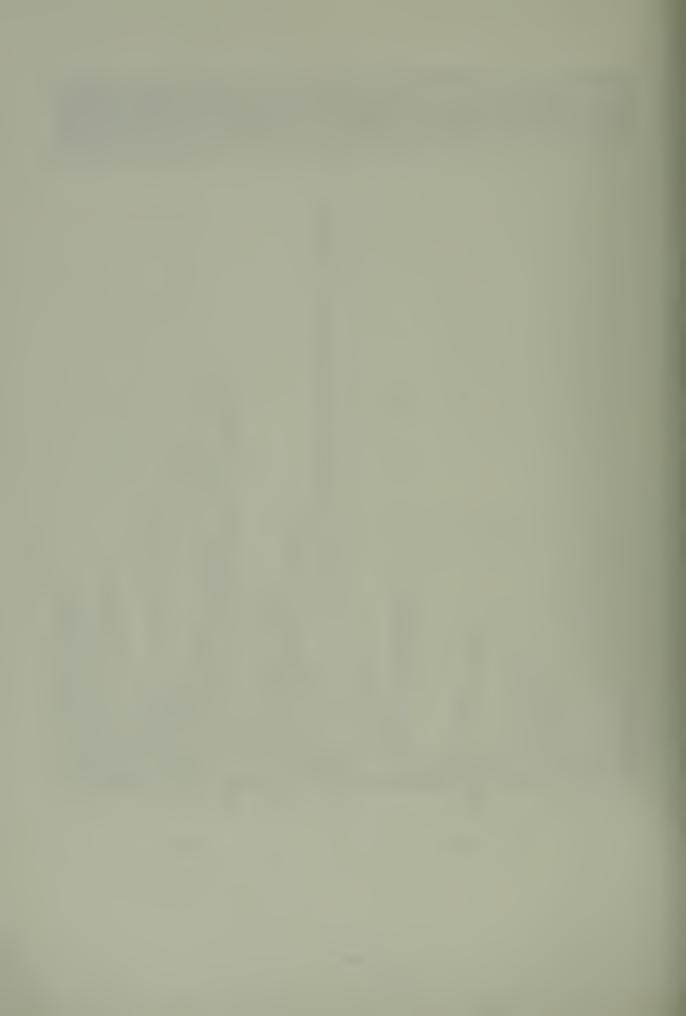
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IF (TOTAL.GT.-EPS) CHECK=

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IF (CHECK.GE.EPS) KTIME

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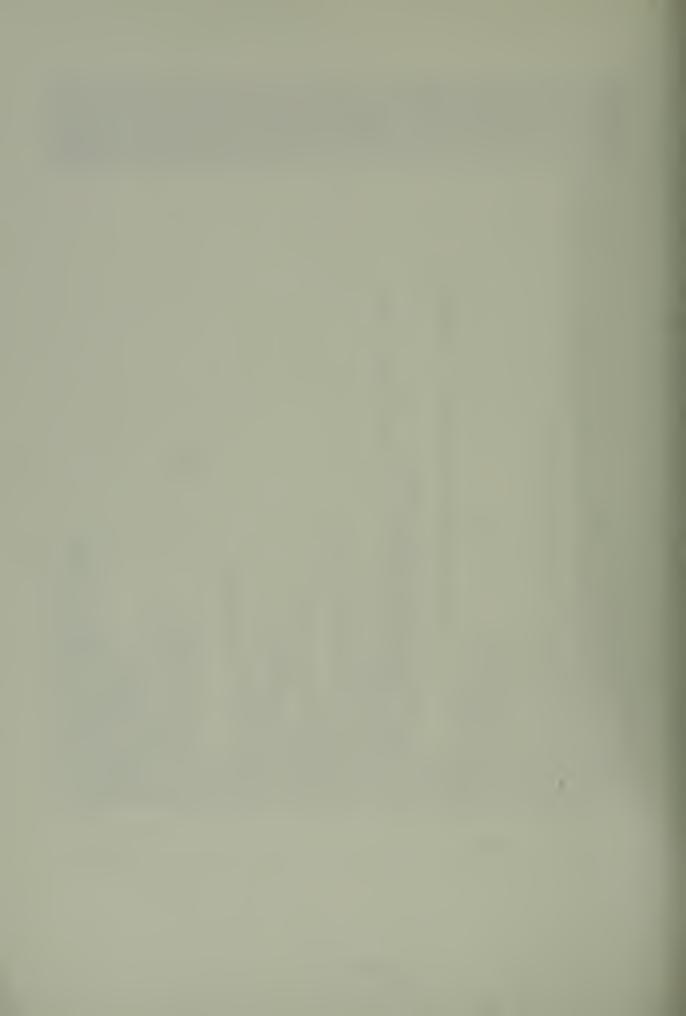
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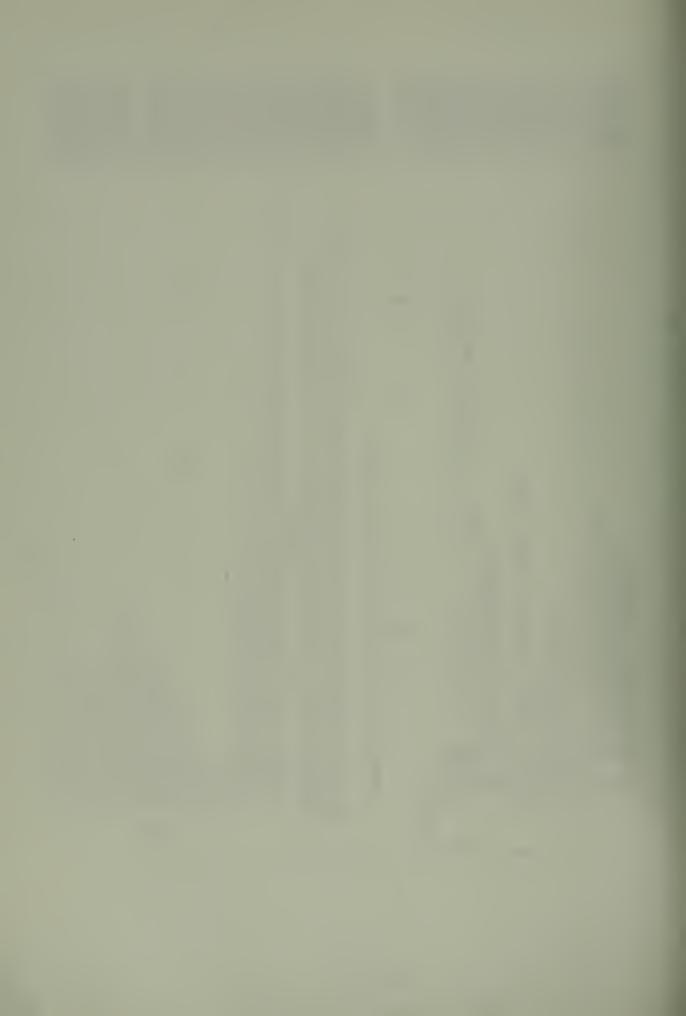
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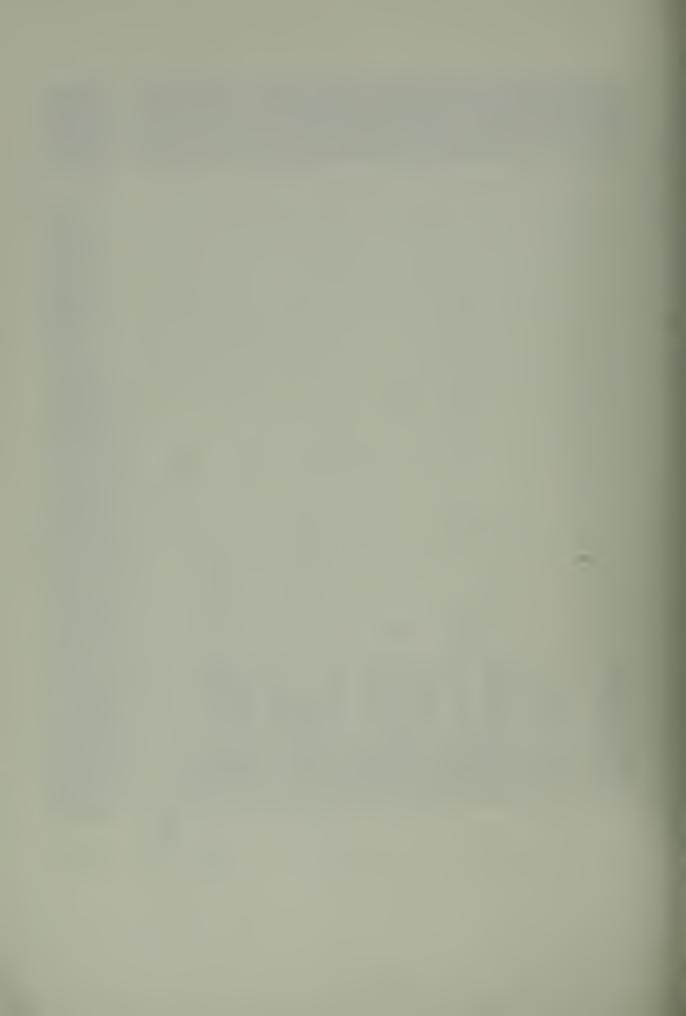
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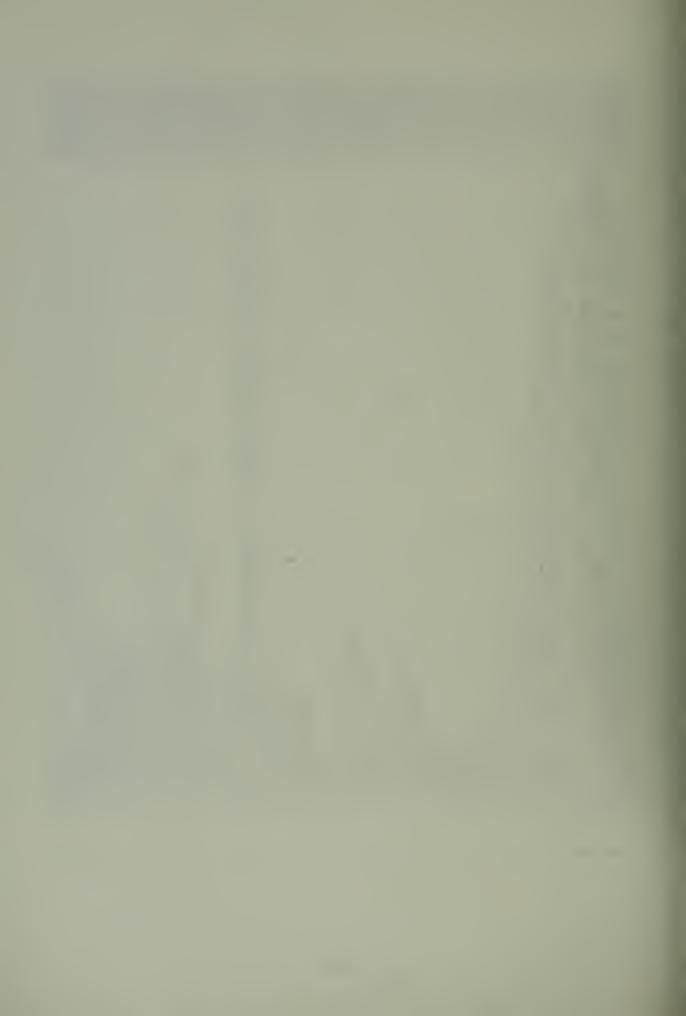
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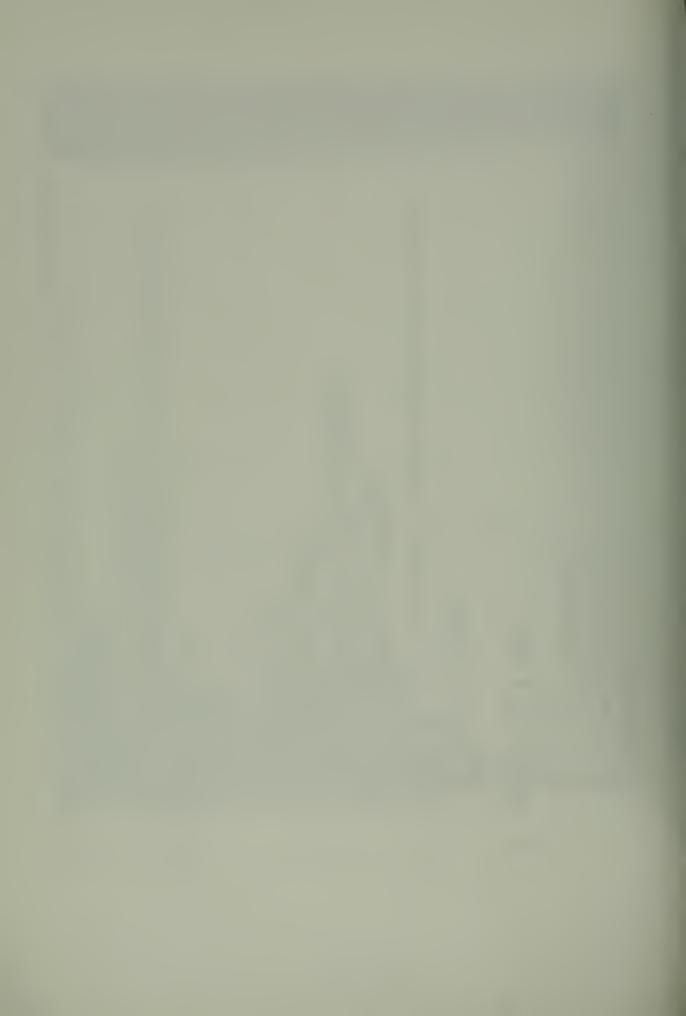
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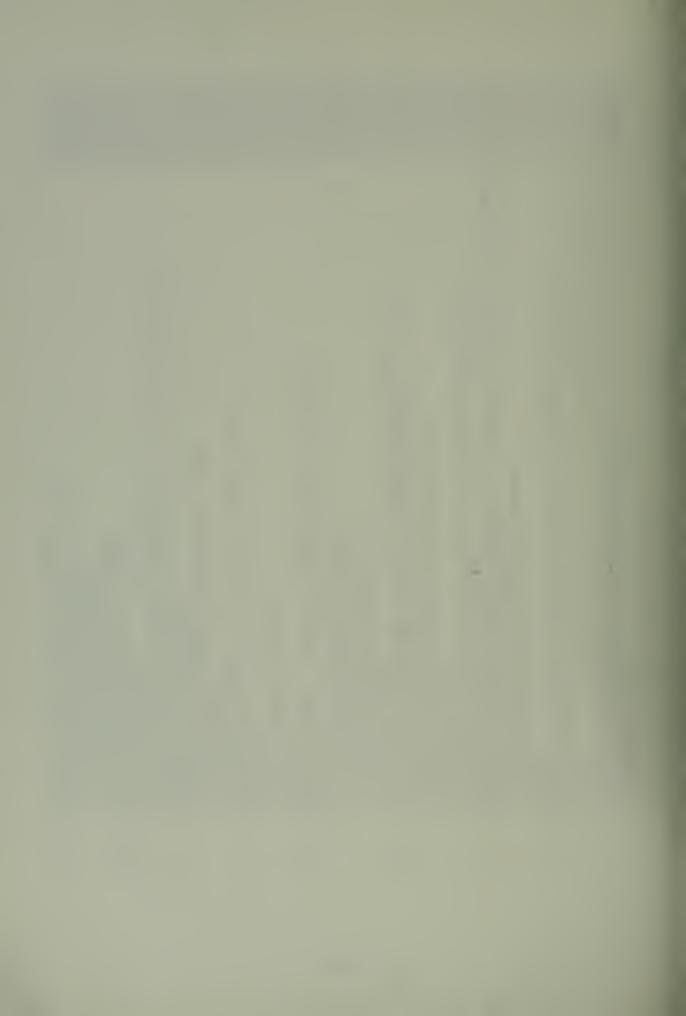
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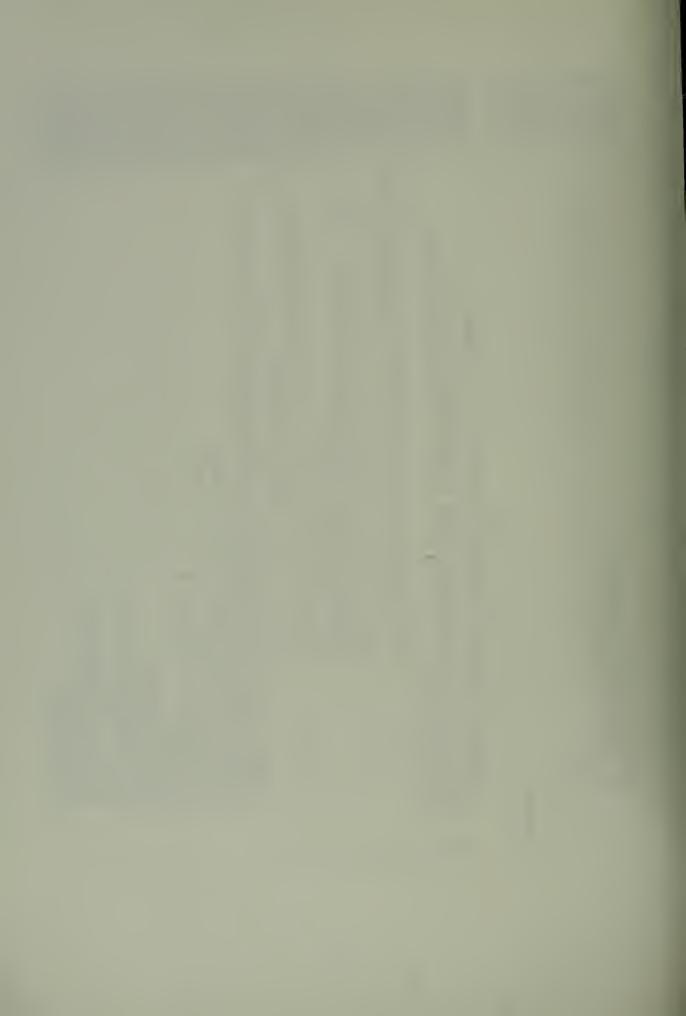
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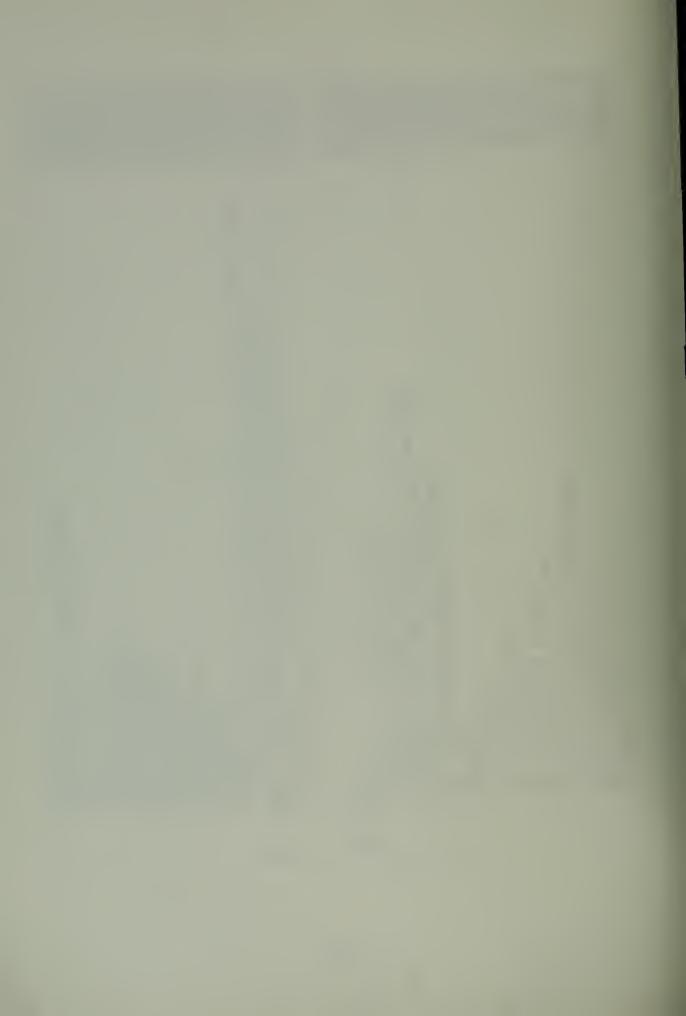
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L*EXPON*APP/2.
M=',14,', K=
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                                                                                                                                                                                                                                                                                                                                                                                                                                               SUBROUT I NE
MOUT=M-1
IF (KOUT-E
SOLN=TOTAL
FORMAT (*
RETURN
END
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SIGHT
                                                                                                                                                                                                                                                                     IMAX, JMAX, II MX, JJMX, I JMX, ALPHA, SIZE, EPS, MODE, BOX, SD, IX, Z
                                                                                                                                                                                                                                            0F
                                                                                                                                                                                                                                            FOR A PARTICULAR LINE SUBROUTINE FUNCT.
                                                                                                                     F (MODE.GT.2) GO TO 3
ALL SHEET (G,GA,XO,YO,PHISYM,NOF)
O TO 4
ALL READ (Z,XO,YO,PHISYM,NOF,IMAX,JMAX,G)
F (DGN.GE.2) WRITE (6,39)
ETURN
CETURN
CORMAT (' GARRAY RETURNS')
                                                                                                                                                                                                                            SUBROUTINE GOLF (R, XI, GIJ, NOF, DGN, NUMB)
DD 1 I=1,IMS
RI=I
I = IMAX+1-I
R=(RI-.5)*DELR-HS
CALL GOLF (R,XI,GIJ,NOF,DGN,NUMB)
G(I,J)=GIJ
IF (ISYM.EQ.2) G(II,J)=GIJ
IF (ISYM.EQ.2) GO TO 1
G(II,J)=GIJ
G(II,J)=GIJ
G(II,J)=GIJ
G(II,J)=GIJ
CONTINUE
GO TO 4
IF (MODE.GT.2) GO TO 3
CALL SHEET (G,GA,XO,YO,PHISYM,NOF)
GO TO 4
CALL SHEET (G,GA,XO,YO,PHISYM,NOF)
                                                                                                                                                                                                                                            S THE FUNCTION G(R,XI)
FUNCTION CONTAINED IN
                                                                                                                                                                                                                                                                                                                                                                                             FUNCT(XS, YS, F, NOF, DGN, NUMB, IJ+F
                                                                                                                                                                                                                                                                    COMMON IMAX, JMAX, ZEROEO.
LMAX=IMAX*3
RLMAX=LMAX
DELXP=SIZE/RLMAX
SXI=SIN(XI)
CXI=COS(XI)
DELYS=DELXP*CXI
DELXS=DELXP*CXI
DELXS=XP*CXI-R*SXI
XP=DELXP*CXI
CIJ=O.
IL=I,LMAX
RL=L
CALL FUNCT(XS, YS, GIJ=GIJ+F
                                                                                                                                                                                                                                           COMPUTES
A KNOWN F
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, RA, NO, NA, N1, N2
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                         EQUATION USED.
                                                                                                                                                                                                                                                                                                                                                                                                                   GIJ=1, F8
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                            N AT POSITION (
IDENTIFIES THE
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                    IJMX, ALPHA, SIZE, Q, S, T, U, V, W, RO
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                          FUNCT (XS, YS, F, NOF, DGN, NUMB
                                                                                                                                                                                            2
XI=1, F8.3,
                                                                                                                                                                                            0
                                                                                                                                                                                                                                                                                                                                           R,XI,GIJ
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                            INPUT FUNCTION
SYSTEM. NOF IL
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O1) E, P
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IZE/2.
RT(XS**2+YS**2)/HS
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NOF.LE.0) GO TO 11
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CCMMON /EQPARA/ A,B

DIMENSION RO(101), R,

AA=A

BB=B

CC=C

DD=D

EE=E

PP=P

IF (NUMB.LE.1) GO TO

AA=S

BB=T

CC=U

CC=U

CC=U

CC=U

PP=Q

PP=Q
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                            FUNCT EVALUATES AS
SECTION COORDINATE
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I INPUT ARRAY READ IN N POINT VALUES. (101) THE FUNCTION.
                                                                 F= AA
                                                                                                                                                                                           FUNCT ION
                                                                                                                                                                                                                                                                                                                               MAXIMA
                                                                                        DISPLACABLE ELLIPTICAL GAUSSIAN:
IF (NOF.6T.3) SO TO 4
F=AA*EXP(-1.*((XS-DD)/BB)**2+((YS-EE)/CC)**2))
GO TO 11
                                                                 • BB) • AND • (ABS (YS-EE) • LE • CC))
                                                                                                                                                                                           RAMP
                                                                                                                                                                                                                                                                                                                              88
                                                                                                                                                                                           DISPLACABLE ELLIPTIC RAI
TO 6
/BB)**2+((YS-EE)/CC)**2
                                                                                                                                                                                                                                                                               B) **2+((YS-EE)/CC)**2
                                     FUNCT I ON:
                                                                                                                                                                                                                                                            STEP FUNCTION
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                                                                                                                                                                                                                                                                                                                                                                           BY
TO
                                                                                                                                                                                                                                                                                                                             CIRCULAR COSINE-SQUARED FUNCTION IF (NOF.GT.7) GO TO 8 F=AA*COS((2.*BB-1.)*PIE*R/2.)**2 GO TO 11
                                                                                                                                                                                                                                RBC.LT.1.) F=AA*((1.-RBC)**PP)
0 11
                                                                                                                                                                                                                                                                                                                                                                            REQUIRES
FOLLOWED B
IS ADDED T
                                     STEP
IF (NOF.GT.1) GO TO 2
F=AA*EXP(-1.*(R*HS/BB)**2)
GO TO 11
                                  ADJUSTABLE RECTANGULAR
IF (NOF.GT.2) GO TO 3
F=PP
IF ((ABS(XS-DD).LE.BB).AN
GO TO 11
                                                                                                                                                                                                                                                          DISPLACABLE ELLIPTIC S

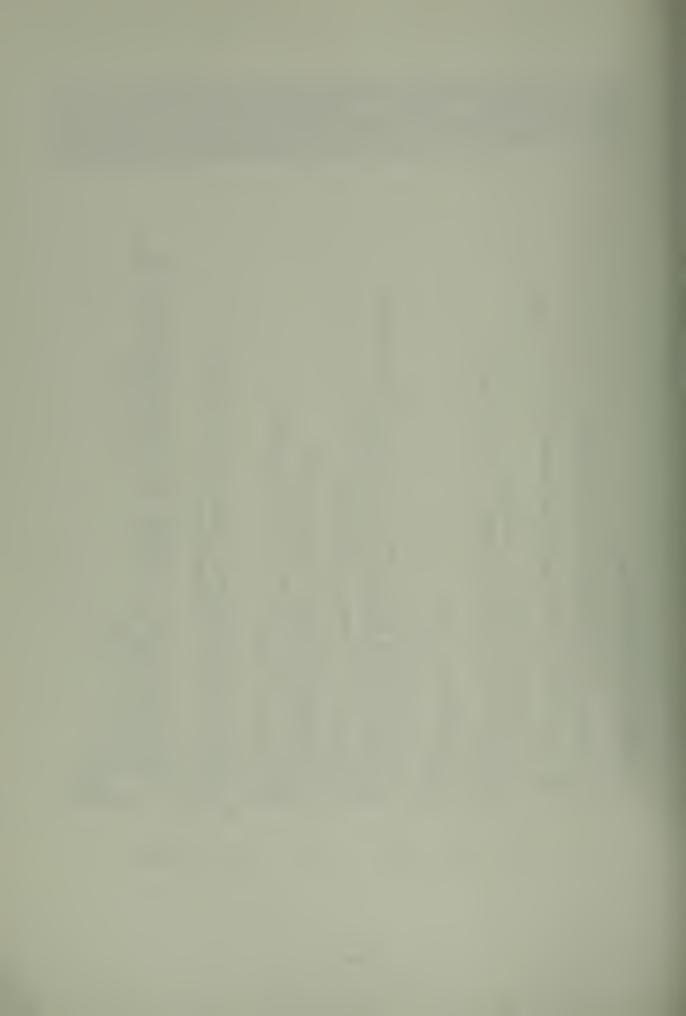
IF (NOF.6T.6) 60 TO 7

RBC=SQRT(((XS-DD)/BB)**2-

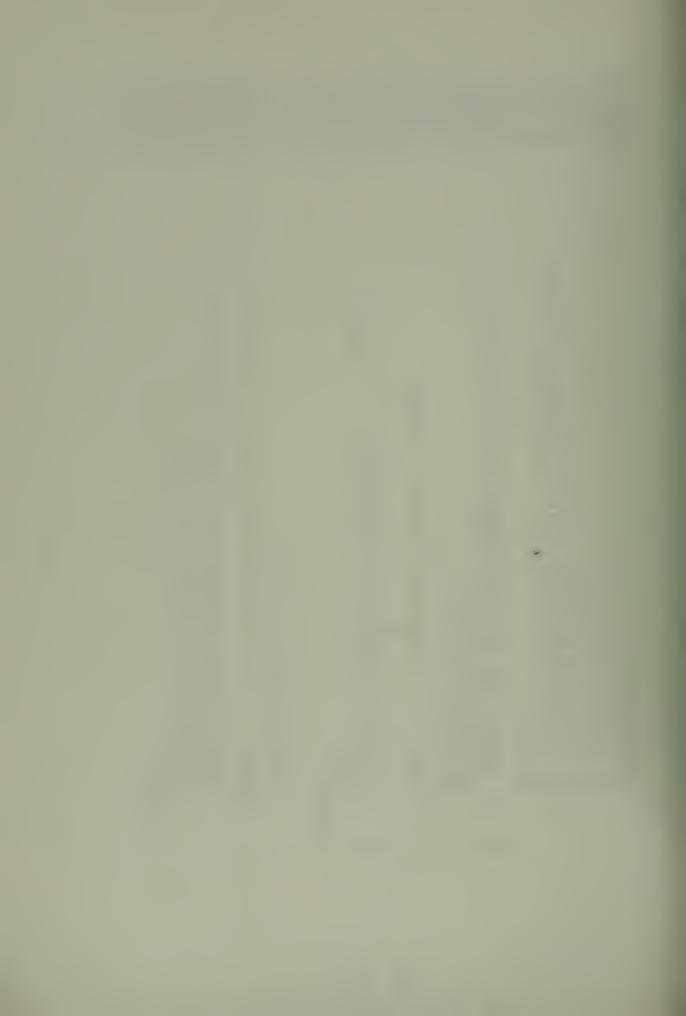
F=0.

IF (RBC.LT.1.) F=AA

GO TO 11
                                                                                                                                                                                                                                                                                                                                                                            FUNCTION:
FREAD: N:
VALUE AA
S) GO TO 9
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GO
DD)
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IF (NOF.GT.5) G
RBC=SQRT(((XS-DD
F=0.
                                                                                                                                            . CONSTANT:
IF (NOF.GT.4) (F=AA
GO TO 11
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SUBROUTINE
A CONSTANE
IN (NOF 6T.8)
IN (NUMB CT.8)
IN NM=N-1
NMM=N-2
RN=N-2
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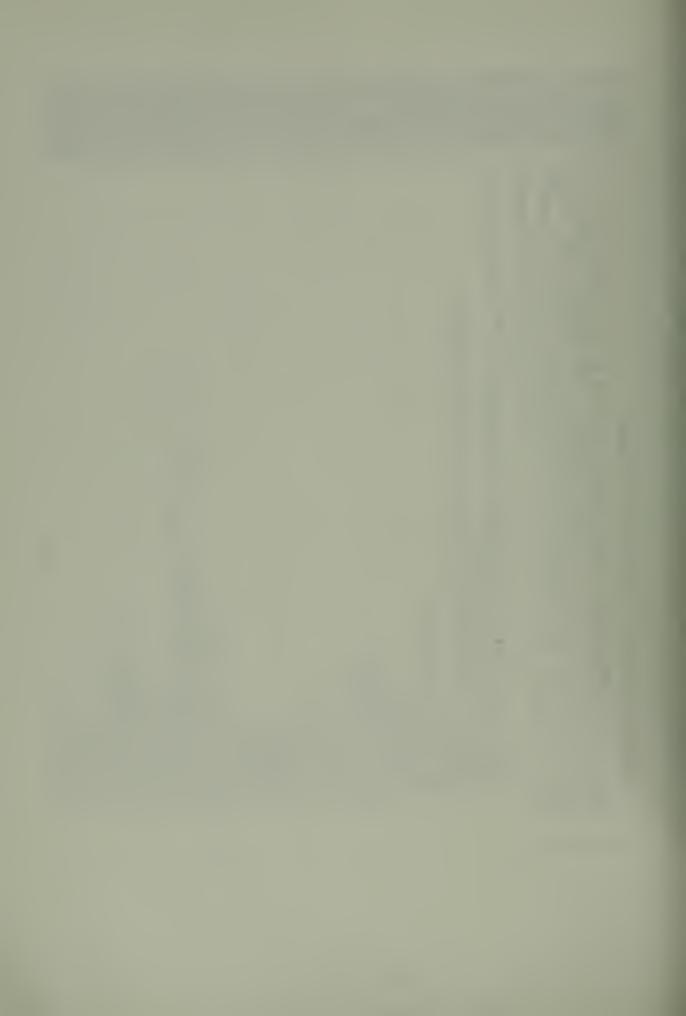


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DIMENSION RO(101),RA(101)
F=0.
IF ((ABS(XS).LE.B).AND.(ABS(YS).LE.C)) F=A
RETURN
END
                                                AND
                                                                                                                                       FUNCTION
                          RO(IR+1)-RO(IR
RA(IR+1)-RA(IR
                                                OCCAS ION
                                                                                                F=1, F8.3
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                          LE.1)
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                                                SB
                                                                                                                              (XS, YS, F)
                                                                          BEYOND
                 1) F=RO(IR)
1) F=RA(IR)
1. AND. (NUMB-L). AND. (NUMB-G)
                                                A: MAY
COUTINE
                                                                                                                                       ROUTINE
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                                                SUBRC
GO TO
                                                                                            WRITE
= , F8
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                                                                                                                                       SPECIAL
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                                                FUNC
RI=R*(RN-1.)+1
RIR=INT(RI)
DI=RICAT(IR)
DI=RI-RIR
IF (NUMB.CE.1)
IF ((IR.NE.NE.N)
IF ((IR.NE.NE.N)
F=F*AA+BB
                                                                                                                              SUBROUTINE
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IF (NOF.G
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GO TO 11
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F=0.
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| No. 
                                           • OR. (XMN.NE.O.) • OR. (YMX.NE.O.) • OR. (YMN.NE.O.)) MXY=OGO TO 3
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                   READ(IN1,58) XD(I), YD(I), XG(I), YG(I), D(I, J), RR(I),
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                        FOR CODE
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                XY(I)=2.
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                  )-(XG(I)-XD(I))*(YO-YD(I))/DEN
)-XD(I)))
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                              ,YD(I),XG(I),YG(I),RR(I),XI(I)
MX,YMN,NOF,I,IM)
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                        DETERMINE CODE, CALCULATE RADIUS & ANGLE
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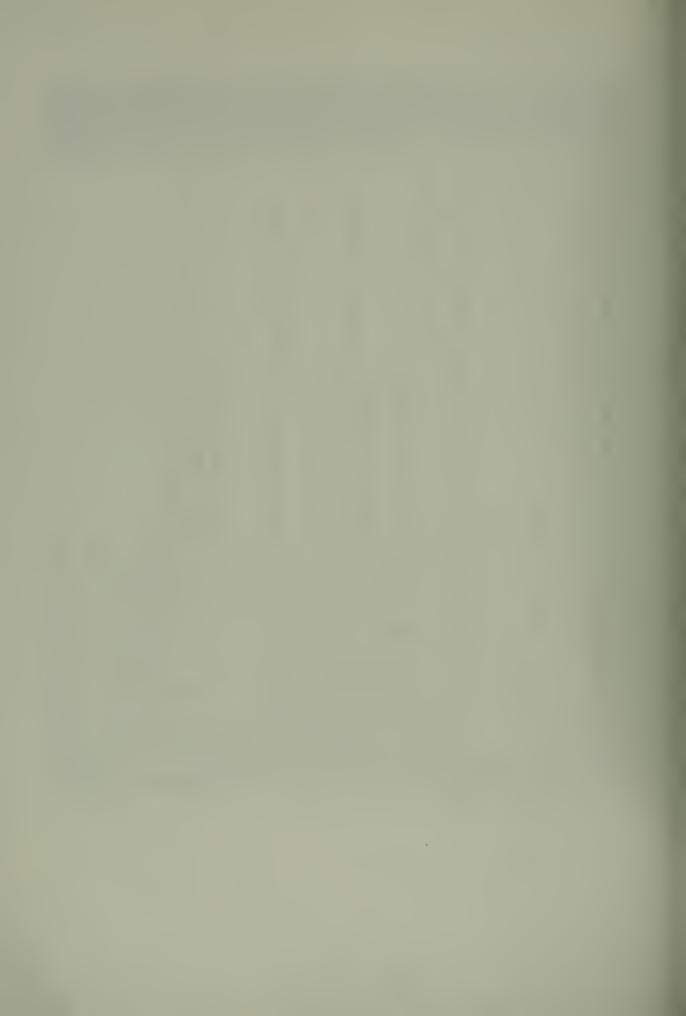
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                                                                                                                                                                                                                                                                                                                          COD
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                             FOR AI
                                                                                                                                                                                                                                                                                                                                                                                                                                     I) = ATANM((YH-YD(I)),(XH-XD(I)))
I) = RRH*SIN(XI(I)-XIH)
TINUE
S AND RADII ARE NOW FILLED FOR I
E THE SET OF VECTORS TO BE USED
IO I=1,IM
RRH=100C0.

YH=RRH×SIN(XIH)

YH=RRH×SIN(XIH)

YTX=-RR(IMT)*SIN(XI(IMT))-YO

YTX=-RR(IMT)*CS(XI(IMT))-YO

XTN=RR(IMT)*CS(XI(IMT))-XO

UG=TAN(XI(IMT))

YH=XH*SQTI(XH-XO)**2)

XIH=RAPUA+UB

RNH=SQTI(XH-XO)**2)

XIH=RAPUN(XI(XH-YO))**2)

XIH=RAPUN(XI(XH-YO))**2)

XIH=RAPUN(XI(XH-YO))**2)

XI(I)=XH-XANM(YH-YO))

XI(I)=XH-XO(I)

XI(I)=XH-XO(I)

XI(I)=XH-XO(I)

XI(I)=0.

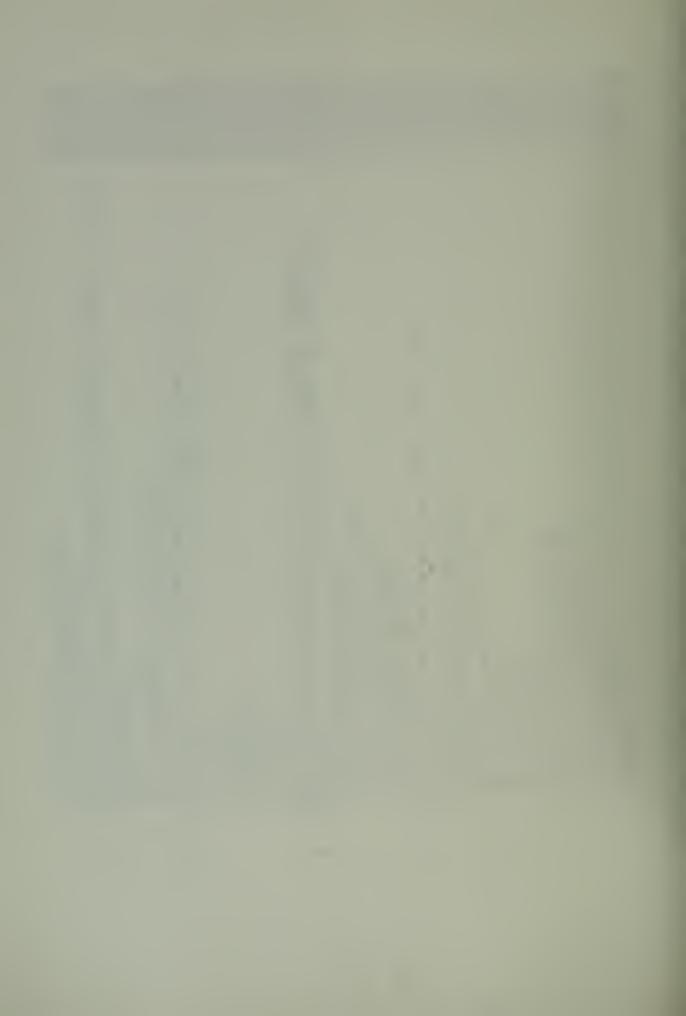
YG(I)=0.

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SMOOTH INTERPOLATION
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H ROW REGULARLY OVER THE ANGLES.
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XI (J) = XD(JX)

XI (J) = DXI/2. -PIE-PHIS,

XI (J) = DXI/2. -PIE-PHIS,

XI (J) = DXI/2. -PIE-PHIS,

XI (J) = XD INE (XG,YD,JM2,XX)

Z4 (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) | XD + XI (J) + G(I),

XI (J) = XI (J) + G(I) + G(I) + G(I),

XI (J) = XI (J) + G(I) + 
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ATANM=ATAN(Y/X)
RETURN
ATANM=PIE+ATANM
ATANM=-PIE+ATANM
                                                                                                                                                                                                                                                                                                                                                                               THE
                                                                                                                                                                                                                                                                                                                                                                               AVERAGE
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12=P16/2

ATANM=S16N(P12, Y)

IF(X.NE.0.)

IF(X.GE.0.)

IF(Y.GE.0.)

IF(Y.GE.0.)

ATAN
                                                                                                          IF (ISYM.EQ.2) 35

J2=JMAX/2+1-J
J3=JMAX/2+1-J
J4=JMAX+1-J
J6=JMAX+1-J
G(II,J2)=G(I,J)
G(II,J3)=G(I,J)
G(II,J3)=G(I,J)
G(II,J4)=G(I,J)
G(II,J4)=G(I,J)
G(II,J4)=G(I,J)
G(II,J4)=G(I,J)
G(II,J4)=G(I,J)
G(II,J4)=G(I,J)
G(II,J4)=G(I,J)
G(ST=(G(I,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G(II,J1)+G
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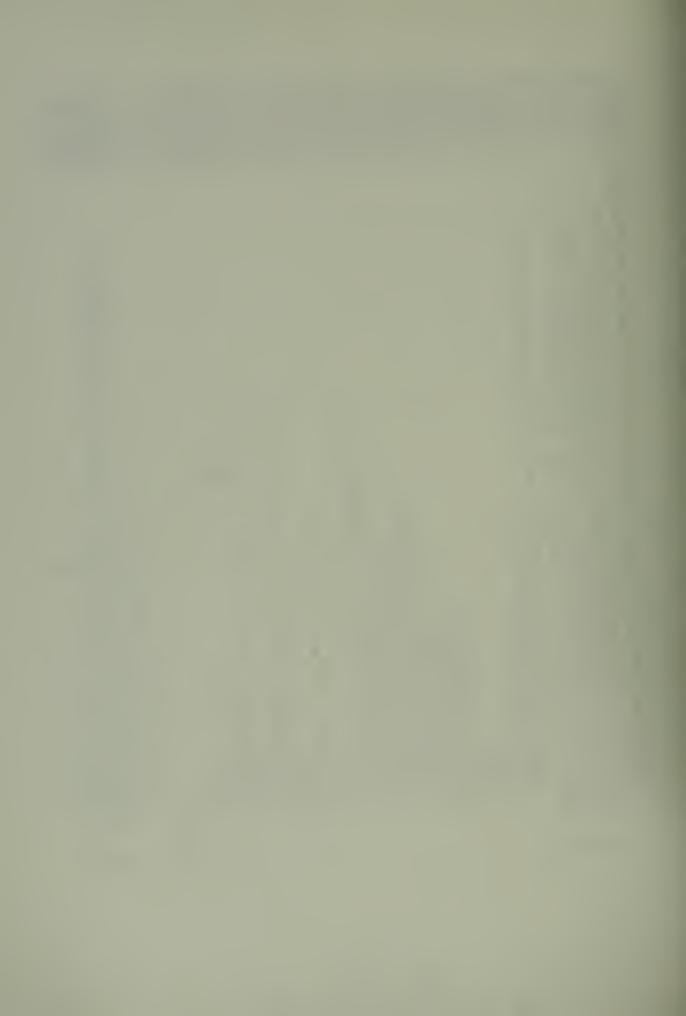


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SUB000070
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SUBROUTINE SIM (XD,YD,XG,YG,D,R,XI,XY,XO,YO,PS,XM,XN,YM,YN,10G,NF)

SIM SIMULATES THE FRINGE NUMBER DATA ONE WOULD OBTAIN FROM THHOLOGRAPHIC INTERFEROGRAM PROCESS FOR A KNOWN FUNCTION AS CONTAINED IN SUBROUTINE FUNCT. THE GRID BOX DIMENSIONS MUST EXCEED THE INVERSION CIRCLE SIZE, AND APERATURE POINTS SPECIF MUST FALL BETWEEN XI = -40 DEGREES, AND XI = +130 DEGREES.
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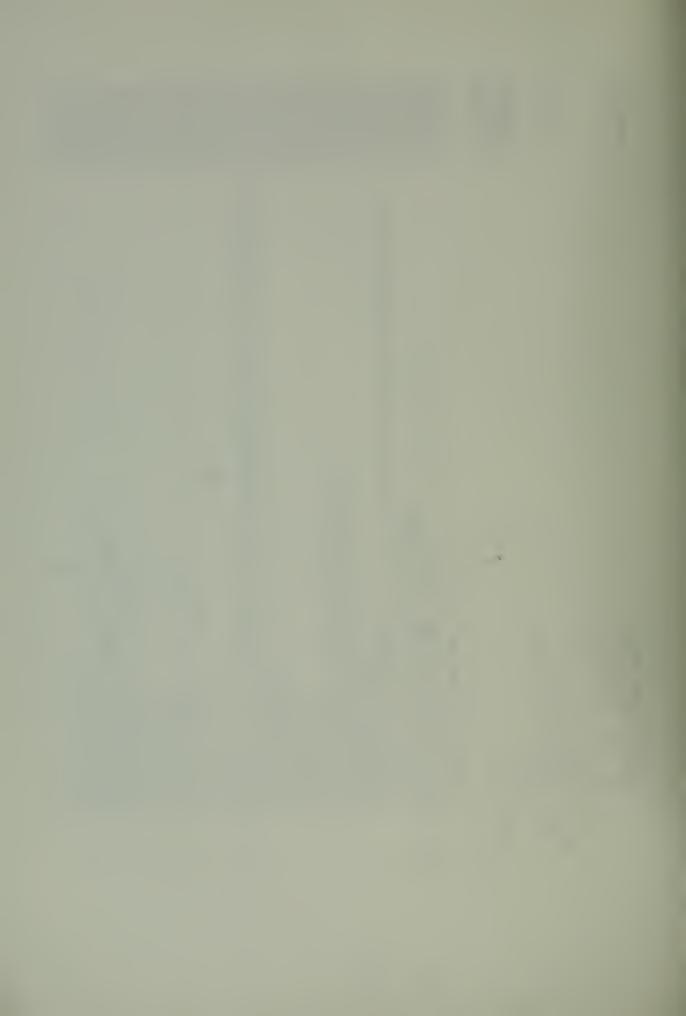
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DIMENSIONS
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                                                                                                                                                                                                                                              (6),
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                      SUBROUTINE GPUNCH (2, XO, YO, PHS, NOF, IMX, JMX, G
                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                                  COMMON /SYM/ ISYM, JSYM, MSYM, FCU, IMS, JMS, QSYM
                                                                                                                                                                                                                                                                                                                                                                                                                                               FUNCTION GARRAY
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X)
MX, JMX, ISM, JSM, IMS, JMS
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IGT=IGB-IB+IT
IGT=IGB-IB+IT
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JMAX) FCU=JMAX
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DO 1 I=1,24
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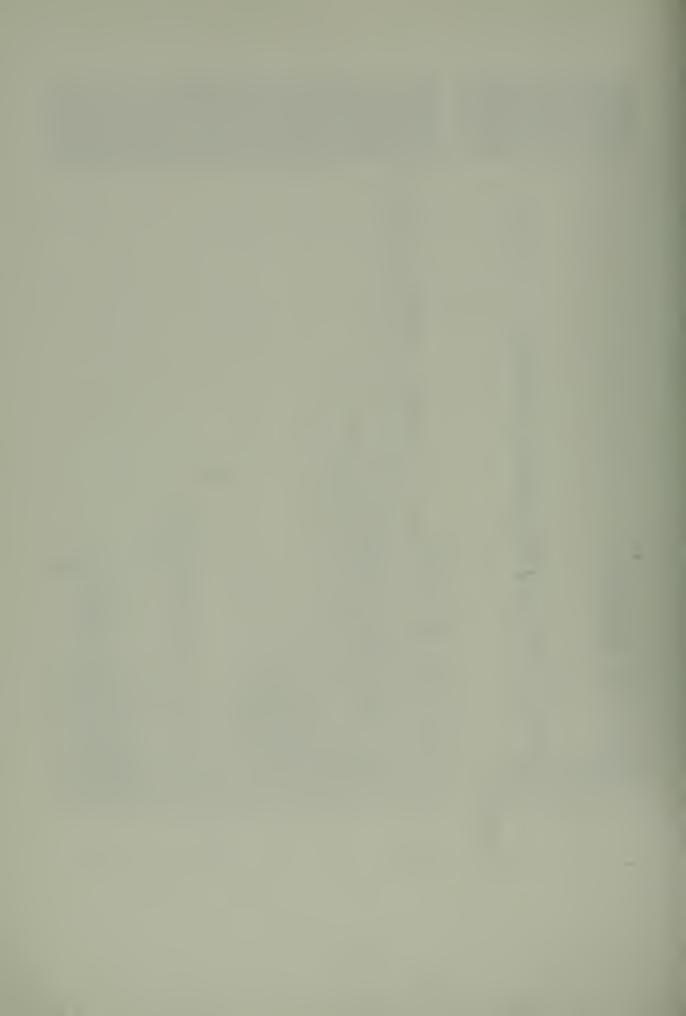
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CCMMON /TAB2/ INDEX(7); JS
DIMENSION G(IMAX, JMAX); G
DIMENSION G(IMAX, JMAX); G
DATA BL, PL, ST, DH, EX/IH,
JMS2=JMAX/2+1
JMS2=JMS2+JMS2=1
JMS3=JMS2+JMS2=1
JMS3=JMS2+JMS3
WRITE(6,67) (ST, I=1,120)
A(I)=G(I; J)
C(I)=GA(I; J)
A(I)=G(I; J)
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A(I)=G(I; J)
A(I)=G(I; J)
BS=0
CALL INTERP (C, IMAX, AS, B
CALL INTERP (C, IMAX, AS, B
CALL INTERP (C, IMAX, AS, D
WRITE (6,69) J
SMALL=0
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IF(C(I)=GT-SMALL) SMALL=
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ICON=1
IF(BAND.LT.0
IF(BAND.LT.0
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WRITE(6,49)
CALL MTMP11
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SPLINN(X,Y,M,XINT,YINT
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RAT=(RIM-1.+2.)

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BI=I

AI=RAT*(BI+BS)

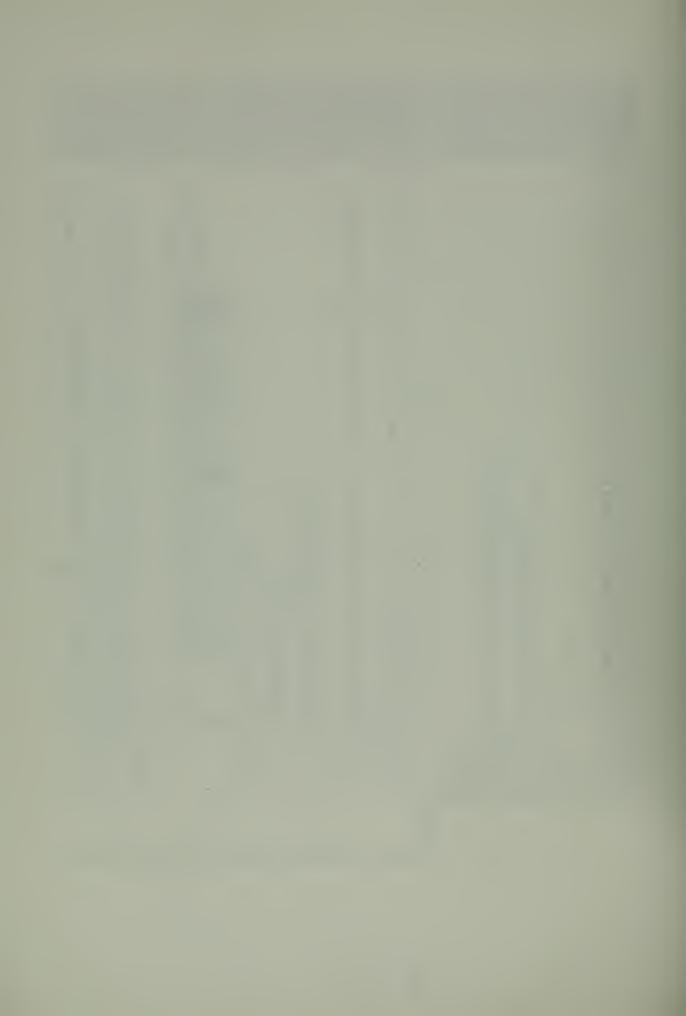
IA=AI-FLOAT(IA)

IF(IA.EQ.0).0

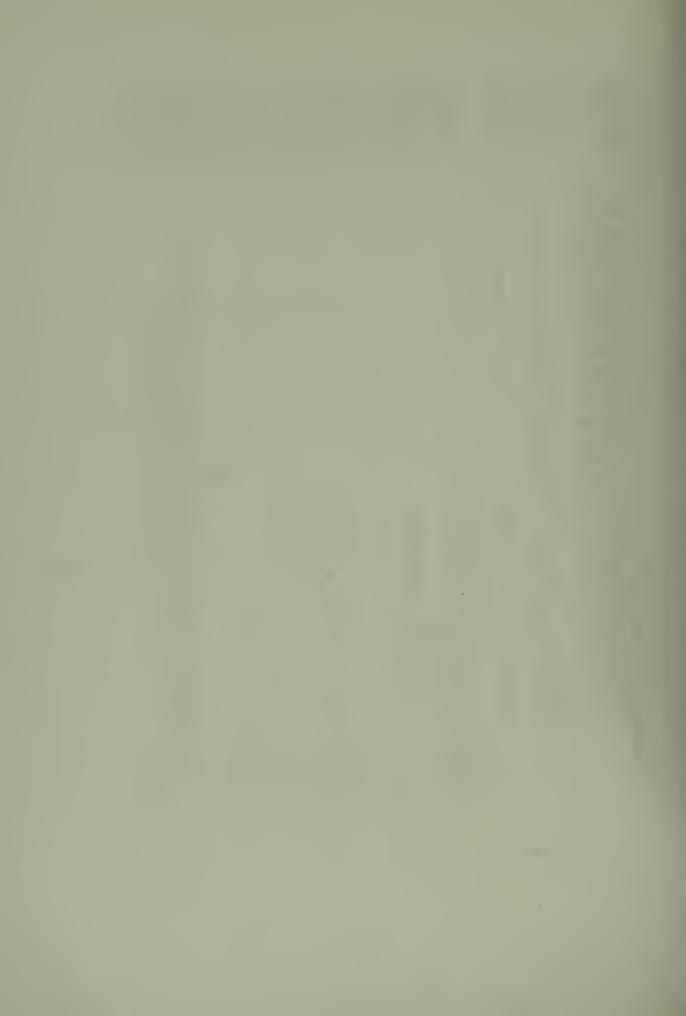
IF(IA.EQ.0).0

IF(IA.EQ.0) B(IA.EQ.0)

IF(IA.EQ.0) B(IA.EQ.0) B(IA.EQ.0)
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- NUMBER OF COLUMNS IN Y.

- ARRAY FOR PLOT TITLE. REAL*4. T(24).

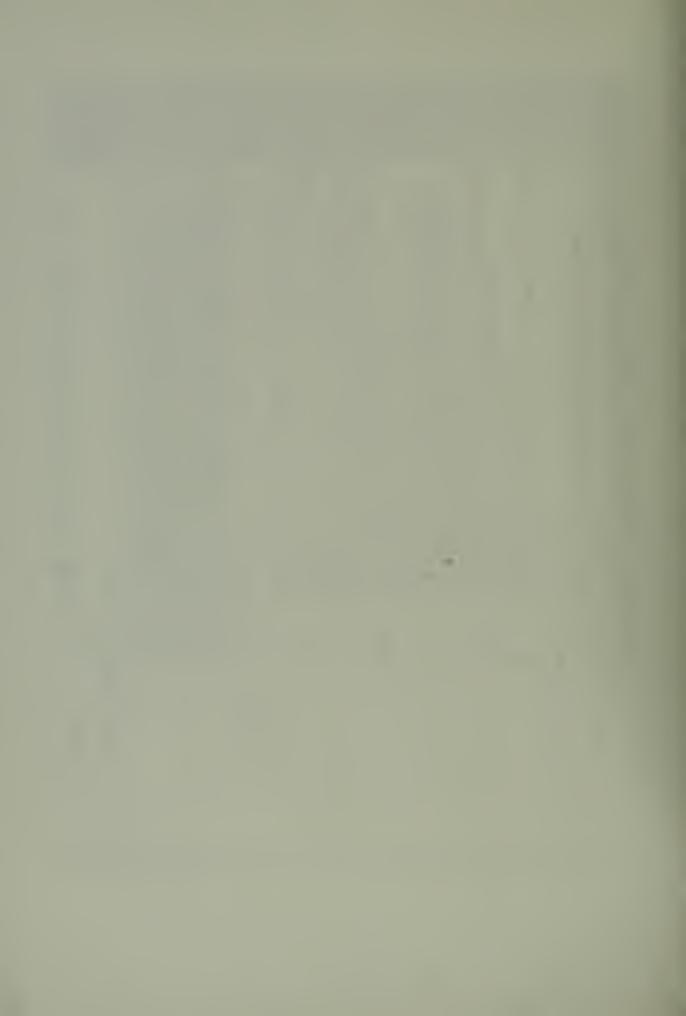
- BANDWIDTH FOR THE CONTOURING. IF BND IS ZER BND=(MAX(Y)-MIN(Y))/15.

- A LINEAR TRANSFORMATION MAYBE PERFORMED ON THE AZ=0 THEN AZ WILL BE COMPUTED SUCH THAT IN AX(Y)1, MIN(Y)1, MILL BE LESS THAN II, AND BZ WILL BE LEFT AS INPUT.

- SEE UNDER AZ WHICH COUTOURING WILL BEGIN. I AMIN(Y) THEN AMIN WILL BE CALCULATED BY AMIN(Y) THEN AMIN WILL BE CALCULATED BY BND.

- THE LEVEL AT WHICH COUTOURING WILL BE CALCULATED BY AMIN(Y) THEN AMIN(Y) AS DETERMINED BY BND.

- IF IJT=O AMIN WILL BE CALCULATED AS DESCRIB ARRAY WILL BE PRINTED IN THE PLOT FORMAT.
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IF NECESSARY THE MAP WILL BE SEGMENTED COLUMNWISE.
THE ROWS AND COLUMNS ARE NUMBERED ALONG THE EDGES.
THAT A SEGMENTED MAP MAYBE EASILY JOINED TOGETHER.
ONLY THREE SIGNIFICANT FIGURES WILL BE PRINTED AT
EACH POINT. THE POSITION OF THE FIRST SIGNIFICANT
DIGIT WILL BE DETERMINED BY MAX(|MAX(Y)|, |MIN(Y)|)
THE PLOT WILL BE PRODUCED ON A INCH INCH GRID. IT
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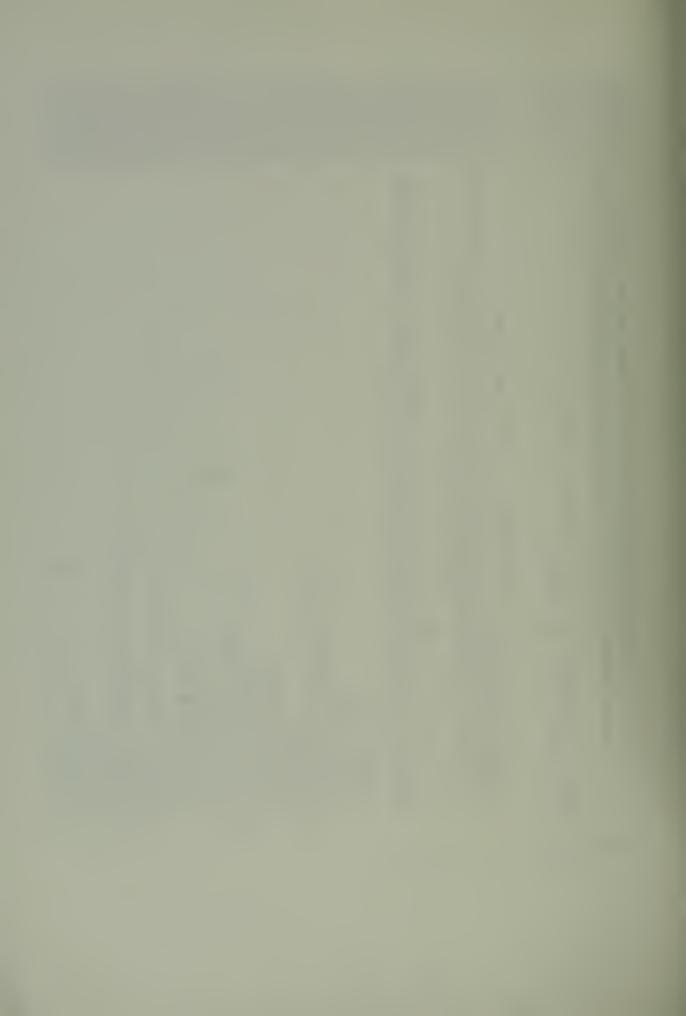
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) B (140), C (140), D (140), IH (20), Y (N, M), TP (10), BTM (10), BTX (10), BT (10), KG (10), T (24), F (140), G (140), H (140)
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YMAX=Y(1,1)
YMAX=Y(1,1)
DO 20 1=1;N
DO 10 J=1;N
DO 10 J=1;N
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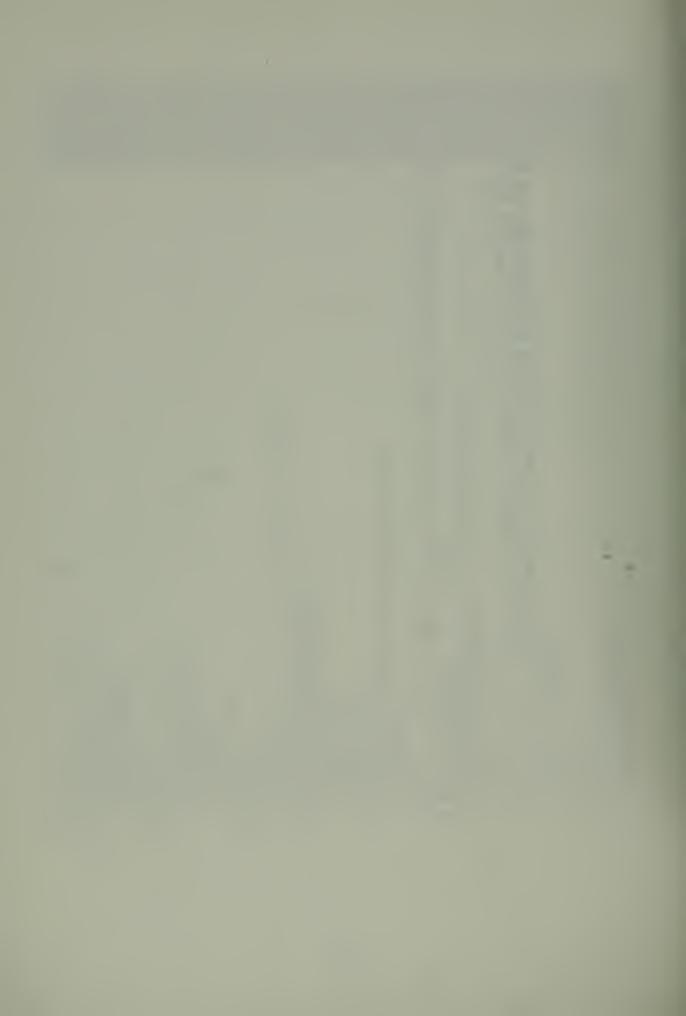
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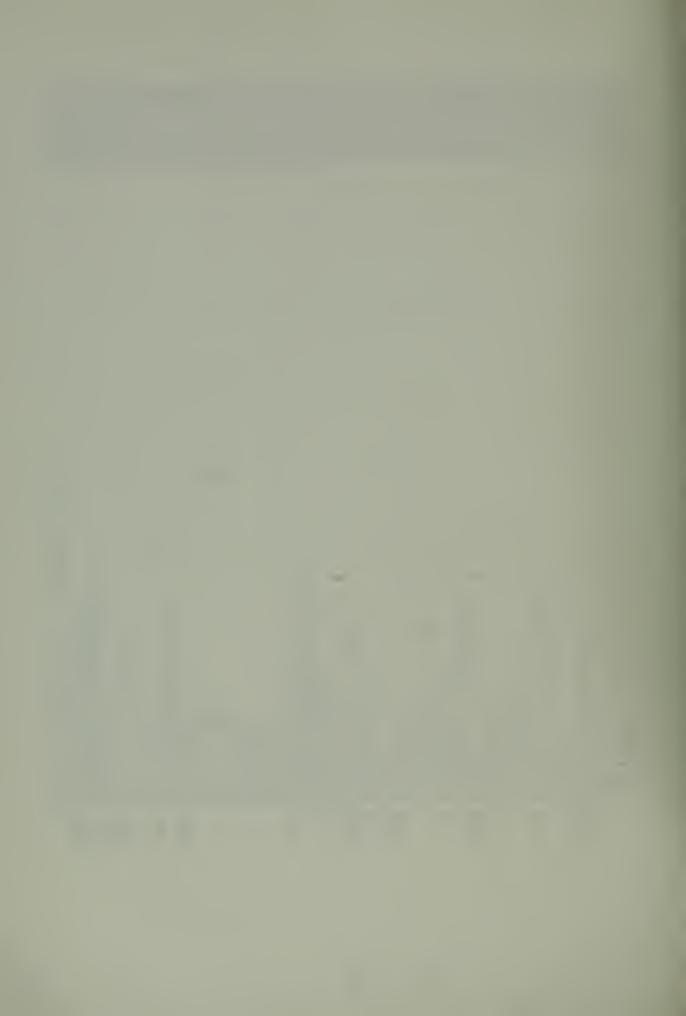
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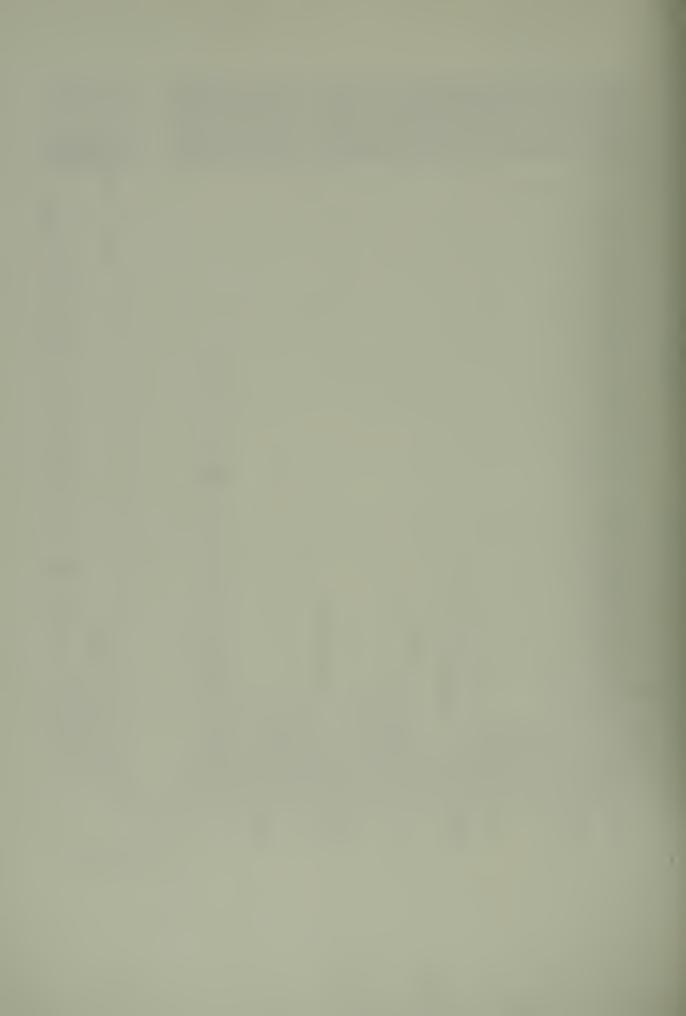
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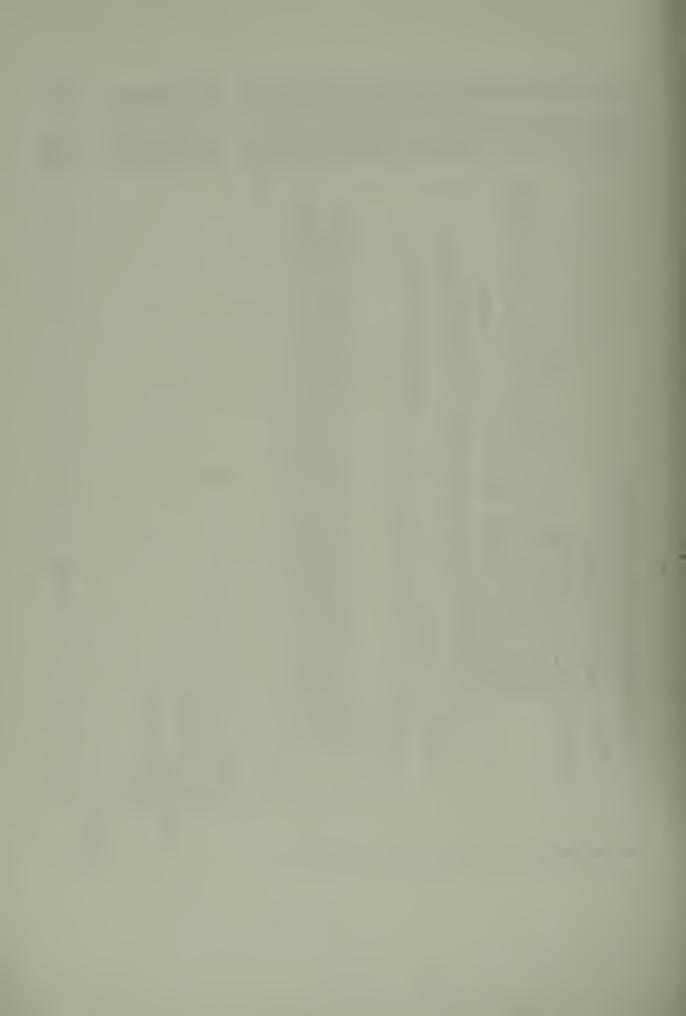


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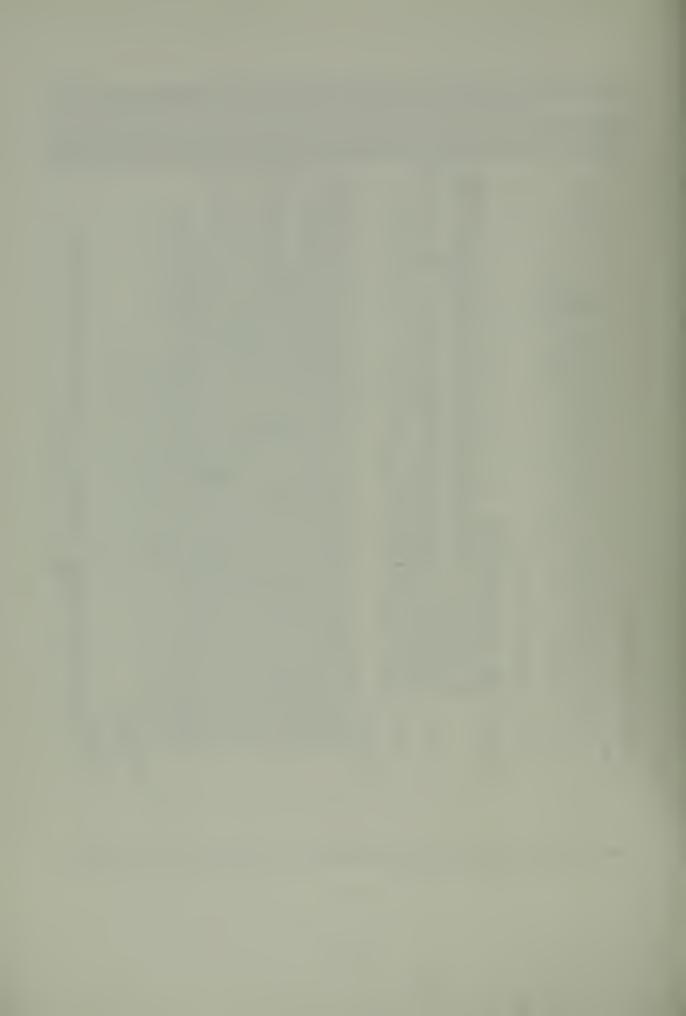


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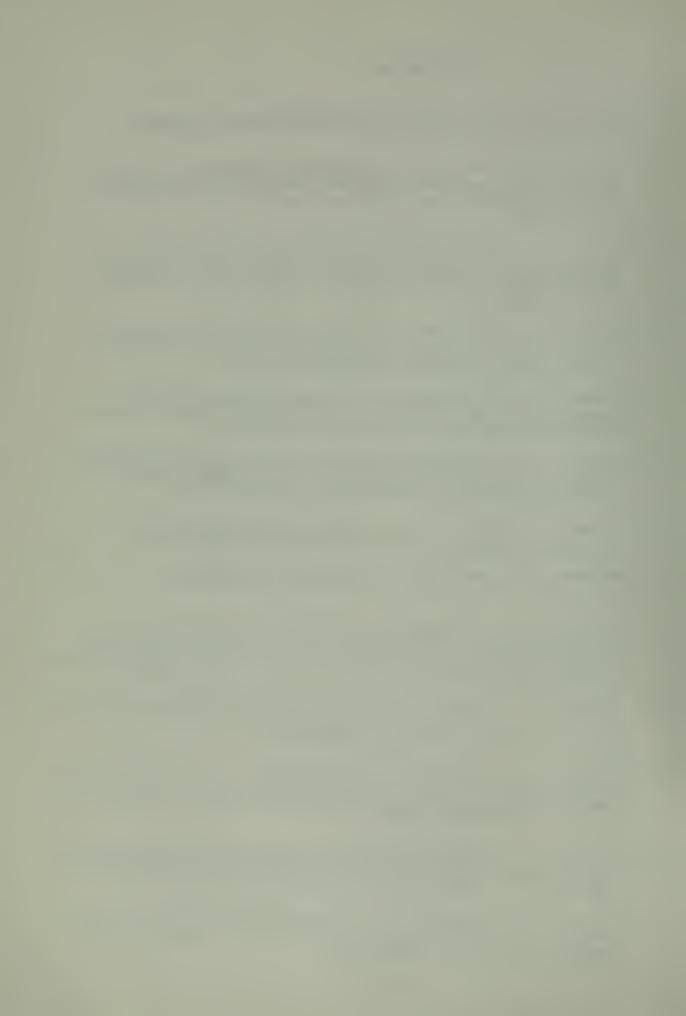
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This study was an attempt to map the density field in a fin-flat plate junction three-dimensionally using holographic interferometry. This investigation has extended the density studies by Matulka [12, 13] and Jagota [3, 4] to include, for the first time, interferometric fringe information obtained through a transparent model in supersonic flow. The fringe information was then inverted by a FORTRAN computer program to produce a plot of the density field around the model. feasibility of the method was demonstrated.

The factors which are thought to have limited the success of the experiment include vibration of the model, fluctuations of the tunnel flow and the fact that the model was somewhat too large in relation to the size of the wind tunnel test section. Schlieren photography was used to look through and around the model and to verify that the same flow was established as was reported by Thomas [23, 24] and Winkelmann [26, 27].

The data reduction of holographic interferograms was, for the first time, accomplished using photographic enlargements. This technique is considered to be much easier and more accurate than the one used in the previous investigations. However, the data reduction step, because of the time and labor involved, is considered to be the rate controlling process of the whole analysis.

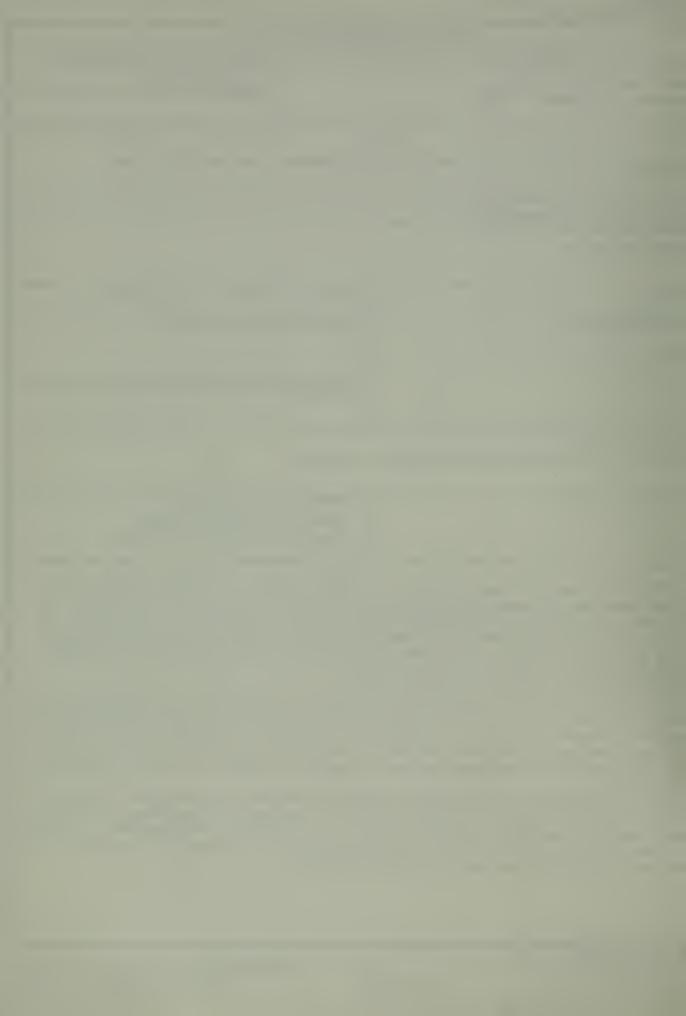
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